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## REDUCED COMPUTATIONAL COST, TOTALLY SYMMETRIC ANGULAR QUADTRATURE SETS FOR DISCRETE ORDINATES RADIATION TRANSPORT

THESIS

Joseph M. Oder, Captain, USAF AFIT/GAP/ENP/97D-07

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# REDUCED COMPUTATIONAL COST, TOTALLY SYMMETRIC ANGULAR QUADRATURE SETS FOR DISCRETE ORDINATES RADIATION TRANSPORT

### THESIS

Presented to the Faculty of the School of Engineering

of the Air Force Institute of Technology

Air University

Air Education and Training Command

In Partial Fulfillment of the Requirements for the

Degree of Master of Science

Joseph M. Oder, B.S., M.S.

Captain, USAF

November 1997

Approved for public release; distribution unlimited

AFIT/GAP/ENP/97D-07

## **REDUCED COMPUTATIONAL COST,**

### TOTALLY-SYMMETRIC

## ANGULAR QUADTRATURE SETS FOR

## DISCRETE ORDINATES RADIATION TRANSPORT

Joseph M. Oder, B.S., M.S. Captain, USAF

Approved:

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<u>1 Dec 199</u>7 date

24 Nov 97 date

24 Nov 97 date

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### Abstract

Several new quadrature sets for use in the discrete ordinates method of solving the Boltzmann neutral particle transport equation are derived. These symmetric quadratures extend the traditional symmetric quadratures by allowing ordinates perpendicular to one or two of the coordinate axes. Comparable accuracy with fewer required ordinates is obtained.

Quadratures up to seventh order are presented. The validity and efficiency of the quadratures is then tested and compared with the LQ<sub>n</sub> level symmetric quadratures relative to a Monte Carlo benchmark solution. The criteria for comparison include current through the surface, scalar flux at the surface, volume average scalar flux, and time required for convergence. Appreciable computational cost was saved when used in an unstructured tetrahedral cell code using highly accurate characteristic methods. However, no appreciable savings in computation time was found using the new quadratures compared with traditional  $S_n$  methods on a regular Cartesian mesh using the standard diamond difference method. These quadratures are recommended for use in three-dimensional calculations on an unstructured mesh.

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### I. Introduction

This research developed a set of angular quadratures that are computationally efficient when used with the discrete ordinates method to solve the three-dimensional Boltzmann neutral particle transport equation. The quadrature sets contain directions perpendicular to one or more cardinal directions. These sets are tested for accuracy and computational efficiency. Performance comparisons are made with traditional (level symmetric) quadrature sets. Types of problems and conditions where these quadratures are most applicable are discussed.

### Background

The foundation of transport theory is the Boltzmann transport equation (BTE). This equation, formulated over a century ago, was originally developed for the study of the kinetic theory of gases (1: 1). Preliminary study in the field was primarily of the diffusion of light by the atmosphere. Then study began in the early part of the twentieth century on investigating the diffusion of energy through the atmosphere of a star (2: 1). The scale of these problems are such that they can be modeled as semi-infinite media with one-dimensional geometry and therefore the methods of their solution are of limited application (1: 1).

In the 1940s, interest in the military and industrial application of nuclear energy stimulated a tremendous amount of research into neutral particle transport. The incredible urgency involved in the research of nuclear energy induced by the events of World War II necessarily resulted in the development and use of approximate methods for solving the linearized transport equation (2:1).

#### Motivation

Military and industrial research into the use of nuclear energy using actual nuclear material has decreased substantially in recent years. The comprehensive nuclear test ban treaty (CTBT), if ratified, will eliminate our ability to obtain any further real data on new weapons designs or systems survivability in or near a real threat environment. Some aspects of the radiation environment resulting from a nuclear detonation are simulated at various test sites. These tests can only approximate the actual post detonation environment and are very costly (17). It is currently politically undesirable for private industry to perform research using significant of amounts nuclear material or to build new nuclear research facilities. These and other factors have greatly increased the need for accurate computer modeling of nuclear material and effects. The high performance computers needed for this modeling are very expensive, and therefore anything that can increase the efficiency of these machines will translate directly into

substantial savings by increasing their productivity and extending their useful life. Reduction in computation time also increases the practicality of analyzing a large variety of similar scenarios for threat analysis or system optimization.

### The Boltzmann Transport Equation

The Boltzmann transport equation (Equation I-1) is a conservation equation for the flux of neutral particles (1: 24). The particles can be neutrons, photons, or any other neutral particle given the nuclear data. The angular flux,  $\psi$ , is dependent on position ( $\mathbf{r}$ ), direction of motion ( $\hat{\Omega}$ ), on speed or energy (v, E), and time (t). Equation (I-1) represents a balance between the loss rate (right side) and gain rate (left side) of particles that exist at each point of this seven dimensional phase space (3: I-2);

$$\begin{bmatrix} \frac{1}{v} \frac{\partial}{\partial t} + \hat{\Omega} \cdot \nabla + \sigma_{t} (\vec{r}, E, t) \end{bmatrix} \psi (\vec{r}, E, \hat{\Omega}, t) =$$

$$\begin{bmatrix} dE' \int d\Omega \sigma_{s} (\vec{r}, E \to E', \hat{\Omega} \cdot \hat{\Omega}, t) \psi (\vec{r}, E', \hat{\Omega}, t) + s(\vec{r}, E, \hat{\Omega}, t) \end{bmatrix}$$

$$(I-1)$$

where the variables are defined in Table 1.

Variable	Description
v	magnitude of the velocity
t	time
r	position
E	energy
$\sigma_t$	total macroscopic cross-section for
	interaction (absorption and scatter)
$\sigma_{s}$	macroscopic scattering cross-section
S	total non-scattering source
Ψ	angular flux: a distribution function
	of particles at point $\vec{r}$ , with energy E,
	moving in direction $\hat{\Omega}$ at time t
Ω	unit vector aligned along the streaming direction of particles: it is often shown as three components, or
	direction cosines, $\mu$ , $\eta$ , $\xi$ defined by
	$\boldsymbol{\mu} = \boldsymbol{\hat{\Omega}} \cdot \boldsymbol{\hat{e}}_{x}, \qquad \boldsymbol{\eta} = \boldsymbol{\hat{\Omega}} \cdot \boldsymbol{\hat{e}}_{y}, \qquad \boldsymbol{\xi} = \boldsymbol{\hat{\Omega}} \cdot \boldsymbol{\hat{e}}_{z}$
	as shown in Figure I-1

Table I-1: Variable Definitions for the BTE, Equation (I-1)

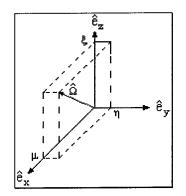


Figure I-1: Direction Cosines

Though discrete ordinates is valid for time dependent problems, this treatment will assume steady state conditions with all time dependence

suppressed; therefore the first term in Equation (I-1) (representing the time rate of change of particles in the phase space) vanishes. The remaining terms represent the steady state balance equation. The second term in the brackets on the left of Equation (I-1) is the streaming operator and represents the loss rate due to particle divergence. The final term in the brackets on the left is the collision operator and represents the loss rate due to particle interaction with the medium. This interaction could be absorption (destroying the particle), or scatter (changing the particle's energy or direction) (3: I-2).

The first term on the right of Equation (I-1) is the gain rate due to particles traveling in other directions that scatter into the given direction,  $\hat{\Omega}$ . As a consequence of the isotropic material assumption, the distribution of particles scattering into a given direction is not a function of the incident angle,  $\hat{\Omega}'$ ; however it is a function of the angle between the incident direction and final direction,  $(\hat{\Omega} \cdot \hat{\Omega}')$ . Despite the isotropic material assumption, this dependence on scattering angle means that scattering may be anisotropic (1). The final term represents the gain rate from production of particles by any source mechanism. The source can be internal, such as radioactive decay and fission or external such as solar x-rays or an incident beam. A more detailed discussion of the BTE and definition of its components is presented in section II.

The macroscopic cross sections are functions of position, through spatially varying number density or changes in material, and of energy

through the microscopic cross sections in which fundamental properties of the isotopes are involved. In general the microscopic cross sections could have some additional implicit spatial variation. For example, this could be due to a temperature-dependence through Doppler broadening (1: 6). This research assumes each material is uniform and isotropic and therefore no such implicit spatial dependence exists.

The physical quantities most often of interest to be found via the BTE are the scalar flux  $\phi$  (zeroth angular moment of the vector flux) and the vector current **J** (first angular moment). The scalar flux is of interest as it represents the total expected particle path length traveled per unit volume at a given location in the medium. This will determine the reaction rates for such things as fission and neutron activation. The vector current will determine the leakage rate from one region to the next or through boundaries (3: I-3).

Only a limited number of analytic and semi-analytic solutions exist for the BTE. Most of these solutions are for highly idealized problems. Many ingenious methods such as discrete ordinates, Monte Carlo, even-parity, finite-elements, and Green functions have been developed to solve the transport equation and to extend the application of such knowledge. The diffusion equation can be derived from the BTE using several simplifying assumptions regarding the angular dependence. For problems with nearly isotropic scatter, this can yield approximate results (16: 2). Of particular

interest are methods capable of providing solutions to the broad range of geometrical configurations found in nuclear reactor and radiation shielding applications (1: 1). From a military perspective, in addition to the application to power generation, accurate modeling of neutron flux allows for more precise modeling of the yield and effects of nuclear weapons.

The advent of high-speed computing and large storage capacity has led to the refinement and use of two primary numerical methods of attaining a solution to the BTE: the method of discrete ordinates and Monte Carlo (1: 2). Other less used methods will not be examined. The theory and development of the Monte Carlo method will not be discussed in detail; however, Monte Carlo solutions will be used as benchmarks.

Since its evolution from the angular segmentation formulation by Carlson in 1957 discrete ordinates has become a widely used method for solving the integrodifferential form of the transport equation. The discrete ordinates method involves enforcing the transport equation only at discrete angular directions called ordinates. These ordinates are selected such that the flux moments may be evaluated accurately by a weighted sum (1: 118). For example the scalar flux (zeroth flux moment) is

$$\phi(\vec{\mathbf{r}},\mathbf{E},\mathbf{t}) = \int d\Omega \psi(\vec{\mathbf{r}},\mathbf{E},\hat{\Omega}\mathbf{t}) \approx \sum_{\mathbf{n}} \mathbf{w}_{\mathbf{n}} \psi(\vec{\mathbf{r}},\mathbf{E},\hat{\Omega}_{\mathbf{n}},\mathbf{t}).$$

A more detailed discussion of the flux moments is found in Chapter II. The advantages of the discrete ordinates method are the relatively simple derivation and the subsequent ease of transformation into algorithms of good computational efficiency (1: 116). It also lends itself well to the discretizing of energy into multiple energy groups. A distinct advantage over the Monte Carlo method is that it provides flux and current data everywhere in the problem rather than only at a limited number of locations. A quadrature set is the combination of discrete angles and weights used in a weighted sum to evaluate the flux moments.

There are two independent angular directions for  $\hat{\Omega}$ . The directions are parameterized by three direction cosines that obey the relationship

$$\mu^2 + \eta^2 + \xi^2 = 1 \tag{I-3}$$

where  $\mu$ ,  $\eta$ , and  $\xi$  are shown in Figure I-1.

Once the angular approximation has been made, a spatial discretization scheme must be used. Computational cost and storage requirements are directly proportional to the number of spatial cells and discrete ordinates used. A large number of spatial schemes have been formulated for use in discrete ordinates calculations. They include linear methods such as diamond difference, linear discontinuous, and linear characteristic and non-linear methods such exponential characteristic (9).

The most serious drawback to the discrete ordinates method is the buildup of truncation errors due to the discretized angular and spatial representations (1: 131). The truncation errors can result in random error, which limits the accuracy of the results, or may lead to physically unrealistic results such as negative fluxes or sources. Systematic truncation error may lead to what are known as ray effects (21, 8, 3). These are errors caused by the discrete ordinates method of limiting particle motion to discrete directions or rays. Flux due to unscattered particles will only be found to occur at points where a line can be drawn from a source to the point in the direction of a discrete ordinate. This causes the scalar flux to be calculated higher than expected at points along discrete ordinate directions and lower between. The method is not well suited to geometry with a strongly peaked flux in a given direction. Generally, a separate method must be used to calculate the first scatter source for such a problem. In order to increase accuracy of the discrete ordinates method and minimize the negative consequences, it is either necessary to increase the number of directions in the angular quadrature, thus increasing the computation time and storage requirements of the computer system, or to develop an alternative quadrature that produces less error with fewer directions. This research concentrates on increasing the accuracy of the discrete ordinates approximation while minimizing the number of angles.

The primary drawback of allowing motion perpendicular to one or more cardinal directions in a quadrature set is that mathematical instability may result when using current computer codes. The resulting zero components of flux often generate run-time errors. The advantages of using directions parallel to cell boundaries are (due to one or two of the direction cosines being zero) the discrete ordinates equations simplify significantly and fewer directions are required for the same order of anisotropy, thus allowing for increased computational efficiency.

Except for the simple case of isotropic scatter, the cross section for scattering will be a function of the scattering angle as well as energy. Separation of the angular and energy dependence is assumed. Traditionally, cross sections are then expanded in orthogonal Legendre polynomials (1: 13). The order of this expansion is another limit to the accuracy obtainable by the discrete ordinates method.

The high performance computers needed to perform these calculations are very expensive, and therefore anything that can increase the efficiency of these machines will translate directly into substantial savings by increasing their useful life. Reduction of computation time also increases the practicality of analyzing a large variety of similar scenarios for threat analysis or system optimization.

### Statement of the Problem

The objective of this research is to develop and evaluate new quadrature sets that produce accurate discrete ordinates solutions to the BTE using a minimum number of ordinates. The viability of including ordinates perpendicular to one or two cardinal directions is examined. This includes deriving and implementing the appropriate quadrature angles and weights, and comparing the results with those obtained with standard levelsymmetric LQ<sub>n</sub> quadratures and with Monte Carlo benchmark solutions.

### Scope

This research includes the derivation and implementation of discrete ordinates quadrature sets that include directions perpendicular to one or two cardinal directions. Demonstration of the method including comparison to traditional level-symmetric quadratures with regard to computational cost (execution time) and accuracy of results based on a benchmark calculation is performed. The test problems use three-dimensional Cartesian coordinates with no time or energy dependence. The test problems were run using TETRAN (13), an unstructured mesh tetrahedral cell code developed at the Air Force Institute of Technology and the THREEDANT code of the RSICC Computer Code Collection, DANTSYS 3.0 (14) from Los Alamos National Laboratory (LANL) using a rectangular parallelepiped mesh. LANL's MCNP (Monte Carlo Neutron Photon) transport code package (15) provided

benchmark solutions. Due to current limitations in the TETRAN code still in development, the test problems are defined as one energy group, isotropic scatter transport problems. Multiple levels of spatial mesh refinement are used. The method is tested to identify any variation in performance and determine an optimal usage. The scope of the comparison using the output of the THREEDANT module is limited due to the requirement to modify the developed quadratures in order for the module to run. See chapter IV for a discussion of the modifications. No code changes ore new modules were written to augment THREEDANT to obtain a more accurate comparison of the quadrature sets.

### General Approach and Sequence of Presentation

In chapter II the integrodifferential form of the Boltzmann transport equations is discretized over angle. A brief discussion of spatial and energy discretization is included. The consequences of discretization are enumerated. The method of generating the new quadrature sets is developed in chapter III. Several quadrature sets of various order are presented. The method is implemented using two test problems and the results are presented in chapter IV. The geometry of each test problem has been selected to, in the first case, exacerbate, then in the second, mitigate the problem of ray effects. Traditional level symmetric quadratures are used on the same test problems. Benchmark calculations were performed on each

test problem using a Monte Carlo simulation. The methods are compared for accuracy and computational efficiency and potential advantages or disadvantages of the method identified. Consideration is given to smoothness, pointwise and global accuracy, ray-effects, and other systematic errors.

Once the method has been tested and analyzed, recommendations for use and for further research are given in the final chapter. Appendices contain complete derivations of the equations used to generate the quadratures as well as any mathematical routines used to solve them. Also, pertinent portions of input and output files of the test problems are included

## II. Theory

This chapter will present a development of the discrete ordinates method, the criteria for selecting a quadrature, discuss spatial discretization, and some the consequences of applying these approximations.

The steady state assumptions reduces the Boltzmann transport equation (Equation I-1) to:

$$\begin{split} \hat{\Omega} \cdot \vec{\nabla} \psi \left( \vec{\mathbf{r}}, \mathbf{E}, \hat{\Omega} \right) + \sigma(\vec{\mathbf{r}}, \mathbf{E}) \psi \left( \vec{\mathbf{r}}, \mathbf{E}, \hat{\Omega} \right) = \\ \int d\mathbf{E}' \int d\Omega' \sigma_s \left( \vec{\mathbf{r}}, \mathbf{E} \to \mathbf{E}', \hat{\Omega} \cdot \hat{\Omega}' \right) \psi \left( \vec{\mathbf{r}}, \mathbf{E}', \hat{\Omega} \right) + s \left( \vec{\mathbf{r}}, \mathbf{E}, \hat{\Omega} \right) \end{split}$$
(II-1)

In order to yield a convenient normalization over all angles, the incremental solid angle is defined as

$$d\Omega = \frac{d\omega}{2\pi} \frac{d\theta \sin \theta}{2} = \frac{d\omega}{2\pi} \frac{d\mu}{2}$$
(II-2)

so that

$$\int d\Omega = \int_{0}^{2\pi} \frac{d\omega}{2\pi} \int_{-1}^{1} \frac{d\mu}{2} = 1.$$
 (II-3)

ç

Definitions of the above angles are shown in Figure II-1 (1: 11). Equation (II-1) contains terms that are functions of position, scattering angle, and energy. The energy dependence in the discrete ordinates approximation is most often accounted for by dividing the energy range of interest into a number of intervals. It is assumed that for each interval, cross-sections are given as average values over the interval (18: 109). The transport equation is then

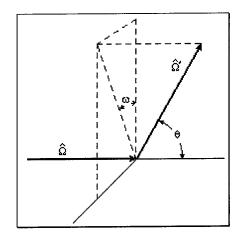


Figure II-1: Particle Entering from Direction  $\hat{\Omega}$ , Scattering into direction  $\hat{\Omega}'$ solved in each energy interval as a mono-energetic equation with particle contributions scattered from outside the energy interval added as a source term and particles scattering from the interval of interest into other energy intervals treated as losses. The resulting equations are known as the *multigroup equations* (1: 61). For clarity, the remainder of the derivations given will be for a mono-energetic system. The Boltzmann equation then becomes the mono-energetic transport equation:

$$\hat{\Omega} \cdot \vec{\nabla} \psi \left( \vec{\mathbf{r}}, \hat{\Omega} \right) + \sigma(\vec{\mathbf{r}}) \psi \left( \vec{\mathbf{r}}, \hat{\Omega} \right) = \int d\hat{\Omega}' \sigma_{s} \left( \vec{\mathbf{r}}, \hat{\Omega} \cdot \hat{\Omega}' \right) \ \psi \left( \vec{\mathbf{r}}, \hat{\Omega}' \right) + s \left( \vec{\mathbf{r}}, \hat{\Omega} \right).$$
(II-4)

In the discrete ordinates method each term of the integrodifferential transport equation is assumed to be a separable function of space and angle and the dependencies are dealt with separately (3: I-3). Both the spatial and angular variables are required to be discretized before the problem may be solved numerically. The angular discretization, being the focus of this research, will be considered first and discussed in some detail while the spatial discretization will be dealt with later.

#### Angular Discretization

The majority of the derivation that follows is an extension into threedimensions of the procedure presented by Lewis and Miller (1). The differential scattering cross sections are expanded in orthogonal Legendre polynomials  $P_1(\hat{\Omega} \cdot \hat{\Omega}')$  where  $\mu_0 = \hat{\Omega} \cdot \hat{\Omega}'$  is the cosine of the scattering angle. The differential scattering cross section may be expressed as

$$\sigma_{\rm s}\left(\vec{\mathbf{r}},\hat{\boldsymbol{\Omega}}\cdot\hat{\boldsymbol{\Omega}}'\right) \approx \sum_{l=0}^{\rm L} (2l+1)\sigma_{\rm sl}(\vec{\mathbf{r}}) P_l\left(\hat{\boldsymbol{\Omega}}\cdot\hat{\boldsymbol{\Omega}}'\right). \tag{II-5}$$

The scattering moments  $\sigma_{\rm sl}$  are found from

$$\sigma_{\rm sl}(\vec{\mathbf{r}}) = \int_{-1}^{1} \frac{\mathrm{d}\,\mu_{\rm o}}{2} \sigma_{\rm s}(\vec{\mathbf{r}},\mu_{\rm o}) P_{\rm l}(\mu_{\rm o}), \tag{II-6}$$

having taken advantage of the orthogonality property of the Legendre polynomials. The expansion in Equation (II-5) has been truncated at L+1 terms, assuming and adequate level of approximation, taking into consideration the degree of anisotropy and the availability of cross section data. For the case of isotropic scatter only the first term of Equation (II-5) is used (L = 0) resulting in

$$\sigma_{\rm s}(\vec{\mathbf{r}}, \hat{\Omega} \cdot \hat{\Omega}') \cong \sigma_{\rm s}(\vec{\mathbf{r}}). \tag{II-7}$$

This derivation will not assume isotropic scatter. Combining Equations (II-4) and (II-5) yields

$$\hat{\Omega} \cdot \vec{\nabla} \psi \left( \vec{\mathbf{r}}, \hat{\Omega} \right) + \sigma(\vec{\mathbf{r}}) \psi \left( \vec{\mathbf{r}}, \hat{\Omega} \right) = \sum_{l=0}^{L} (2l+1) \sigma_{sl}(\vec{\mathbf{r}}) \int d\Omega' P_l \left( \hat{\Omega} \cdot \hat{\Omega}' \right) \psi \left( \vec{\mathbf{r}}, \hat{\Omega}' \right) + s \left( \vec{\mathbf{r}}, \hat{\Omega} \right)$$
(II-8)

which can be simplified using the Legendre addition theorem (1: 367)

$$\mathbf{P}_{1}\left(\hat{\boldsymbol{\Omega}}\cdot\hat{\boldsymbol{\Omega}}'\right) = \frac{1}{2l+1} \sum_{m=-1}^{l} \mathbf{Y}_{lm}^{*}\left(\hat{\boldsymbol{\Omega}}\right) \mathbf{Y}_{lm}\left(\hat{\boldsymbol{\Omega}}'\right) \tag{II-9}$$

where the  $Y_{lm}(\hat{\Omega})$  are the spherical harmonics and the asterisks signifies the complex conjugate. Using this, the sum on the right hand side of Equation (II-8) becomes

$$\sum_{l=0}^{L} \sigma_{sl}(\vec{r}) \sum_{m=-l}^{l} Y_{lm}^{*}(\hat{\Omega}) \int d\Omega' Y_{lm}(\hat{\Omega}') \psi(\vec{r},\hat{\Omega}').$$
(II-10)

The angular integral in expression (II-10) is now just the coefficients resulting from the expansion of the angular flux in spherical harmonics

$$\psi(\vec{\mathbf{r}}, \hat{\boldsymbol{\Omega}}) \approx \sum_{l=0}^{L} \sum_{m=-l}^{l} \phi_{lm}(\vec{\mathbf{r}}) Y_{lm}^{*}(\hat{\boldsymbol{\Omega}}).$$
(II-11)

with

$$\phi_{lm}(\vec{\mathbf{r}}) = \int d\Omega' Y_{lm}(\hat{\Omega}') \ \psi(\vec{\mathbf{r}}, \hat{\Omega}'). \tag{II-12}$$

Substituting this into expression (II-10) gives

$$\sum_{l=0}^{L} \sum_{m=-1}^{l} Y_{lm}^{*}(\hat{\Omega}) \sigma_{sl}(\vec{r}) \phi_{lm}(\vec{r}).$$
(II-13)

Replacing the sum in Equation (II-8) with expression (II-13) yields a form of the transport equation that is convenient for discretization in angle

$$\hat{\Omega} \cdot \vec{\nabla} \psi \left( \vec{\mathbf{r}}, \hat{\Omega} \right) + \sigma(\vec{\mathbf{r}}) \psi \left( \vec{\mathbf{r}}, \hat{\Omega} \right) = \sum_{l=0}^{L} \sum_{m=-1}^{l} Y_{lm}^{*} \left( \hat{\Omega} \right) \sigma_{sl}(\vec{\mathbf{r}}) \phi_{lm}(\vec{\mathbf{r}}) + s(\vec{\mathbf{r}}, \hat{\Omega}). \quad (II-14)$$

In the discrete ordinates approximation, Equation (II-14) is enforced only for a set of discrete directions yielding

$$\hat{\Omega}_{n} \cdot \vec{\nabla} \psi_{n}(\vec{\mathbf{r}}) + \sigma(\vec{\mathbf{r}}) \psi_{n}(\vec{\mathbf{r}}) = \sum_{l=0}^{L} \sum_{m=-l}^{l} Y_{lm}^{*}(\hat{\Omega}_{n}) \sigma_{sl}(\vec{\mathbf{r}}) \phi_{lm}(\vec{\mathbf{r}}) + s(\vec{\mathbf{r}}, \hat{\Omega}_{n}), \quad (\text{II-15})$$

where  $\psi_n(\vec{r}) \equiv \psi(\vec{r}, \hat{\Omega}_n)$ . The scalar flux is approximated by

$$\phi(\vec{\mathbf{r}}) = \sum_{n=1}^{N} \mathbf{w}_n \psi_n(\vec{\mathbf{r}}), \qquad (\text{II-16})$$

and the flux moments of Equation (II-12) are approximated by

$$\phi_{lm}(\vec{\mathbf{r}}) = \sum_{n=1}^{N} \mathbf{w}_n \mathbf{Y}_{lm}(\hat{\boldsymbol{\Omega}}_n) \boldsymbol{\psi}_n(\vec{\mathbf{r}}).$$
(II-17)

By defining the right hand side of Equation (II-15) as the emission density or scattering source  $q_n(\vec{r})$  and substituting Equation (II-17) for the flux moments, Equation (II-15) becomes

$$\hat{\Omega}_{n} \cdot \vec{\nabla} \psi_{n}(\vec{r}) + \sigma(\vec{r}) \psi_{n}(\vec{r}) = q_{n}(\vec{r})$$
(II-18)

where

$$q_{n}(\vec{r}) = s(\vec{r}, \hat{\Omega}_{n}) + \sum_{l=0}^{L} \sum_{m=-l}^{l} Y_{lm}^{*}(\hat{\Omega}_{n}) \sigma_{sl}(\vec{r}) \sum_{n'=1}^{N} w_{n'} Y_{lm}(\hat{\Omega}_{n'}) \psi_{n'}(\vec{r}) \quad (II-19)$$

The most common way of solving Equation (II-18) is the *iteration on the* scattering source form of Von Neuman's series solution (1: 80). The method's usefulness is derived from the fact that if the right side of Equation (II-18) is known, solving for  $\psi(\mathbf{\bar{r}})$  is usually straightforward. The iteration is defined by modifying Equation (II-18) to

$$\left[\hat{\Omega}_{n}\cdot\vec{\nabla}+\sigma(\vec{r})\right]\psi_{n}^{i+1}(\vec{r})=q_{n}^{i}(\vec{r})$$
(II-20)

where i is the iteration index. Equation (II-19) likewise becomes

$$q_{n}^{i}(\vec{r}) = s(\vec{r}, \hat{\Omega}_{n}) + \sum_{l=0}^{L} \sum_{m=-1}^{l} Y_{lm}^{*}(\hat{\Omega}_{n}) \sigma_{sl}(\vec{r}) \sum_{n'=1}^{N} w_{n'} Y_{lm}(\hat{\Omega}_{n'}) \psi_{n'}^{i}(\vec{r}).$$
(II-21)

The system of Equations (II-20) and (II-21) must be solved to convergence at each  $\vec{r}$  of interest.

Before proceeding to the methodology of choosing a quadrature set, a brief discussion of the spatial discretation which allows for the approximation of the those terms which are now function of the spatial variable only will be helpful.

# **Spatial Discretization**

Discretization of the spatial variable in three dimensions can take many forms. The most common method is to divide each of the three Cartesian spatial directions, x, y, and z, into i, j, and k intervals respectively resulting in a three dimensional grid containing  $i \times j \times k$  rectangular parallelepiped cells. Another method gaining popularity is to generate an unstructured mesh of tetrahedra. All cross sections are taken to be piecewise constant and therefore not allowed to vary inside a given cell. Various methods exist then to calculate the angular flux  $\psi_n$  for each value of *n* in each cell given appropriate boundary conditions. THREEDANT uses diamond difference (DD) with negative flux fix up in a structured Cartesian mesh, and TETRAN has the capability of using linear characteristic (LC), exponential characteristic (EC) or step characteristic (SC) on an unstructured tetrahedral mesh. Each of these methods and many more are discussed and derived in detail elsewhere (1,2,3,8,9,16).

## **Consequences of Discretization**

Several problems arise when using discrete ordinates. Any time computations are performed on a computer, truncation errors will occur. The effects of truncation error will accumulate with each computation. Truncation errors therefore limit the accuracy achievable by computational methods. With each refinement of the spatial or angular mesh, the mathematical model more closely resembles the continuous analytical solution, however the number of computations also increases. Though truncation error associated with each calculation should decreases with mesh refinement, there is a limit to the accuracy gained by continued mesh refinement as the number of calculations gets increasingly large. This effect is most visible in three-dimensional problems where the number of cells increases as the cube of the linear refinement. Another, more systematic problem arises due to the angular discretization called ray effects. The phenomenon is most evident in problems with localized sources and small scattering cross sections (1:195).

When the scattering cross section is small, a substantial percentage of particles traveling from a localized source to an area of interest will be uncollided. This will result in a peaked distribution about the discrete ordinates. Clearly these results are physically unrealistic. Ray effects

II-9

manifest as oscillations (1: 197), peaks about a region where an ordinate can be traced from a source region and valleys in between. As the order of the angular approximation increases so does the number of oscillations, which also tend to decrease in the amount of deviation from a smooth curve. If the oscillations are uniform about the correct curve, then the integral of the scalar flux over the boundary will still yield good results (1: 200). When the scattering source makes a large contribution to the scalar flux, ray effects tend to be mitigated. The scattering source tends to be distributed over a large area and is often nearly isotropic. This gives a more uniform angular distribution of neutrons.

Numerical diffusion is a consequence of the spatial discretization and truncation errors due to spatial differencing. For example, if a beam of neutrons were to enter the lower left corner of a pure absorbing cube of material traveling along the cube diagonal, one would expect the attenuated beam to exit only at the upper right corner. In the spatial walk of the discrete ordinates calculation, each cell is considered to have a distribution of flux through out the cell and flux can only enter and exit the faces of the cells. For a mesh consisting of regular parallelepiped cells, this means the beam entering the cell in the bottom left corner will exit that cell through the top and sides. The fraction entering the adjacent cells is determined by the incident angle of the beam. This will continue through the spatial walk resulting in a smearing out of the beam. This effect, to a small extent, may

II-10

mitigate ray-effects. A more complete discussion of numerical diffusion can be found in reference 3 where is referred to as *quasi-ray effect*.

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# III. Method

The primary desire when selecting a quadrature set is to maximize accuracy while minimizing the number of ordinates. The primary source of computational cost of obtaining a discrete ordinates solution to the BTE is in the spatial integration. A complete walk through the spatial mesh is required for each ordinate. Thus minimizing the number of ordinates reduces the computational cost by minimizing the number spatial walks required. Other concerns are the mitigation of numerical artifacts both systematic and nonsystematic such as truncation errors, ray effects, and numerical diffusion. With these concerns in mind, this work will present a method of generating simultaneous sets of polynomial equations, which produce quadrature sets based on exact integration of spherical harmonics. Quadrature sets up to order seven are presented.

#### Zero Components

The key difference in the quadratures presented here from those seen elsewhere is the addition of ordinates with zero components for one or two of the direction cosines. There have been several reasons for not using zero components in the past, primarily stemming from arguments in one or twodimensional geometry. In these problems, vertical vector flux does not propagate through the problem. The infinite path length resulting from the reciprocal cosine term in the spatially discretized mesh must also be dealt

III-1

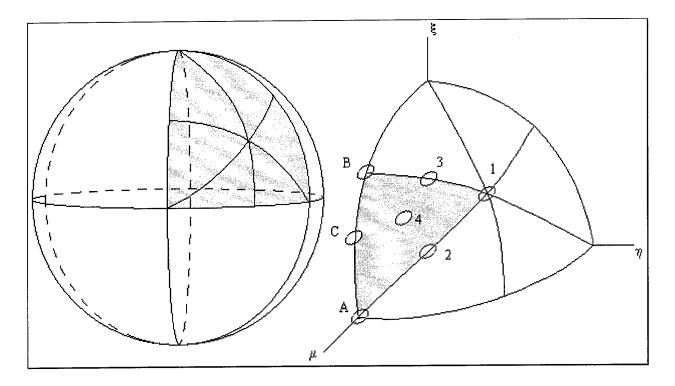
with. Discontinuity in the vector flux at a vacuum boundary can lead to ambiguity in the meaning of  $\psi(\mu = 0)$  at the boundary when using a regular Cartesian mesh. Also, it can be shown that  $\psi(\mu = 0)$  is not truly an independent variable when discretized in 2-D and therefore contributes little to the accuracy of the solution (19: 132).

When three-dimensional quadratures are developed, they are often an extension of one and two-dimensional cases. The use of even order in traditional quadrature sets, when applied to 3-D problems has invariably been due to the even order used in the lower dimensional base. The use of diamond difference (DD) is also pervasive in old computer codes. DD codes will not accept zero components without special handling routines being developed. There is often a desire when developing quadrature sets to generate results that will run on the legacy codes without modification. This is unfortunate, as modern parallel computers are capable of handling far greater complexities than the computers for which these codes were developed. The development of computer codes capable of performing discrete ordinates calculations on an unstructured mesh eliminates many of the problems associated with special directions. More accurate characteristic methods of dealing with the spatial integration do not have the same problems dealing with zero components as the DD method.

III-2

#### **Quadrature Derivation**

The derivation of these polynomial equations for the generation of quadrature sets is based on methodology developed by Dr. Kirk Mathews (25) as an extension of work by B.G. Carlson and others (5). The basic quadrature sets presented here are defined over the entire unit sphere, but can be represented in the principal octant or its edges. The principal octant is where the components of  $\hat{\Omega}$  are all positive. We require the quadrature set to meet the total symmetry condition. Total symmetry, sometimes referred to as cubic symmetry, requires the quadrature set to remain invariant under all axis exchanges, ninety-degree rotations about a cardinal axis, and reflections across the x - y, x - z, or y - z planes. An axis exchange operation is the same as a reflection across any x = y, x = -y, x = z, x = -z, y = z, or y = -z plane. The discrete ordinates can be represented as points on the surface of the unit sphere. These points represent where the tips of the unit vectors corresponding to each ordinate lie on the unit sphere. A base set refers to those ordinates in the original hextant of the principal octant. The original hextant is a spherical triangle covering a one sixth area of the principal octant. Figure III-1 shows the unit sphere with the principle octant shaded and an expanded view of the principle octant with the original hextant shaded. A complete quadrature set can be built by choosing points in any hextant and reflecting the points by performing successive axes exchange





# Cases and Designators

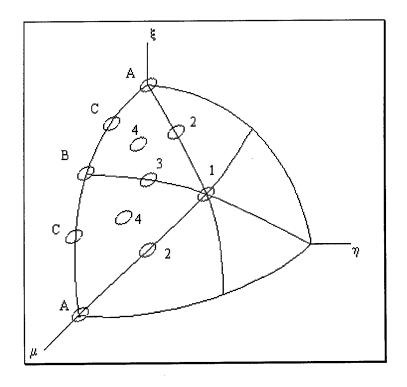


Figure III-2:  $\mu \leftrightarrow \xi$  Exchange Operation

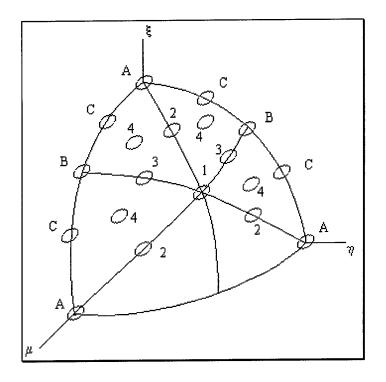


Figure III-3:  $\mu \leftrightarrow \eta$  Exchange Operation

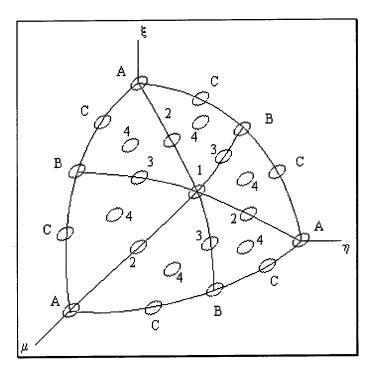


Figure III-4: Complete Principal Octant after  $\, \xi \leftrightarrow \eta \,$  Exchange Operation

operations as shown in Figures III-2 through 4 (25). The remaining ordinates are generated by sequential reflections across the x-y, x-z and y-z planes. The directions  $\hat{\Omega}_n$  constitute a discrete set of values of  $\hat{\Omega}$  over its entire domain (6: 2). If  $\psi$  has a convergent expansion in spherical harmonics, then from Equations (II-15) and (II-17) we see that a necessary condition for choosing the ordinates and weights is that the spherical harmonic orthogonality condition up to the desired order be satisfied, that is

$$\int d\Omega Y_{lm}(\hat{\Omega}) Y_{l'm'}(\hat{\Omega}) = \sum_{n=1}^{N} w_n Y_{lm}(\hat{\Omega}_n) Y_{l'm'}(\hat{\Omega}_n) = \delta_{ll'} \delta_{mm'}$$
(III-1)

for all  $0 \le l, l' \le L$ ,  $-l \le m \le l$ , and  $-l' \le m' \le l'$  with the desired order of precision L. The l = l' = m = m' = 0 case provides the normalization condition for the weights,

$$\sum_{n}^{N} w_{n} = 1.$$
 (III-2)

We also require that  $w_n > 0$  to reduce to possibility of obtaining physically unrealistic results.

Because of the symmetry requirements and the exchangeability of coordinates, Equation (III-1) will be satisfied if the moments equations,

$$\int \mu^{i} d\Omega = \sum_{n=1}^{N} w_{n} \mu_{n}^{i} , \qquad i = 0, 1, ..., 2L$$
 (III-3)

are satisfied exactly. In Equation (III-3) i is an exponent not a superscript. This condition is enough to ensure that  $\int \mu^i \eta^j \xi^k d\Omega$  will integrate exactly for all  $0 \le i, j, k \le 2L$  with  $i + j + k \le 2L$  (25). Also due to symmetry, if i, j, or k is odd, the integral is exactly zero. This is because the integration in Equation (III-1) can be represented as the integration of a product Legendre polynomials, which are intern polynomials of  $\mu = \cos(\theta)$ , and  $e^{i(m-m')\phi}$ .

Symmetry guarantees the exponential will ingrate exactly because for every  $\varphi_n$  there exists with equal weight and equal  $\mu_n$  a  $\varphi_{n'}$  that equals  $\varphi_n + \pi$ . The terms in the summation therefore cancel except when m = m', which results in the exponential reducing to unity. The odd powered terms are exactly zero because cosine integrates to zero over the range of -1 to 1.

Finally, the case where i = 0 and the case where i = 2 are not independent. For the i = 0 case, Equation (III-3) becomes

$$\int d\Omega = 1 = \sum_{n=1}^{N} w_n , \qquad (III-4)$$

and for the i = 2 case

$$\int \mu^2 d\Omega = \frac{1}{3} = \sum_{n=1}^{N} w_n \mu_n^2 \quad . \tag{III-5}$$

Because of the symmetry requirements, the  $\mu_n^2$  values with equal weights will always exist in sets of three that sum to one (this can readily be seen by examining the most general case (case 4) in appendix B with k = 1). Equation (III-5) therefore becomes

$$\frac{1}{3} = \sum_{n=1}^{N} \frac{1}{3} w_n \quad . \tag{III-6}$$

Equation (III-6) and Equation (III-3) are not linearly independent. We can therefore replace the system of Equations (III-3) with

$$\int \mu^{2k} d\Omega = \frac{1}{2k+1} = \sum_{n=1}^{N} w_n \mu_n^{2k} , \quad k = 1, 2, ..., L.$$
(III-7)

The system of Equations (III-7) may have a solution or solutions providing the number of equations, L, is equal to the total number of free parameters (degrees of freedom): the total number of independent values for  $w_n$  and  $\mu_n$ . It is important to make clear that the criteria used for the selection of a quadrature set is that it will evaluate, without quadrature error, integrals of the sort

$$\int Y_{lm}^{*}(\hat{\Omega})f(\hat{\Omega})d\Omega$$

for any  $N \ge l$ ; which means

$$\int Y_{lm}^{*}(\hat{\Omega}) Y_{l'm'}(\hat{\Omega}) d\Omega$$

must integrate with out error. This is why Equation (III-3) requires  $\mu^i$  to integrate exactly for i = 0 to 2L. Symmetry assures getting all the odd powers exactly while the degrees of freedom in the weights and angles are used to get the even powers. Thus five degrees of freedom integrate the coefficients of a fifth order expansion. P<sub>5</sub> anisotropic scatter needs this order of expansion. Chapter II discusses the expansion of the scattering cross section. With highly anisotropic scattering, a high order expansion may be needed to get accurate results. This requires the discrete ordinates order to be at least as high. Computational efficiency with highly anisotropic scattering therefore becomes of even greater concern. Therefore, obtaining the desired order of expansion with fewer ordinates is very desirable.

The objective is then to use the most reasonable value of L and minimize N, maintain total symmetry, and produce accurate transport results. The desired (or available) precision of the arithmetic, computation cost, and the order of the Legendre expansion of the scattering cross section will determine the most reasonable value of L. Figure III-4 shows the principle octant and the various possible cases for points to be located and satisfy all symmetry requirements. The four cases not on the edge of the octant correspond with the cases 1 through 4 presented by Carlson (5) and have 1, 3, 3, and 6 points in the octant (hence 8, 24, 24, 48 points on the unit sphere) with 1, 2, 2, and 3 degrees of freedom respectively. The three additional cases, A, B, and C, are shared by two or more octants. Case A points are on a primary axis. Case B points are in a zero plane with the remaining two components equal. Case C points are in a total of 6, 12, and 24 points over the unit sphere with 1, 1, and 2 degrees of freedom for cases A, B, and C respectively. Table III-2 provides a summary of pertinent information regarding each case.

Table III-1: Summary c	f Quadrature Ca	ase Data
------------------------	-----------------	----------

Case	Total Points on	Degrees of	Parameters to	Ordinates per
	Unit Sphere	Freedom	be Determined	Degree of Freedom
1	8	1	W	8
2	24	2	w,µ	12
3	24	2	w,µ	12
4	48	3	w,μ,η	16
A	6	1	W	6
В	12	1	w	12
С	24	2	w,µ	12

Cases A and 1 have the fewest ordinates per degree of freedom and are therefore preferred. Case 4 is the most expensive and also the most difficult to calculate and is to be avoided. The remaining cases are equivalent in this regard.

To distinguish these quadrature sets from those previously developed the notation  $MQ_n$  is adopted for a quadrature of order n. To uniquely solve for a quadrature set of order n, the total degrees of freedom in the equation set used must also equal n. This is accomplished by selecting a combination of cases from Table III-1 such that the sum of the degrees of freedom equals n. Equation (III-7) is then used to determine the free parameters. Care must be taken to uniquely identify the quadrature being referenced because there may be multiple quadrature sets of the same order. For example, to find an  $MQ_5$  quadrature case 1 + case 2 + case A + case B will provide the needed degrees of freedom. A different  $MQ_5$  quadrature results from selecting case 4 + case C. Further discussion of notation is presented later.

When selecting the cases, those cases with no free angles may only be selected once; otherwise a case may be used multiple times. Each free parameter must be uniquely identified. For example, for case 3 + case 3 + case A, the free parameters would be  $\{(\mathbf{w}_3, \mu_3, \eta_3), (, \mathbf{w}_{3'}, \mu_{3'}, \eta_{3'}), (\mathbf{w}_A)\}$ resulting in an MQ<sub>7</sub> quadrature.

Because for each case only one point on the sphere is independent, the others resulting from exchange and reflection operations, we are able to write a set of equations for each case that depends only on the free parameters of that point. This series of equations is then inserted into the system of Equations (III-7). For case 1,  $\mu_1 = \eta_1 = \xi_1 = \frac{1}{\sqrt{3}}$  in the primary octant with  $w_1$  to be determined and a total of eight points over the sphere. Therefore, the contribution of case 1 to Equation (III-7) is

$$\mathbf{w}_{1} \left[ \left( \frac{1}{\sqrt{3}} \right)^{2k} + \left( \frac{1}{\sqrt{3}} \right)^{2k} + \left( \frac{1}{\sqrt{3}} \right)^{2k} + \left( \frac{1}{\sqrt{3}} \right)^{2k} + \left( -\frac{1}{\sqrt{3}} \right)^{2k} + \left( -\frac{1}{\sqrt{3}} \right)^{2k} + \left( -\frac{1}{\sqrt{3}} \right)^{2k} \right] = 8 \mathbf{w}_{1} \left( \frac{1}{3} \right)^{k}$$

$$\mathbf{k} = 1, 2, \dots, \mathbf{L}.$$
(III-8)

The contributions of the other cases are found analogously. A complete derivation of each case is given in Appendix A and is summarized in Table III-2. Note that cases 2 and 3 have the same form of contribution equation. If the angle solved for is less than  $\frac{1}{\sqrt{3}}$  then this is a case 2 ordinate and the angle is  $\mu$ , if the angle is greater than  $\frac{1}{\sqrt{3}}$  then a case 3 ordinate results and the angle is  $\eta$ . The column labeled "Cosines" gives the equations or values needed to find the initial ( $\mu$ ,  $\eta$ ,  $\xi$ ) triplet. The rest of the quadrature is then generated by the operations described previously. As an example, for the

Case	Contribution	Cosines $(\mu, \eta, \xi)$
1	$8w_1\left(\frac{1}{3}\right)^k$	$\left(\frac{1}{\sqrt{3}},\frac{1}{\sqrt{3}},\frac{1}{\sqrt{3}}\right)$
2	$8w_{2}\left[\mu_{2}^{2k} + 2\left(\frac{1-\mu_{2}^{2}}{2}\right)^{k}\right]$	$\left(\mu_{2}, \sqrt{\frac{1-\mu_{2}^{2}}{2}}, \sqrt{\frac{1-\mu_{2}^{2}}{2}}\right)$
3	$8w_{3}\left[\eta_{3}^{2k} + 2\left(\frac{1-\eta_{3}^{2}}{2}\right)^{k}\right]$	$\left(\sqrt{\frac{1-{\eta_3}^2}{2}}, \ \eta_3, \ \sqrt{\frac{1-{\eta_3}^2}{2}}\right)$
4	$16w_{4}\left[\mu_{4}^{2k}+\left(1-\mu_{4}^{2}-\eta_{4}^{2}\right)^{k}+\eta_{4}^{2k}\right]$	$\left(\mu_{4}, \eta_{4}, \sqrt{1-{\mu_{4}}^{2}-{\eta_{4}}^{2}}\right)$
Α	2w <sub>A</sub>	(1,0,0)
В	$8w_B\left(\frac{1}{2}\right)^k$	$\left(\frac{1}{\sqrt{2}}, 0, \frac{1}{\sqrt{2}}\right)$
С	$8\mathbf{w}_{C}\left[\mu_{C}^{2k}+\left(1-\mu_{C}^{2}\right)^{k}\right]$	$\left(\mu_{\rm C}, 0, \sqrt{1-\mu_{\rm C}^2}\right)$

Table III-2: Contribution of Each Case to Equation (III-7)

 $MQ_5$  quadrature using case 1 + case 2 + case A + case B, the following equations result

$$2w_{A} + 8w_{B} \left(\frac{1}{2}\right)^{k} + 8w_{1} \left(\frac{1}{3}\right)^{k} + 8w_{2} \left[\mu_{2}^{2k} + 2\left(\frac{1-\mu_{2}^{2}}{2}\right)^{k}\right],$$
  
k = 1, 2, 3, 4, 5. (III-9)

These equations were entered into a Mathematica<sup>™</sup> (20) notebook and solved using built-in functions. A listing of the notebook is found in Appendix B. The solution for Equations (III-9) is shown in Table III-3. Including all unique permutations of the cosines will complete the principal octant. For example, there are two additional case A ordinates with the same weight as the one shown and having cosines (0, 1, 0) and (0, 0, 1) respectively. There are no additional case 1 ordinates as there are no more unique permutations of the cosines. Figure III-3 shows the distribution of these points on the principal octant. Both the exact solution and the decimal fractions of the weights are given. The weights are all positive, as required, which reduces the likelihood of physically unrealistic results such as negative flux. They are also similar in magnitude, providing for better conditioning of the problem. From Table III-1 we see that there are fifty total directions and five degrees of freedom for this quadrature. For comparison, an LQ10 quadrature also has five degrees of freedom but has 120 total directions over the unit sphere requiring 2.4 times the computational cost (assuming both calculations converge in the same number of iterations). The LQ<sub>6</sub> level symmetric quadrature has 48 directions but only three degrees of freedom. Table III-3

also gives the base sets for an  $MQ_3$  and an  $MQ_7$  quadrature set. Table III-4 shows the number of discrete ordinates required to obtain a desired number of degrees of freedom for various  $MQ_n$  and  $LQ_n$  quadrature sets.

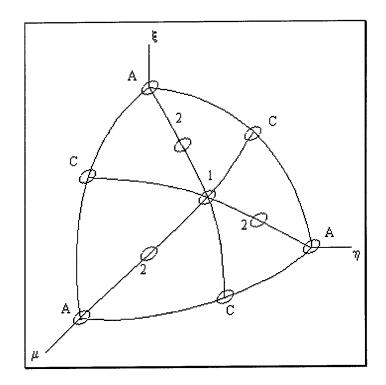


Figure III-5: Example MQ<sub>5</sub> Quadrature Layout

Order	Class	weight	μ	η	ξ
n = 3	А	1/21 .047619047619	1	0	0
	В	9/280 .0321428571429	$1/\sqrt{2}$	0	1/√2
	1	4/105 .0380952380952	1/√3	1/√3	1/√3
n = 5	А	4 / 315 .0126984126984	1	0	0
	В	64 / 2835 .0225749559083	$1/\sqrt{2}$	0	1/√2
	1	27 / 1280 .0210937500000	1/√3	1/√3	1/√3
	2	14641 / 725760 .020173335379	3/√11	1/√11	1/ √11
n = 7	Α	.00904818883016	1	0	0
	В	.02103246043743	$1/\sqrt{2}$	0	$1/\sqrt{2}$
	С	.00645149153857	.954580866940172	0	0.29795195665031
	1	.01827941392342	1/√3	1/√3	1/√3
	2	.01634375972737	.875317087598172	.34192104070871	.34192104070871

Table III-3: Some Possible MQn Quadrature Base Sets

Table III-4: Comparison of Computational Cost of  $MQ_n$  versus  $LQ_n$ 

Degrees of Freedom	MQn Cases Used	LQn Required	MQ <sub>n</sub> : Total Number of Ordinates on Unit Sphere	LQ <sub>n</sub> : Total Number of Ordinates on Unit Sphere
1	Α	$\mathbf{S}_2$	6	8
3	A, B, 1	$S_6$	26	48
5	A, B, 1, 2	$S_{10}$	50	120
7	A, B, C, 1, 2	${ m S}_{12}$ / ${ m S}_{14}$ *	74	168 / 224
9	A, B, 1, 2, 4	$S_{16}$	110	288

 $^*S_{12}$  only has six degrees of freedom and  $S_{14}$  has eight. No level symmetric  $S_n$  quadrature has 7 degrees of freedom.

For the MQ7 quadrature in Table III-3, the exact values are not given,

because Mathematica was not able to solve this set of equations exactly.

Instead a numerical solutions was obtained using a numerical solving built-

in, NSolve. As the order of the quadrature increases the complexity of the polynomial system of equations increases substantially. Only for the lowest order quadrature sets was I able to obtain exact solutions.

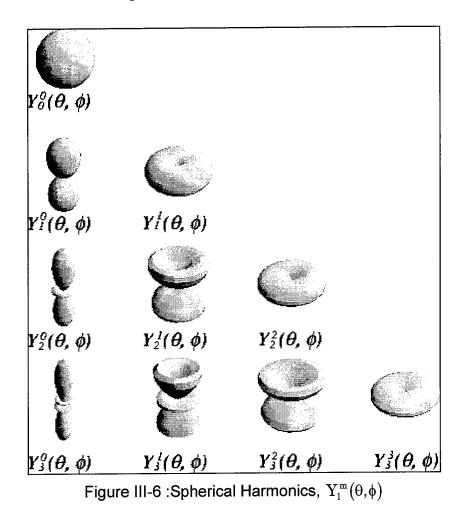
These quadrature sets were tested with a computer algorithm to verify they met the orthogonality condition in Equation (III-1) to double precision. Many more combinations of cases are available than are presented here. Quadrature sets can be tailored using Table III-2 and Equation 3 to best fit the geometry of the problem. Appendix C shows several additional quadrature sets.

Generating the sets of equation to solve is a straightforward matter, solving them is often a challenge. Some of the equation sets do not have real solution that I have been able to find. The requirement of positive weights also eliminates some possibilities. Appendix B contains some of the work I did in attempting to get valid quadrature sets. Table III-6 summarizes a significant number of possible quadrature set combinations and the results of my efforts. Only those combinations using at least one of the new cases are shown. The "2(3)" notation in the Cases column is intended to show that the resulting equations when selecting a case 2 or a case 3 are mathematically equivalent as discussed above. An entry in the results column of one valid found means a valid quadrature set was found for the resulting system of equations, no valid means all the solutions had either negative weights, imaginary values or cosines greater than one. No solution means that the system of equations did not provide any solutions, good or bad. Empty spaces are quadrature combinations I have not tried and those marked with "tried", means they were attempted, but I was unable to solve the equations. The letters in the order column are a method to distinguish different quadratures of like order, for example  $M_{5b}$ . Valid quadrature sets are presented in Appendix C

Or	der	Cases	Results
3	a	AB1	One valid found
	b	C1	One valid found
	с	AC	No solution
	d	BC	No solution
5	a	AB12(3)	One valid found
	b	AC2(3)	One valid found
-	с	BC2(3)	One valid found
	d	C4	One valid found
	е	ABC1	No solution
	f	C12(3)	No valid
	g	AB4	Tried
	h	A23	
	i	B23	
	j	A13	
	k	B13	
7	a	ABC12(3)	One valid found
	b	AC23	One valid found
	С	ABC4	Tried
	d	AC14	
	е	AB14	Tried
	f	AB42(3)	
	g	AB123	No valid
9	a	ABC123	Tried
	b	ABC2(3) 4	Tried
	С	Etc.	

Table III-3: Case Combination for Quadrature of Order n

Though trial and error can yield valid quadrature sets it may be possible to gain some direction when selecting the cases to use. Since the method is based on the exact integration of spherical harmonics, selecting cases based on features of the spherical harmonics of the order being matched may yield systems of equations that are more readily solved. Figure III-6 (12) shows the first few spherical harmonics.



The three quadratures show in Table III-3 where tested and compared with level symmetric quadratures. These quadratures were chosen because they are the sets with the fewest ordinates for the given order. Also they have ordinates which are reasonable dispersed over the unit sphere and the weights are all about the same order of magnitude. The  $MQ_{5a}$  quadrature was shown in Figure III-5, the  $MQ_{3a}$  and  $MQ_{7a}$  quadrature sets are shown in Figures III-7 and III-8.

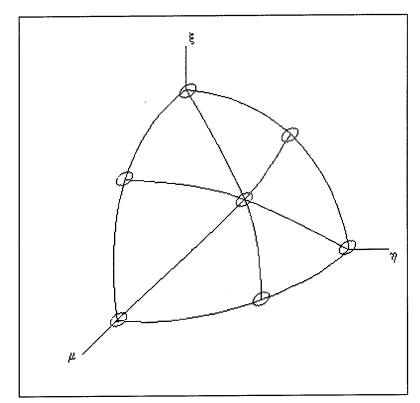


Figure III-7: MQ<sub>3a</sub> Quadrature Layout

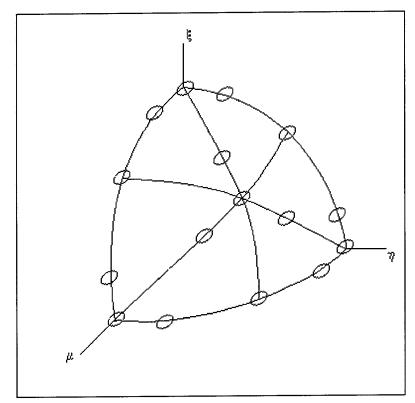


Figure III-8: MQ7a Quadrature Layout

# IV. Results

Some of the new quadrature sets were tested on two problems and compared with the level symmetric LQn quadratures. The first problem is a cube source inside a cube shield region and was designed to exacerbate ray effects. The second problem is a spherical source region in a spherical shield. This problem was selected to reduce ray-effects. Due to the symmetry of this problem, the solution should also by symmetric; therefore any deviation from smooth behavior must be due to the method of calculation. Each test problem was run on the TETRAN code using the linear characteristic (LC) method for handling the spatial computations. The MQ3, MQ5, and MQ7 quadrature sets from Table III-3 are compared with the standard LQ<sub>4</sub>, LQ<sub>10</sub>, and LQ<sub>16</sub> level symmetric quadrature sets. For brevity, no letter subscript will be used with the MQ<sub>n</sub> quadratures. Each problem was also run on the THREEDANT module of the DANTSYS code package using Diamond Difference with negative flux fix-up (DZ). Some quadratures with zero ordinates would cause the code to produce nan (not a number) values for some of the output. The case A ordinates did not cause problems, but the case B and C would. The input file for THREEDANT requires  $\mu$ ,  $\eta$ , and the weights to be entered. The program calculates the value of  $\xi$ . For the case B and C ordinates, when  $\xi$  should be zero, instead a nan would be reported. Error handling routines prevent the code from failing but the output is not useful. By reducing the precision of the quadrature angles input into THREEDANT, the code does

run and provide output. Presumably because the value of  $\xi$ , though very near zero, is sufficiently large so as not to cause problems.

In order to overcome the limitations of THREEDANT with regard to the expectation that the ordinates not lie on the edges of an octant and to the total number of directions in a quadrature set, it was necessary to make further modifications. THREEDANT defines a quadrature only in the principle octant, the remaining angles generated by reflection operations. This also requires the weights that should lie on a boundary to be divided by the number of octants among which it is shared. For example, in order for the normalization of the weights to be correct the weight of a case B or C ordinate must be divided by two. Also, the number of ordinates in a quadrature must match the number of ordinates expected. The variable isn in the input file must be set to the value of n for the LQn quadrature intended for use (there are provisions for other quadrature types not of relevance here). The  $LQ_n$  level symmetric quadratures are selected by default if no quadrature information is given. If a quadrature set is entered, the number of ordinates supplied for the principle octant must match the number in a  $LQ_n$  quadrature for the value of n supplied. This number is given by N = n(n+2)/8. For the MQ<sub>3</sub> quadrature there are seven angles in the principle octant. The closest LQn quadrature with at least that many ordinates is  $LQ_8$  with ten ordinates. Therefore, to run the  $MQ_3$  quadrature, the variable isn must be set to eight and the first seven ordinates in the

**IV-2** 

quadrature array are filled with the  $MQ_3$  values. The remaining three slots in the quadrature array are fill with dummy angles with the weights set to zero. This fix will maintain the quadrature set's ability to integrate properly. The disadvantage of this fix is a loss of computational efficiency because the code still has to perform computations using the dummy values that do not contribute to the solution. Valid comparisons of computational efficiency between the  $MQ_n$  and  $LQ_n$  sets using THREEDANT are therefore limited.

The TETRAN calculations were performed on the IBM SP located at the major shared resource center (MSRC) of Wright Patterson AFB. These problems were run on a single node, consisting of an RS/6000 P2SC model 595 processor with a clock speed of 135MHZ and 1Gb of memory (23). The THREEDANT code was run on an IBM RS-6000/590 workstation using the AIX 3.2.5 operating system with 256 Mb of memory and operating at 67MHz. The convergence tolerance for each problem was 10<sup>-6</sup>.

A benchmark solution was obtained for each problem using a Monte Carlo simulation generated with MCNP (15). MCNP is a general threedimensional simulation code widely used and accepted. MCNP results are the average of the computed quantity over all histories. Each history is a simulation of one particle's motion through the media. MCNP also provides the estimated statistical relative error, R, at the one standard deviation level defined as the sample standard deviation divided by the sample mean. **Problem Definition** 

The cube problem uses the geometry shown in Figure IV-1. The mesh shown is for clarity and not used in calculation. This problem is a threedimensional extension of the two-dimensional problem presented by Lathrop (21) and discussed in references 1 and 3. The source region is a 2x2x2 cm

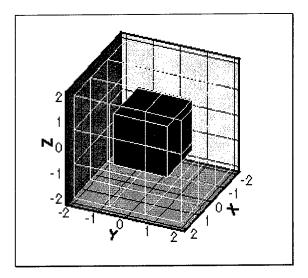


Figure IV-1: Geometry for the Cube in Cube Problem

cube centered in a 4x4x4 cm cube with the source normalized to one over the source volume. The second problem, shown in Figure IV-2, is a spherical source region in a spherical shield region. Problem two used a source region with a radius of 1.5 cm normalized to one over the volume and a shield region with a radius of 3 cm. Figure IV-2 shows the finest tetrahedral mesh used with TETRAN. The source and shield regions have the same nuclear data for all test problems. Vacuum boundaries are used outside all shield regions.

The total cross section,  $\sigma_t$ , is  $\frac{3}{4}$  cm<sup>-1</sup> with a scattering to total cross section

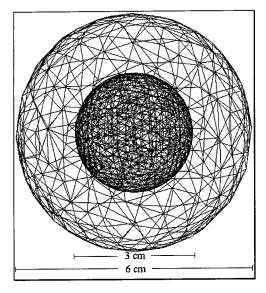


Figure IV-2: Geometry for the Sphere in Sphere Problem

ratio of 
$$\frac{\sigma_s}{\sigma_t} = \frac{2}{3}$$
. This results in a mean free path of  $\frac{4}{3}$  cm. The MCNP

benchmark solution gives the volume average scalar flux in each region, the average scalar flux at the outer surface and the net current through each surface.

### Test Problem One - Cube Source in Cube Shield

The first test problem is a simple 4x4x4 cm cube with vacuum boundaries and a uniformly embedded isotropically-emitting source of strength 1.0 cm<sup>-3</sup> sec<sup>-1</sup>. The source is constrained to the 2x2x2 cm center region of the cube. The nuclear data is given above. The MCNP solution for this problem was run with one million histories. The statistical relative error, R as defined above, is .003 for surface data and .0007 for volume data.

#### **Tetrahedral Mesh**

Each tetrahedral mesh was generated using the Pro/Mesh module of Pro/Engineer (22). This problem was run with two levels of mesh refinement. The meshes are shown in Figure IV-3. Table IV-1 gives the tetrahedral mesh information for test problem one. For clarity, only the surface cells are show for the fine mesh.

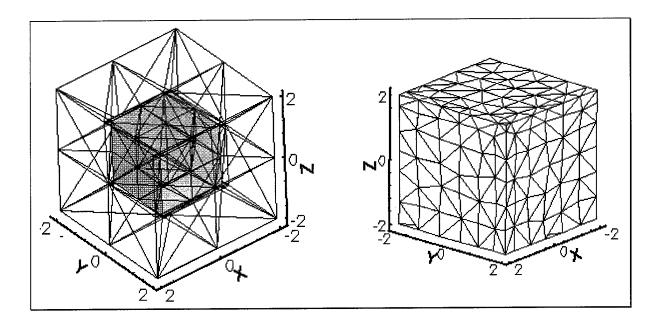


Figure IV-3: Tetrahedral Meshes Used in Test Problem One

Mesh	Total Tetrahedra	Cells in Source Region	Net Volume of Source Region	Average Optical Thickness
Coarse	191	47	8.0	1.203
Fine	1513	162	8.0	0.5997

Table IV-1 : Tetrahedr	al Mesh Data for	Test problem One
------------------------	------------------	------------------

For a right regular problem geometry the tetrahedral mesh does very well in matching the region volumes, even with a relatively coarse mesh. Here the average optical thickness is defined as the mean path length through a cell measured in mean-free paths. The accuracy of each quadrature relative to the benchmark will be examined first.

Figure IV-4 shows contour plots of the surface average scalar flux for the coarse mesh for the MQ<sub>7</sub> and LQ<sub>16</sub> quadrature sets. Figure IV-5 shows the same data for the fine mesh. The plots of the other quadrature sets are very similar and are not shown. The TETRAN code gave the scalar flux at the center of each cell and at the surface of each cell on the boundary. To generate these plots it was necessary to have values at the nodes. Each node value was approximated as the average of the cell center values for each tetrahedron shared by that node. This averaging may effect the variability of the actual surface data. Also from these figures, the flux distribution's dependence on cell geometry and orientation is very evident. These contour plots show little dependence on quadrature. Contour plots of the net current out the cell faces at the surface are very similar to those for the scalar flux and are not presented. These plots indicate that the new quadratures do not cause any significant loss in uniformity and that it is the spatial mesh which is the primary source of surface variation in scalar flux and current. The variability of the data does decreases slightly with quadrature order. For low order quadrature sets the LQ<sub>n</sub> sets have less variability than the MQ<sub>n</sub> sets but this difference is minor. Using the tetrahedral mesh, it is very difficult to get a qualitative comparison of quadrature sets using individual data points. The integral values provide a clearer method of comparison.

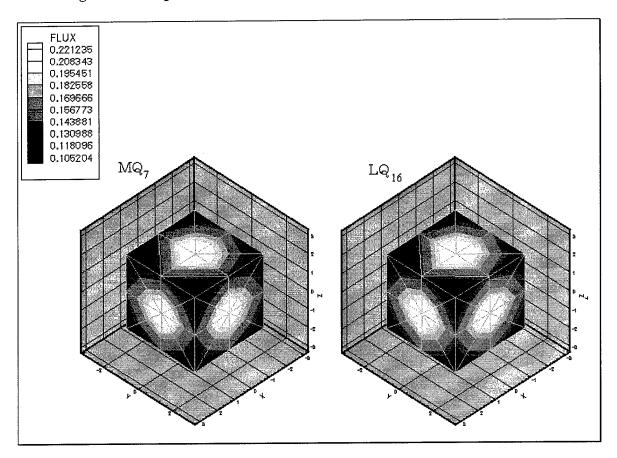


Figure IV-4: Contour Plot of the Surface Average Scalar Flux, Cube Problem,

Coarse Tetrahedral Mesh

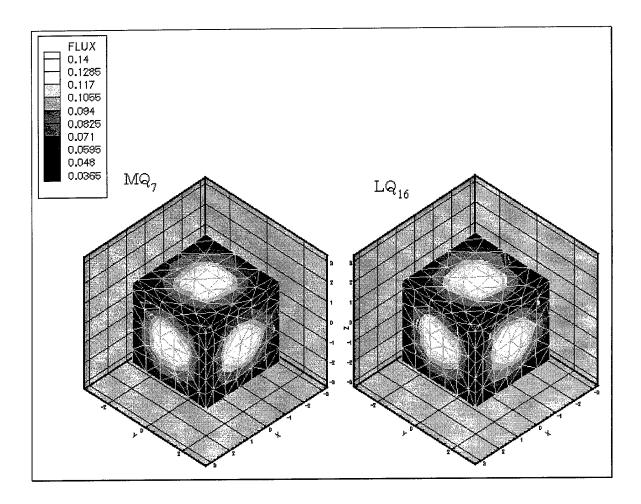


Figure IV-5: Contour Plot of Surface Average Scalar, Cube Problem, Fine Tetrahedral Mesh

The surface average scalar flux for each quadrature and mesh are shown in comparison with the MCNP benchmark calculation in Figures IV-6 and 7. The surface average is calculated as the sum of the product of the flux at the vacuum interface of a surface cell and the area of that cell's face divided by the total surface area,

$$\phi_{\text{surface average}} = \frac{\sum_{\text{cells on surface}} A_{\text{cell}}}{\sum_{\text{cells on surface}} A_{\text{cell}}}$$

The net current out of the cube is similarly shown in Figures IV-8 and 9. The benchmark is shown with error bars indicating the statistical error R. With the exception of the surface flux calculation for the fine tetrahedral mesh, these plot show there is little significant difference in the accuracy achieved by the quadratures tested.

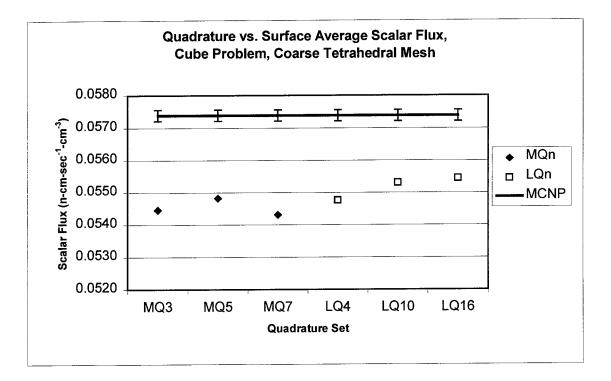


Figure IV-6: Surface Average Scalar Flux, Cube Problem, Coarse Tetrahedral Mesh

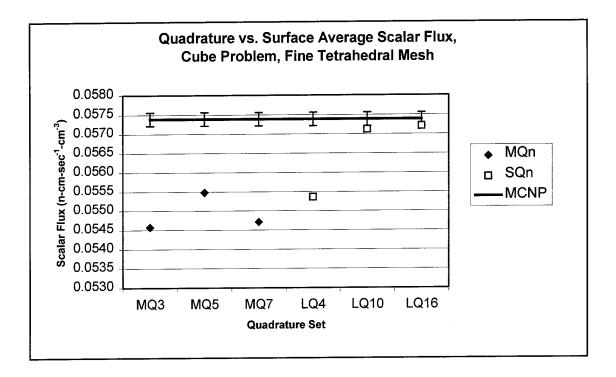


Figure IV-7: Surface Average Scalar Flux, Cube Problem, Fine Tetrahedral Mesh

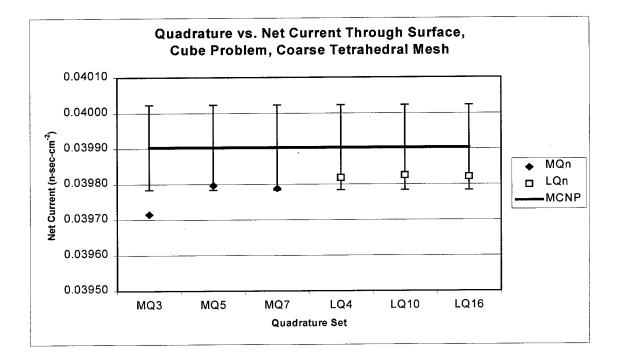


Figure IV-8: Net Current Through the Surface, Cube Problem, Coarse

**Tetrahedral Mesh** 

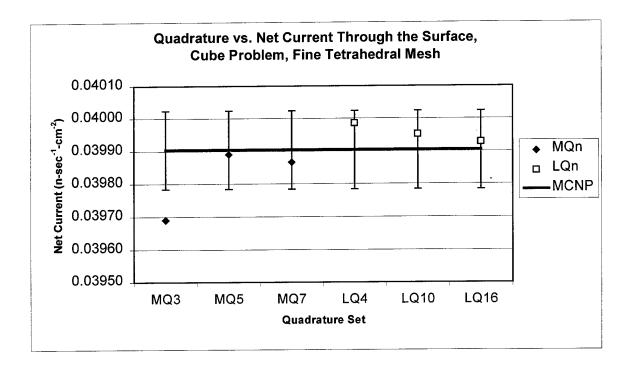


Figure IV-9: Net Current Through the Surface, Cube Problem, Fine Tetrahedral Mesh

The absolute value of the relative error,  $\varepsilon$ , is defined as

$$\varepsilon = \frac{|\phi_{\text{benchmark}} - \phi_{\text{calculated}}|}{\phi_{\text{benchmark}}}$$
(IV-1)

The absolute value of the relative error for the current is found analogously to the flux. The flux, current, and relative error are summarized in Tables VI-2 and 3 for test problem one. The value for  $\varepsilon$  given for the MCNP entry is actually the statistical error R and is shown for reference. Table IV-2: Surface Average Scalar Flux, Net Current, and Relative Error, Cube

Quadrature	Scalar Flux	3	Net Current	3
$MQ_3$	0.054461034	0.05100	0.039715347	0.004732
$\overline{\mathrm{MQ}_{5}}$	0.054827230	0.04462	0.039796675	0.002693
$MQ_7$	0.054314746	0.05355	0.039788637	0.002895
$\overline{LQ_4}$	0.054765258	0.04570	0.039818522	0.002146
$LQ_{10}$	0.055312273	0.03617	0.039825301	0.001976
$LQ_{16}$	0.055447146	0.03382	0.039821360	0.002075
MCNP	0.05739	0.003	0.03990	0.003

Problem, C	Coarse	Tetrahedral	Mesh
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Table IV-3: Surface Average Scalar Flux, Net Current and Relative Error, Cube

Problem, Fine Tetrahedral Mesh

Quadrature	Scalar Flux	3	Net Current	ε
$MQ_3$	0.054574800	0.04902	0.03969063	0.005351
$MQ_5$	0.055480478	0.03324	0.03989043	0.0003442
$MQ_7$	0.054711298	0.04664	0.03986661	0.0009411
$LQ_4$	0.055361991	0.03530	0.03998596	0.002050
$LQ_{10}$	0.057121652	0.004641	0.03995183	0.001194
$LQ_{16}$	0.057210746	0.003090	0.03992792	0.000595
MCNP	0.05739	0.003	0.03990	0.003

For the surface values the  $LQ_n$  quadratures have slightly better accuracy. The difference in relative error between quadrature sets tested is small for the coarse cube. In the fine cube the  $LQ_n$  quadrature sets do better for scalar flux but the  $MQ_5$  and  $MQ_7$  provide comparable results for net current.

Figures IV-10 - 13 show the volume average scalar flux in each region and quadrature with the MCNP solution. The volume average is calculated as the sum of the product of cell flux and cell volume divided by the total region volume,

$$\phi_{\text{average}} = \frac{\sum_{\text{cells in region}} \phi_{\text{cell}} V_{\text{cell}}}{\sum_{\text{cells in region}} V_{\text{cell}}}.$$

The source and shield regions are shown. Tables IV-4 and 5 summarize this data as well as present the relative error.

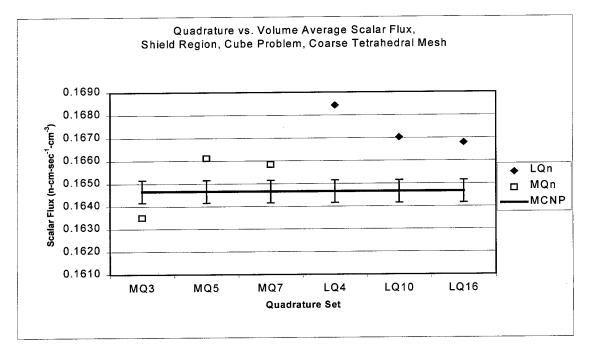


Figure IV-10: Volume Average Scalar Flux, Shield Region, Cube Problem Coarse

**Tetrahedral Mesh** 

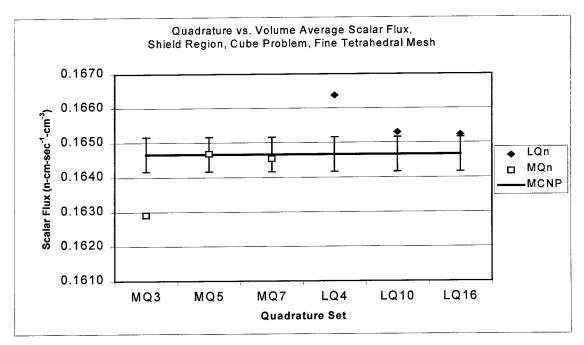


Figure IV-11 Volume Average Scalar Flux, Shield Region, Cube Problem, Fine

**Tetrahedral Mesh** 

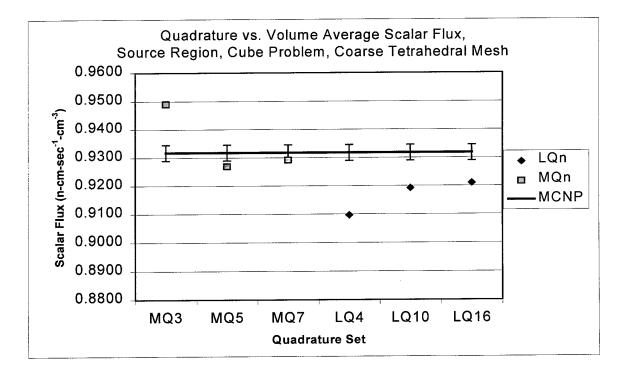


Figure IV-12: Volume Average Scalar Flux, Source Region, Cube Problem,

Coarse Tetrahedral Mesh

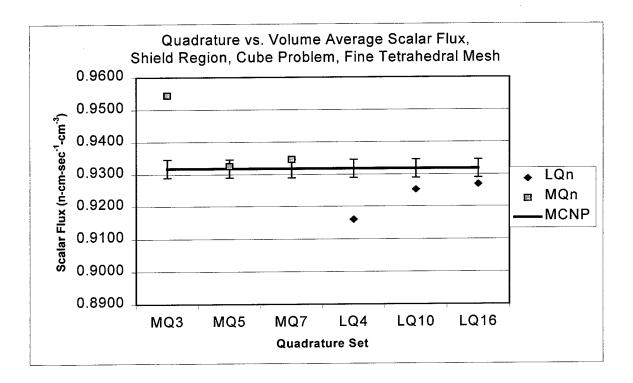


Figure IV-13: Volume Average Scalar Flux, Source Region, Fine Cube

Table IV-4 : Volume Average Scalar Flux and Reletive Error, Cube Problem,

Quadrature	Region	Volume Average	MCNP Benchmark	Relative
		Scalar Flux	+/07%	Error
$MQ_3$	Source	0.9490771	0.93175	0.018596
	Shield	0.1635121	0.16466	0.0069711
$MQ_5$	Source	0.9269118	0.93175	0.0051926
	Shield	0.1661212	0.16466	0.0088739
$MQ_7$	Source	0.9291961	0.93175	0.0027409
· · · · · · · · · · · · · · · · · · ·	Shield	0.1658497	0.16466	0.0072253
$LQ_4$	Source	0.9096752	0.93175	0.023692
	Shield	0.1684340	0.16466	0.022920
$LQ_{10}$	Source	0.9191761	0.93175	0.013495
	Shield	0.1670303	0.16466	0.014395
$LQ_{16}$	Source	0.9209707	0.93175	0.011569
	Shield	0.1668003	0.16466	0.012998

Coarse Tetrahedral Mesh

Quadrature	Region	Volume Average	MCNP Benchmark	Relative
		Scalar Flux	+/07%	Error
$MQ_3$	Source	0.9544397	0.93175	0.024352
	Shield	0.1629156	0.16466	0.010594
$MQ_5$	Source	0.9324943	0.93175	0.00079882
	Shield	0.1646805	0.16466	0.00012450
MQ <sub>7</sub>	Source	0.9346356	0.93175	0.0030970
	Shield	0.1645381	0.16466	0.00074031
$LQ_4$	Source	0.9160339	0.93175	0.016867
	Shield	0.1663769	0.16466	0.010427
$LQ_{10}$	Source	0.9252458	0.93175	0.0069806
	Shield	0.1652950	0.16466	0.0038564
$LQ_{16}$	Source	0.9269132	0.93175	0.0051911
	Shield	0.1652210	0.16466	0.0034068

Table IV-5: Volume Average Scalar Flux, Fine Mesh

The  $MQ_n$  quadrature set performed better than the LQn quadrature set for the volume average data in all but the  $MQ_3$  case on the fine mesh.

Computational efficiency was measured as the user time on the IBM SP taken to solve the problem. Table IV-6 shows the user time for each quadrature and each mesh. Recall the convergence tolerance for all problems was 10<sup>-6</sup>.

Table IV-6 :User Time in Seconds Taken to Solve the Cube in Cube Problem

Quadrature	Coarse Mesh	Iterations to	Fine Mesh	Iterations to
•	(sec)	Convergence	(sec)	Convergence
$MQ_3$	9.51	20	96.75	20
$MQ_5$	18.43	19	187.55	20
$MQ_7$	26.73	19	274.3	20
$\overline{LQ_4}$	9.58	19	92.77	20
$LQ_{10}$	48.15	19	459.17	20
$LQ_{16}$	116.75	19	1100.61	20

For these quadratures to be of value, they must provide increased computational efficiency while maintaining comparable error, or they must provide increased accuracy in comparable time. Figures IV-14 through 17 show plots of user time versus the absolute value of relative error for the surface average scalar flux and net current through the surface. Data points nearer to the origin represent better performance. For these surface averaged values the LQ<sub>n</sub> quadrature sets perform better on this problem except for the current through the surface on the fine mesh, where the MQ<sub>5</sub> and MQ<sub>7</sub> quadratures perform well. Figures IV-18 and 19 show the volume average scalar flux versus relative error. The MQ<sub>n</sub> quadrature sets have consistently better performance, providing less error for the computation cost in all cases except for the MQ<sub>3</sub> on the fine mesh. These plots clearly show the savings in computational costs while maintaining accurate transport results.

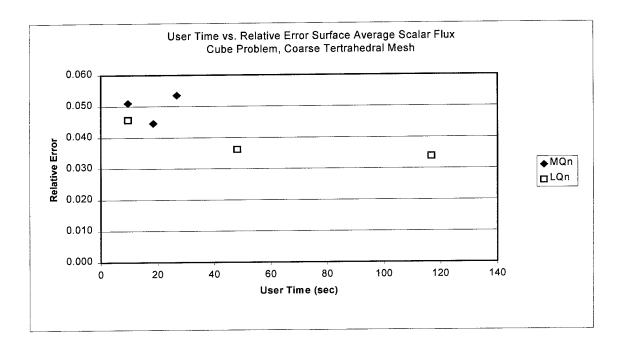
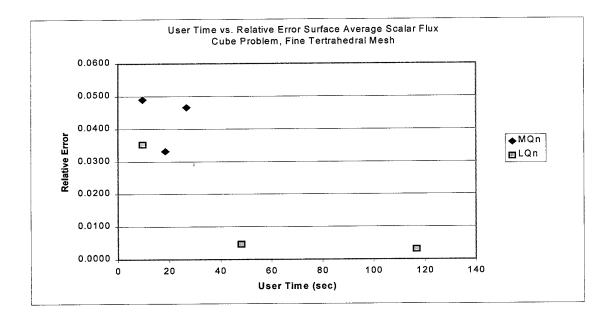
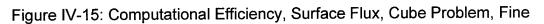


Figure IV-14: Computational Efficiency, Surface Flux, Cube Problem, Coarse

Tetrahedral Mesh





Tetrahedral Mesh

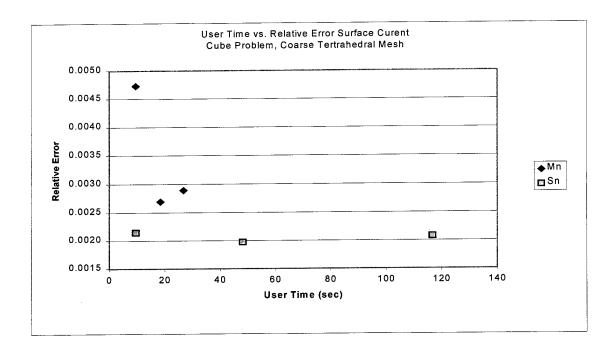


Figure IV-16: Computational Efficiency, Surface Current, Cube Problem Coarse

**Tetrahedral Mesh** 

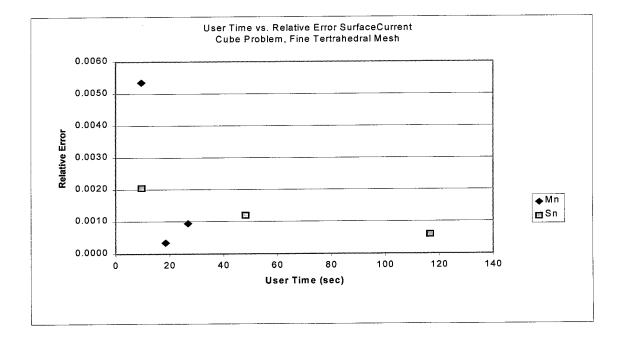


Figure IV-17: Computational Efficiency, Surface Current, Cube Problem Fine

**Tetrahedral Mesh** 

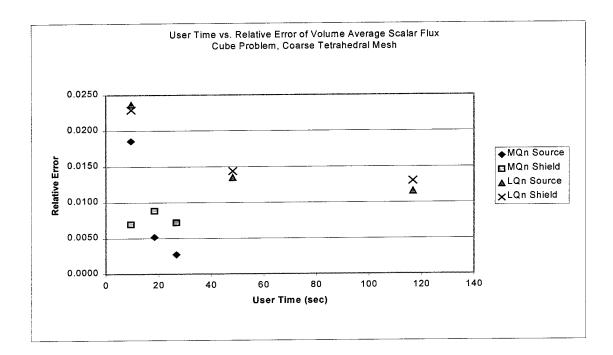


Figure IV-18: Computational Efficiency, Volume Average Scalar Flux, Cube

Problem Coarse Tetrahedral Mesh

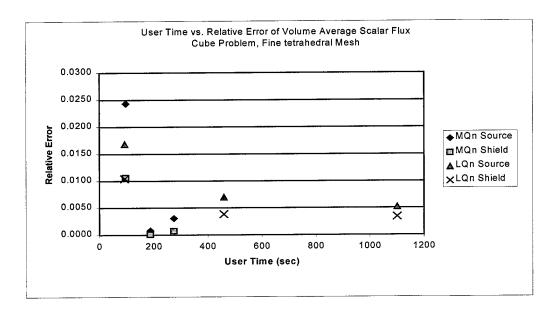


Figure IV-19: Computational Efficiency, Volume Average Scalar Flux, Cube

Problem Fine Tetrahedral Mesh

### Parallelepiped Mesh

This problem was run with one mesh. The mesh is 16x16x16 cells in the principal octant. Reflective boundaries were used on three sides. This results in a 2x2x2 cm cube which is one corner of the test problem containing one eighth of the volume. The solution to the remaining seven eighths of the problem is assumed to be the same by symmetry. The source is normalized to .125 over this source volume which is one eighth of a complete source cube. Data for the parallelepiped mesh is shown in Table VI-7. The volumes of the regions are easily conserved with this meshing method. The number in parenthesis in the volume column is for an equivalent volume if reflective boundaries had not been used.

Table IV-7 : Parallelepiped Mesh Data

Total Cells	Cells in Source	Net Volume of	Optical
	Region	Source Region	Thickness
4096	512	1.0 (8.0)	.09375

Figure IV-20 shows a contour plot of the cell center scalar flux for the base layer of cells (Z= .125 cm plane) using each quadrature set. The flux shows a significant dependency on quadrature for this mesh. The MQ<sub>3</sub> case shows a very pronounced ray effect. The higher order quadrature sets show better behavior.

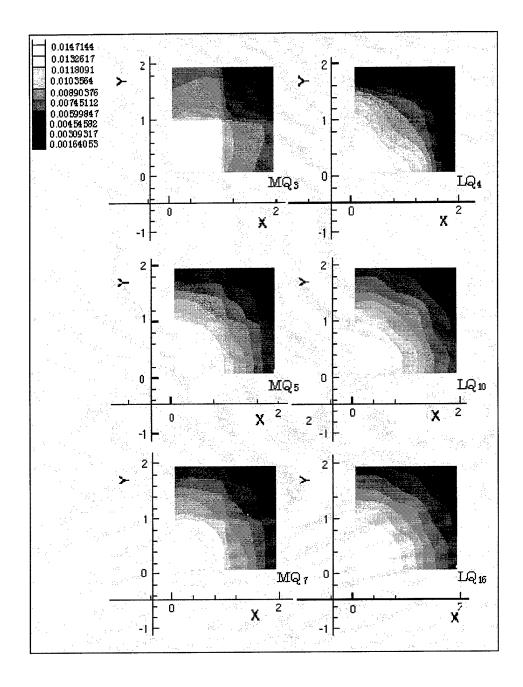


Figure 20: Contour Plot of Scalar Flux in the Z = .125cm Plane, Cube Problem, Parallelepiped Mesh

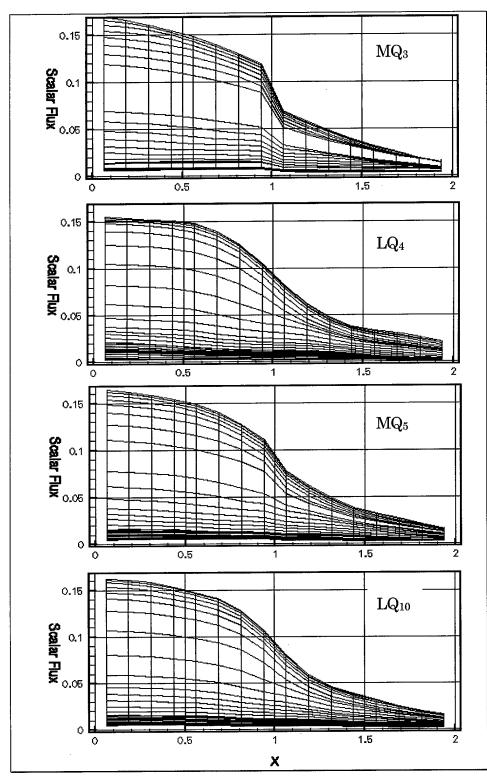


Figure IV-21: Scalar Flux at Cube Surface

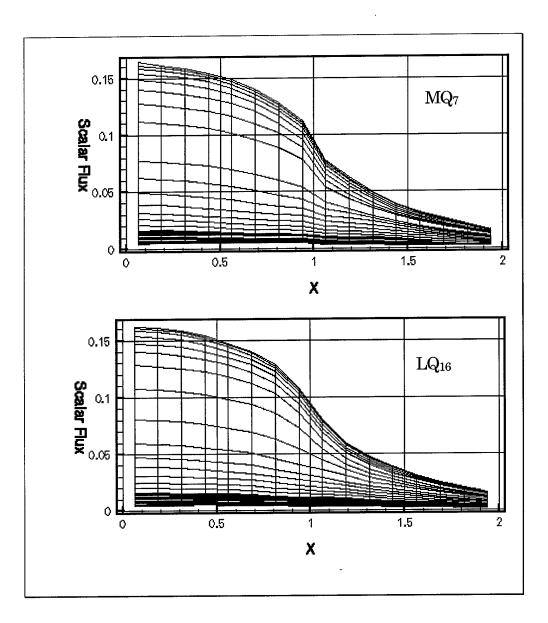


Figure IV-21 (Continued): Scalar Flux at Cube Surface

Figure IV-21 shows a comparison of the flux at the top layer of cells (Z = 1.875 plane). Each successively lower line represent a step of one cell width in the positive y direction starting at y = .125. The lower order MQn quadratures display a substantial drop in flux when crossing from the source to shield region. This is still evident in the MQ7 quadrature but is not as pronounced. This is evidently a ray effect due to the ordinates perpendicular to the axes. Figure IV-22 shows the volume average scalar flux for each quadrature compared with the MCNP benchmark. All quadrature sets performed well, with error less than one percent except for LQ4. MQ5 and LQ<sub>16</sub> performed very well, both with error less than 0.1 percent. The net current through the surface for each quadrature and the MCNP solution are plotted in Figure IV-23. This data is summarized in Table IV-8 with the relative error for each quadrature.

Table IV-8: Parallelepiped Cu	be Data Summary
-------------------------------	-----------------

Quadrature	Flux	3	Net Current	3
MQ3	0.2622893	0.0066899	0.059428	0.489260
MQ5	0.2607804	0.0008985	0.059805	0.498714
MQ7	0.2609637	0.0016024	0.059759	0.497565
LQ4	0.2558341	0.0180855	0.061041	0.529702
LQ10	0.2601124	0.0016650	0.059972	0.502898
LQ16	0.2604749	0.0002739	0.059881	0.500627

Figures IV-24 and 25 shows the relative error vs. computation time. The lines were added for clarity connecting the MQ<sub>n</sub> and LQ<sub>n</sub> quadratures with separate lines by increasing order. When using the THREEDANT code with these quadratures the convergence time is not as direct a function of the number of angles in the quadrature set as it is when using TETRAN. This is primarily due to the added dummy angles required. For Example, both the  $MQ_3$  and  $MQ_5$  quadrature used isn = 8, requiring the input of 10 ordinates. For the  $MQ_3$  case, three of the ordinates were dummy values with zero weights but for the  $MQ_5$  all ten ordinates are used. The  $MQ_7$  quadrature used isn = 12, requiring 21 ordinates, five of which were dummy values. In addition, since THREEDANT uses quadrature sets input by octant, redundant calculations are made for the case A, B, and C ordinates that lie on octant boundaries. Despite this, the  $MQ_n$  quadratures still seem to have better computational efficiency when examining integral values. Table IV-9 shows the computation times for each quadrature sets.

Quadrature	Time (sec)	Iterations to Convergence
MQ3	3.6	10
MQ5	3.03	8
MQ7	11.49	8
LQ4	1.91	9
LQ10	5.56	8
LQ16	24.39	8

Table IV-9 : Time by Quadrature, Parallelepiped Mesh, Cube Probelem

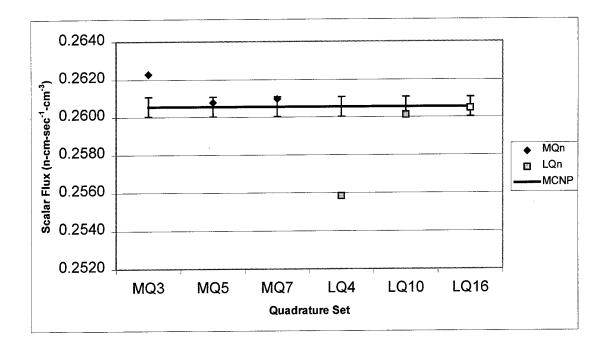


Figure IV-22 : Volume Average Scalar Flux, Cube Problem, Parallelepiped Mesh

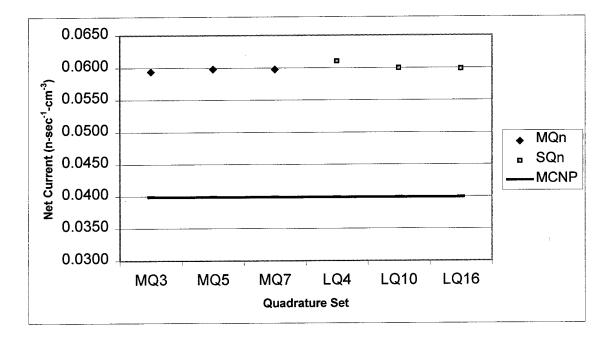


Figure IV-23: Net Current through the Surface, Cube Problem, Parallelepiped

Mesh

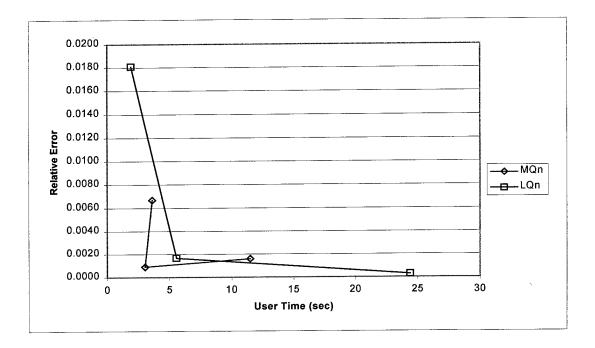


Figure IV-24: User Time vs. Relative Error in Volume Average Scalar Flux, Cube

Problem, Parallelepiped Mesh

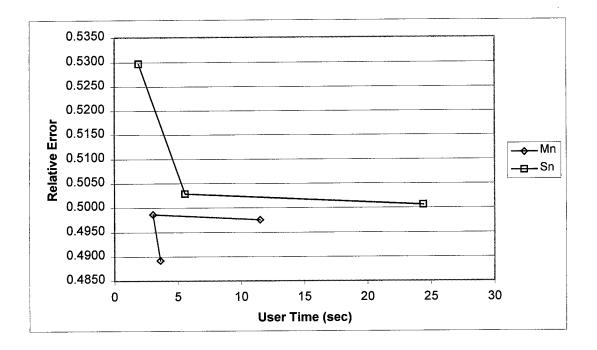


Figure IV-25: User Time vs. Relative Error in Net Current through the Surface,

Cube Problem, Parallelepiped Mesh

# Test Problem Two - Spherical Source in Spherical Shield

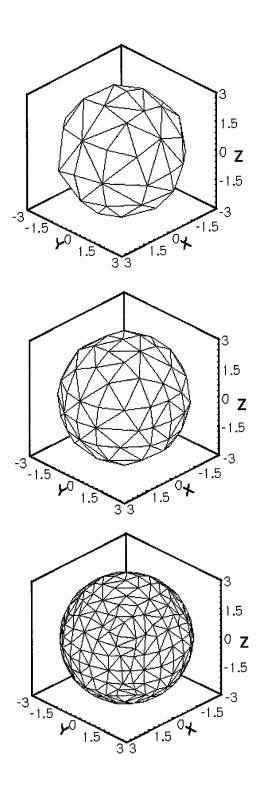
The second test problem is a 3 cm radius sphere with vacuum boundaries and a uniformly embedded isotropically emitting source of strength 1.0 cm<sup>-3</sup> sec<sup>-1</sup>. The source is constrained to a 1.5 cm radius sphere in the center. The nuclear data is the same as the previous problem. The same computer systems and convergence tolerance that were used in problem one were used for problem two.

## **Tetrahedral Mesh**

This problem was run with three levels of mesh refinement. The structure for each mesh is shown in Figure IV-26. Table IV-10 gives the tetrahedral mesh information for test problem two. The volumes shown are the sums of the tetrahedron volumes in each region. This is compared with 14.14 and 98.96 cm<sup>-3</sup> for an actual spherical volume. For curved geometry the Table IV-10 : Tetrahedral Mesh Data

Mesh	Total Tetrahedra	Cells in Source Region	Mesh Volume in Source Region (cm3)	Mesh Volume in Shield Region (cm3)	Average Optical Thickness
Coarse	211	14	5.885	82.24	1.1317
Medium	896	152	11.69	93.48	0.84944
Fine	6632	2033	13.74	96.46	0.40539

tetrahedral mesh does not conserve volume very well until a very fine mesh is used. This is a fault in the design of mesh generators. They are usually used





for finite elements mechanics calculations where volume is not an issue. Figure IV-27 shows the source region for the coarsest mesh. This figure shows how difficult it is to mesh curved regions. The accuracy of each quadrature will again be examined first.

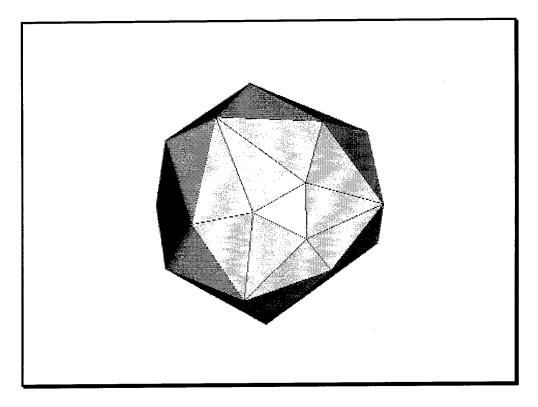


Figure IV-27: Source Region for Sphere Problem, Coarse Mesh

Figure IV-28-30 shows contour plots of the surface average scalar flux at the surface layer of tetrahedra for each mesh and various quadrature sets. Due to the similarity of the plots, not all of the quadratures are presented. The same method of node averaging was used to present this data. These contour plots also show little dependency on quadrature. To better show the quadrature dependence, Figure IV-31 shows contour plots of the fine sphere using MQ<sub>3</sub> and LQ<sub>16</sub> quadrature sets, looking from the *-x*, *x*, and *z* directions respectively. The orientation can be seen from the standard arrowhead  $\odot$ , signifying the positive direction is out of the page, or tail  $\otimes$ , signifying the positive direction is into the page, notation at the origin of each plot. Also from this figure, an unusual asymmetry can be seen. There is a biasing in the positive x direction. This appears to be due to a problem in the TETRAN code that is currently under investigation. The degree of variability in the surface average scalar flux can be seen in Figure IV-32. This figure shows the TETRAN calculated surface flux arbitrarily ordered by magnitude for the medium mesh. Due to the symmetry of the problem the actual flux should be uniform over the surface. The shape of this curve remains the same as the mesh is made more or less refined, however the magnitude of the peak increases to about .11 for the coarse mesh and decreases to about .065 for the fine mesh. These plots are not shown. Figures IV-33 shows the net surface current in the same manner as the scalar flux. This data is ordered by the magnitude of the scalar flux to allow comparison with the previous figure. The current and scalar fluxes have similar trends but do vary independently. Only the  $MQ_7$  and  $LQ_{16}$  data is presented here, the other quadratures have similar behavior.

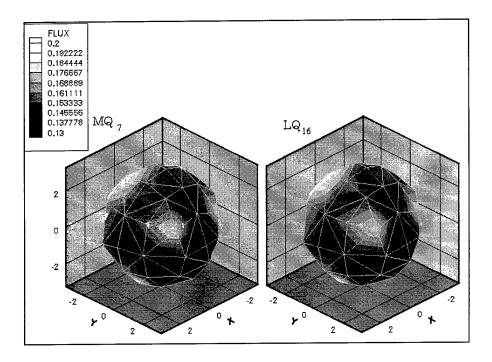


Figure IV-28: Contour Plot of Surface Average Scalar Flux, Sphere Problem,

**Coarse Tetrahedral Mesh** 

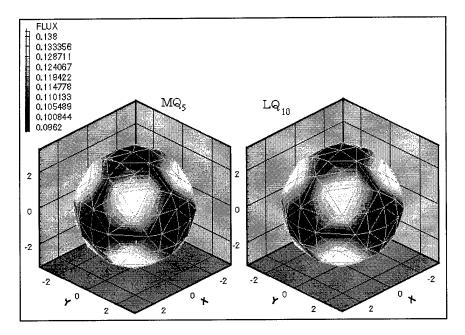


Figure IV-29: Contour Plot of Surface Average Scalar Flux, Sphere Problem,

Medium Tetrahedral Mesh

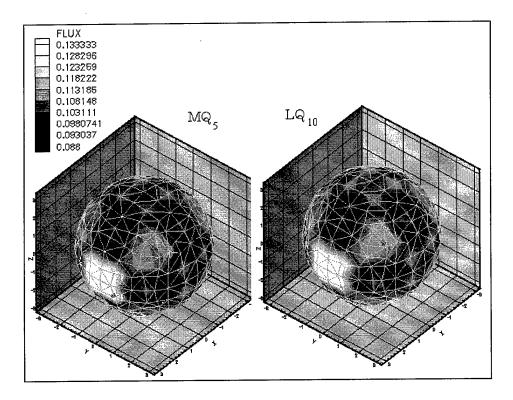


Figure IV-30: Contour Plot of Surface Average Scalar Flux, Sphere Problem,

Fine Tetrahedral Mesh

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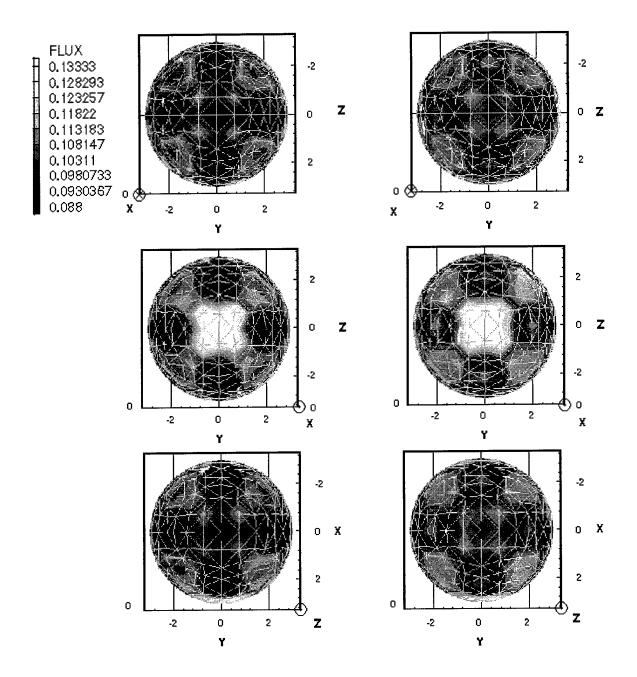


Figure IV-31: Contour Plot of Surface Flux, MQ<sub>5</sub> and LQ<sub>16</sub>, Axial View of Fine Sphere

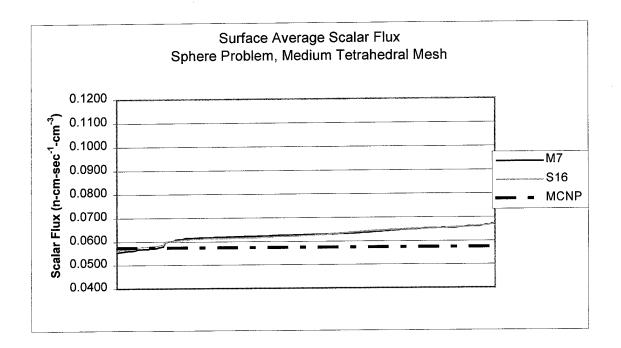


Figure IV-32: Variability of Scalar Flux at the Surface of Sphere Problem,

Medium Tetrahedral Mesh

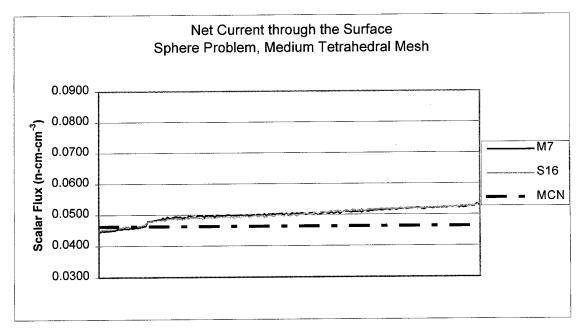


Figure IV-33: Variability in Net Current through the Surface, Sphere Problem,

Medium Tetrahedral Mesh

The surface average scalar flux and net current through the surface show little dependence on quadrature. Figure IV-34 shows this for the scalar flux on the coarse sphere. The small differences in accuracy can be seen in a plot of relative error versus quadrature as shown in Figures IV-35-37 for the scalar flux. Plots of the net current are similar and are not presented. Regardless of mesh, the quadratures have comparable performance in approximating the surface values. The relative error for this geometry is much larger than for the cube problem above. This is primarily attributed to the poor job the mesh does in matching the curved surfaces. Though the MQ<sub>n</sub> sets do appear to have less error for the surface flux on the coarser two meshes, the difference is small and, as can be seen from Figure IV-34, has little significance. This information is summarized in Tables VI-11-13.

Table IV-11: Surface Average Scalar Flux and Net Current, Sphere Problem, Coarse Tetrahedral Mesh

Quadrature	Scalar Flux	3	Net Current	3
$MQ_3$	0.072776076	0.2663	0.05724275	0.2363
$MQ_5$	0.072876460	0.2681	0.05722983	0.2360
$MQ_7$	0.072896331	0.2684	0.05722878	0.2360
$LQ_4$	0.072993465	0.2701	0.05720223	0.2354
$LQ_{10}$	0.073059529	0.2712	0.05723194	0.2360
LQ <sub>16</sub>	0.073047543	0.2710	0.05723134	0.2360
MCNP	0.05747	0.003	0.04630	0.003

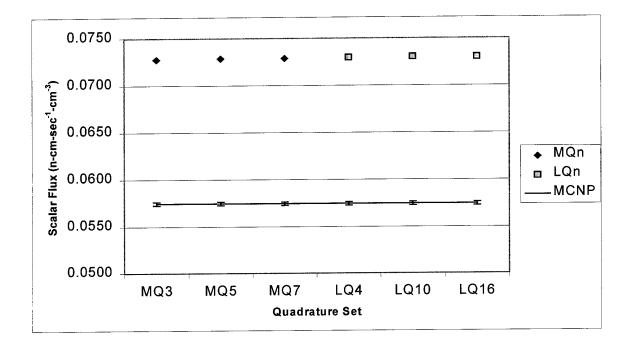


Figure IV-34: Surface Average Scalar Flux, Sphere Problem, Coarse Tetrahedral

Mesh

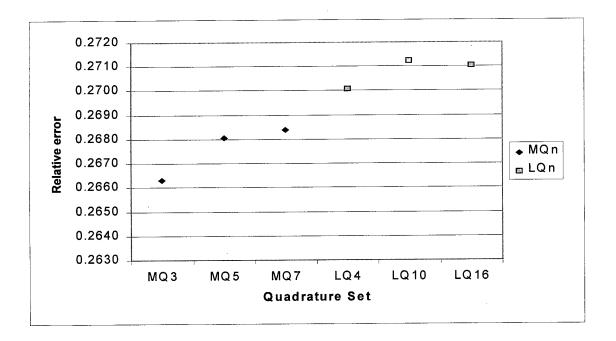


Figure IV-35: Relative Error in Surface Average Scalar Flux, Sphere Problem,

**Coarse Tetrahedral Mesh** 

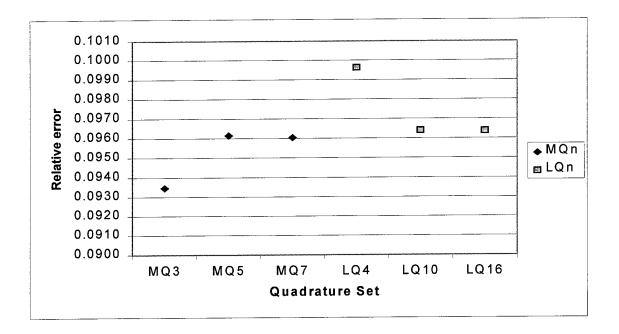


Figure IV-36: Relative Error in Surface Average Scalar Flux, Sphere Problem,

Medium Tetrahedral Mesh

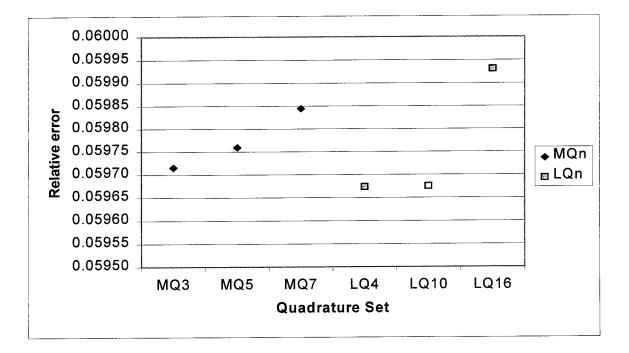


Figure IV-37: Relative Error in Surface Average Scalar Flux, Sphere Problem,

Fine Tetrahedral Mesh

Table IV-12: Surface Average Scalar Flux and Net Current, Sphere Problem,

Quadrature	Scalar Flux	3	Net Current	3
$MQ_3$	0.062842334	0.0935	0.05024186	0.0851
$\overline{MQ_5}$	0.062996823	0.0961	0.05027019	0.0857
MQ <sub>7</sub>	0.062990863	0.0960	0.05026612	0.0856
$LQ_4$	0.063196562	0.0996	0.05028158	0.0859
$LQ_{10}$	0.063011997	0.0964	0.05028019	0.0859
$LQ_{16}$	0.063010129	0.0964	0.05027461	0.0858
MCNP	0.05747	0.003	0.04630	0.003

# Medium Tetrahedral Mesh

Table IV-13: Surface Average Scalar Flux and Net Current, Sphere Problem,

Fine Tetrahedral Mesh

Quadrature	Scalar Flux	3	Net Current	3
$\overline{MQ_3}$	0.060902901	0.0597	0.04857312	0.0490
$\overline{MQ_5}$	0.060905444	0.0598	0.04857166	0.0490
$\overline{MQ_7}$	0.060910336	0.0598	0.04857333	0.0490
$LQ_4$	0.060900515	0.0597	0.04857391	0.0490
LQ <sub>10</sub>	0.060900610	0.0597	0.04856642	0.0489
LQ <sub>16</sub>	0.060915316	0.0599	0.04856830	0.0489
MCNP	0.05747	0.003	0.04630	0.003

Tables IV-14 and 15 summarize the volume average data. All quadratures show nearly equal performance in calculating the volume average scalar flux

Quadrature	Region	Volume	MCNP	3
	Ŭ	Average Scalar	Benchmark	
		Flux	+/2%	
$MQ_3$	Shield	0.259490998	0.17601	0.47430
	Source	2.073454733	1.29600	0.59989
MQ5	Shield	0.260540247	0.17601	0.48026
	Source	2.059669935	1.29600	0.58925
MQ7	Shield	0.260516136	0.17601	0.48012
	Source	2.060075531	1.29600	0.58956
LQ4	Shield	0.261797736	0.17601	0.48740
	Source	2.043972732	1.29600	0.57714
LQ10	Shield	0.260675794	0.17601	0.48103
	Source	2.057621449	1.29600	0.58767
LQ16	Shield	0.26067187	0.17601	0.48101
	Source	2.05772263	1.29600	0.58775

Table IV-14: Volume Average Scalar Flux Data, Coarse Sphere

for a given mesh. By examining the errors in the source region shown in Tables IV-15 and 16 an interesting phenomenon can be detected when refining the mesh from the medium to the fine sphere. The error in the source region increases from about one percent up to about nine percent. The expected trend is for error to decrease with this level of mesh refinement. The source of this increase in error is suspected to be the same unresolved problems in the TETRAN code that induced the biasing mentioned earlier. This only seems to become detectable when running fine mesh problems. It is suspected that a very small biasing in the positive x direction that only accumulates significantly for very fine meshes is caused by a tetrahedron splitting algorithm used by TETRAN. The magnitude of this problem does not appear to vary with quadrature and therefore we can still use this output

for relative comparisons between quadrature sets.

Quadrature	Region	Volume	MCNP	3
		Average Scalar	Benchmark	
		Flux	+/2%	
$MQ_3$	Shield	0.206680395	0.17601	0.17425
	Source	1.313078744	1.29600	0.01318
$MQ_5$	Shield	0.20660516	0.17601	0.17383
	Source	1.312631187	1.29600	0.01283
$MQ_7$	Shield	0.206585308	0.17601	0.17371
	Source	1.312932737	1.29600	0.01307
$LQ_4$	Shield	0.206521942	0.17601	0.17335
	Source	1.312863629	1.29600	0.01301
$LQ_{10}$	Shield	0.20670228	0.17601	0.17438
	Source	1.311474207	1.29600	0.01194
$LQ_{16}$	Shield	0.206639441	0.17601	0.17402
•	Source	1.312184651	1.29600	0.01249

Table IV-15: Volume Average Scalar Flux Data, Medium Sphere

Table IV-16: Volume Average Scalar Flux Data, Fine Sphere

Quadrature	Region	Volume	MCNP	Absolute
v		Average Scalar	Benchmark	Relative Error
		Flux	+/2%	
$MQ_3$	Shield	0.19327010	0.17601	0.09806
······································	Source	1.18027900	1.29600	0.08929
$MQ_5$	Shield	0.19331830	0.17601	0.09834
	Source	1.17998870	1.29600	0.08951
$MQ_7$	Shield	0.19331458	0.17601	0.09832
	Source	1.17996012	1.29600	0.08954
$LQ_4$	Shield	0.19331059	0.17601	0.09829
	Source	1.17996978	1.29600	0.08953
$LQ_{10}$	Shield	0.19336766	0.17601	0.09862
	Source	1.17981243	1.29600	0.08965
$LQ_{16}$	Shield	0.19334904	0.17601	0.09851
· ·	Source	1.17988161	1.29600	0.08960

Computational efficiency was measured in the same fashion as for problem one. Table IV-17 shows the user time for each quadrature and each mesh.

Quadrature	Coarse Mesh	Medium Mesh	Fine Mesh
MQ <sub>3</sub>	13.46	62.73	531.79
$\overline{MQ_5}$	25.78	121.17	1015.44
$\overline{MQ_7}$	37.99	177.87	1494.66
$LQ_4$	12.48	58.87	491.42
$LQ_{10}$	62.47	278.98	2471.91
$LQ_{16}$	149.05	701.37	5788.87

Table IV-17 : User Time in Seconds Taken to Solve the Sphere Problem

Figures IV-38 shows a plot of data from Table IV-14 of user time versus relative error in the volume average scalar flux for the coarse mesh. This curve is nearly flat. Curves for the other levels of mesh refinement and for the net current as similar. The lack of significant features implies the quadratures have already converged to the minimum error obtainable by the discrete ordinates method even for the lowest order quadrature. This is not surprising considering the simple nature of the problem.

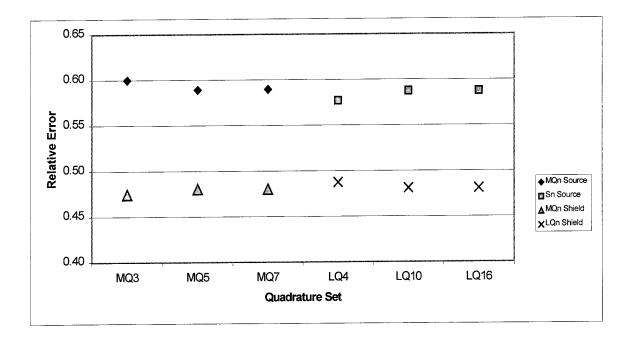


Figure IV-38: Relative Error in Volume Average Scalar Flux, Sphere Problem, Coarse Tetrahedral Mesh

### Parallelepiped Mesh

Only one mesh was used for this problem. As in the cube case, reflective boundaries were used on three sides and the problem was run using a one eighth section of the sphere. The remainder of the problem is assumed to be the same by symmetry. Data for the parallelepiped mesh is shown in Table VI-18. The numbers in parentheses in the volume columns are for an entire sphere if reflective boundaries had not been used. The volumes shown here are compared with 98.96 and 14.14 cm<sup>3</sup> for an actual spherical volume. This method of mesh generation has a difficult time matching curved surfaces. Table IV-18 : Parallelepiped Mesh Data

	Cells in Material	Cells in Source Region		Net Volume of Source Region (cm <sup>3</sup> )
3375	1464	211	10.02 (80.2)	1.688 (13.5)

The cells are cubes .1 cm across corresponding to .075 optical thickness. The source volume was allowed to bulge into the shield region a small amount to allow for closer volume modeling. Figure IV-39 shows contour plots of the scalar flux in the x-y plane cutting through the origin. Despite the rough geometry of the mesh, the results are fairly uniform. Figure IV-40 shows the volume average scalar flux and Figure 41 shows the relative error. All quadrature sets did well, with error less than two percent. This data is summarized in Table IV-19

Table IV-19: Sphere Data Summary, Parallelepiped Mesh

Quadrature	Volume Average	Relative Error	Net Current	Relative Error
-	Scalar Flux			
$MQ_3$	0.3110153	0.0158015	0.047224	0.019887
$\overline{MQ_5}$	0.3111058	0.0155153	0.047246	0.020359
$MQ_7$	0.3111027	0.0155251	0.047224	0.019871
$LQ_4$	0.3107919	0.0165086	0.047302	0.021565
$LQ_{10}$	0.3111415	0.0154021	0.047215	0.019677
$LQ_{16}$	0.3112716	0.0149906	0.047218	0.019755

Figure IV-42 shows the net current by quadrature with the MCNP benchmark. Again, the results show little deviation by quadrature. Figures IV-43 and 44 show the relative error vs. computation time. Neither quadrature method seems to have a clear advantage.

Quadrature	Time (sec)
MQ3	3.24
MQ5	3.3
MQ7	9.58
LQ4	2.92
LQ10	5.81
LQ16	22.97

Table IV-20 : Time by Quadrature, Sphere Problem, Parallelepiped Mesh

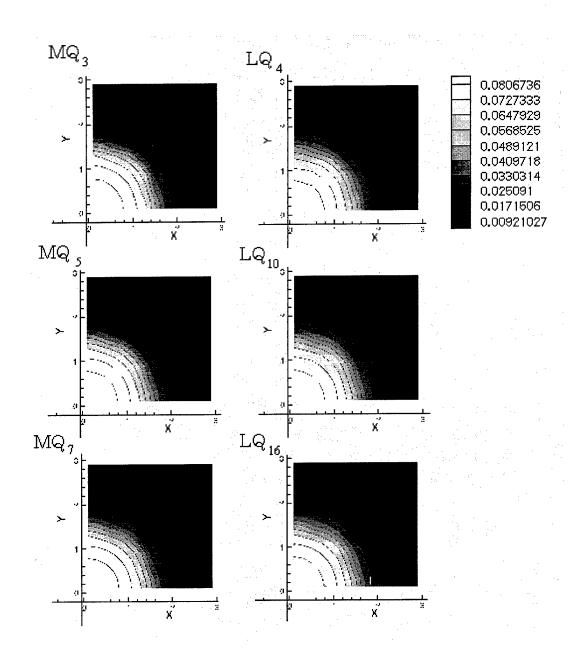


Figure IV-39: Comparison of Contour Plots: Scalar Flux, Z = .125 plane, Sphere Problem, Parallelepiped mesh

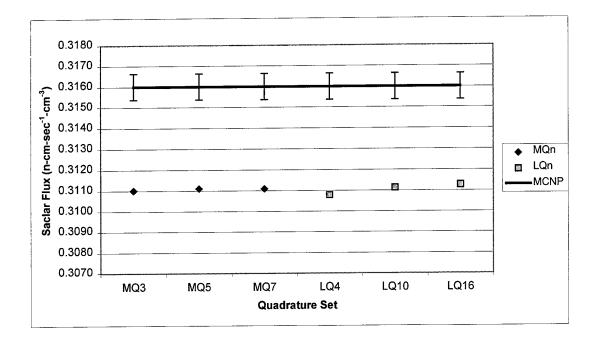


Figure IV-40 : Volume Average Scalar Flux, Sphere Problem, Parallelepiped

Mesh

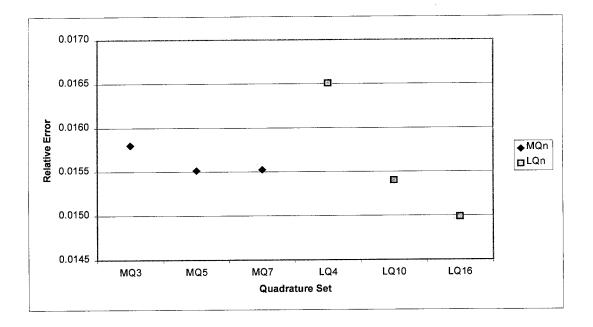


Figure IV-41: Relative Error, Volume Average Scalar Flux, Sphere Problem,

Parallelepiped Mesh

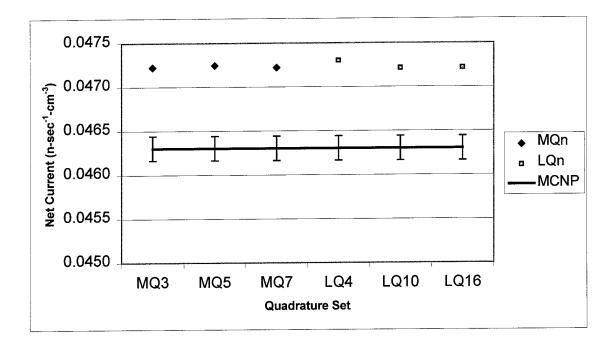


Figure IV-42: Net Current, Sphere Problem, Parallelepiped Mesh

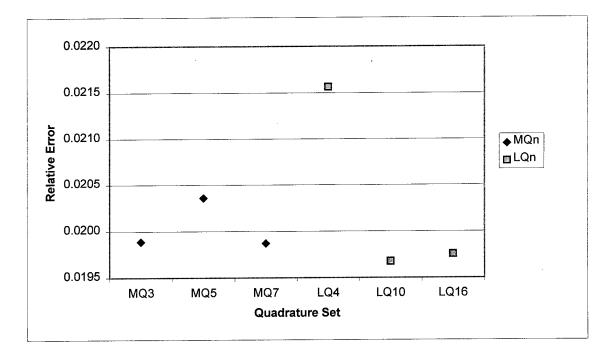


Figure IV-43: Relative Error in Net Current, Sphere Problem, Parallelepiped

Mesh

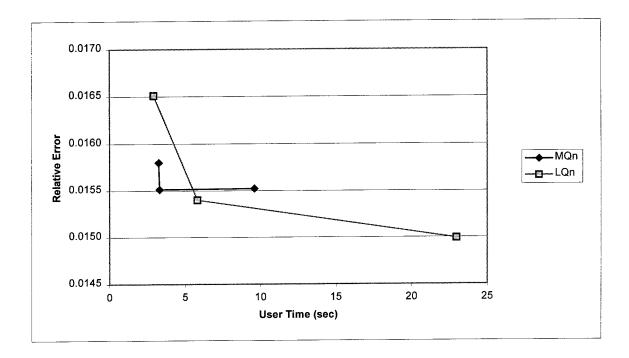
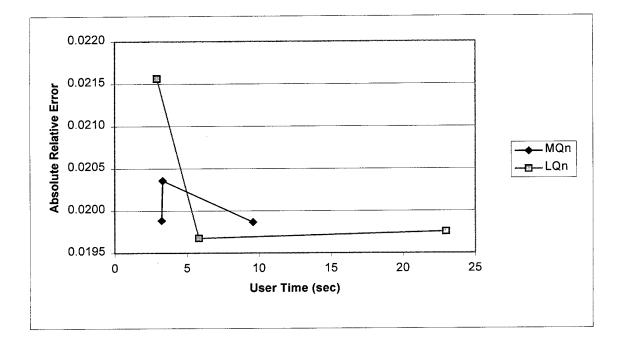
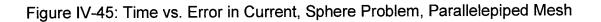


Figure IV-44: Time vs. Error in Flux, Sphere Problem, Parallelepiped Mesh





### V. Conclusion

The quadratures developed here show the potential for use in discrete ordinates transport calculations. The MQ<sub>n</sub> quadratures tested do provide accuracy comparable to traditional methods, and often at a substantial saving in time. However, there is as yet insufficient data to make conclusive statements on the effectiveness of the method.

Decreased computational cost for use on parallelepiped meshes may be obtained by development of new computer codes to take advantage of the one, or two dimensional aspect of the special directions, and by use of more flexible quadrature input modules to allow any number of ordinates.

The MQ<sub>n</sub> quadratures seem to perform best in the unstructured mesh. It is difficult to determine the type problem with regard to mesh type and problem geometry that these quadratures will work best with due to the limited amount of data obtained. The sphere in sphere problem did not adequately differentiate performance between quadrature sets. More work is need on a greater variety of problems before significant conclusions on this aspect of performance can be made.

The reduced number of angles in the new quadratures does reduce the computational cost of the problems solved. Using the tetrahedral mesh, these quadratures had comparable accuracy to the  $LQ_n$  quadratures with regard to smoothness and global accuracy. Their performance was best when used to find integral results. On the parallelepiped mesh, the  $MQ_n$  quadratures

V-1

showed substantial ray effects though still did well when calculating integral results.

I recommend using the  $MQ_n$  quadratures on problems with little symmetry on an unstructured mesh. For structured meshes and problems with high symmetry, more data is need before a recommendation can be made.

### Recommendations for further research

More quadrature sets need to be solved and more data needs to be generated. From an increased database, the optimal quadrature set to use as a function of problem geometry and mesh type may be found. Also, the potential benefits of the higher order of these quadratures as compared to others with the same number of angles needs to be investigated by evaluating them on anisotropic test problems.

Development of code to take advantage of the one- and twodimensional aspects of the MQ<sub>n</sub> quadrature sets may provide an interesting challenge with potentially a great deal of gain. Integrating the new code modules with current programs may allow for easy transition to these new sets where applicable

I have been unable to solve any quadrature above seventh order. I spent many hours on Mathematica, and attempted to write a FORTRAN program using an expansion method, but was unsuccessful. More time in

V-2

developing a method of solving the high order polynomials may yield higher order quadratures.

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### **Appendix A**

Each case A, B, C, 1, 2, 3, and 4 included in a quadrature contribute to the summation

$$\sum_{n=1}^{N} w_n \mu_n^{2k}, \qquad k = 1, 2, \ldots, L.$$

Because of the symmety requirements all of the weights for a particular case are required to be the identical and all of the angles for each case are functions of the angles,  $\mu$  and  $\eta$ , in the base hextant. In some of the cases the angles have restrictions on them so they are not free parameters. The number of degrees of freedom provided by a case is the number of free angles plus one for the weight. Each case below will have a discussion of which parameters are free.

CaseA: This consist of six points over the unit sphere, two on each coordinate axis at ± 1. All of the angles are restricted in this case and only the weight is free, yielding one degree of freedom. The contribution of caseA to the summation is

$$\mathbf{w}[(1)^{2k}+(0)^{2k}+(0)^{2k}+(0)^{2k}+(0)^{2k}+(-1)^{2k}]$$

which simplifies to

2 w

for all  $k \neq 0$ .

CaseB: This case consist of 12 ordiates over the unit sphere. Each ordinate has one direction cosine equal to zero and the other two equal  $\frac{1}{\sqrt{2}}$  in the principal

octant. The only free parameter for this case then is the weight for one degree of freedom. The contribution of caseB to the summations is

$$w \left[ 4 \left( \frac{1}{\sqrt{2}} \right)^{2k} + 4 (0)^{2k} + 4 \left( -\frac{1}{\sqrt{2}} \right)^{2k} \right]$$

which simplifies, for  $k \neq 0$ , to

8 w 2<sup>k</sup>.

CaseC: This case consist of 24 ordinates

over the unit sphere. Each ordinate has one direction cosine equal to zero,

one free, and the other is found from  $\mu^2 + \eta^2 + \xi^2 = 1$ .

This leaves one free angle and the weight for two degrees of freedom. The contribution to the summation is

$$w \left[ 4 \mu^{2k} + 4 \left( \sqrt{1 - \mu^2} \right)^{2k} + 8 (0)^{2k} + 4 (-\mu)^{2k} + 4 \left( -\sqrt{1 - \mu^2} \right)^{2k} \right] .$$

This simplifies to, for  $k \neq 0$ , to

$$8 w \left[ \mu^{2k} + (1 - \mu^2)^k \right]$$
  
A-1

Casel: This case consist of eight ordinates over the unit sphere. Each direction cosine is equal to  $\frac{1}{\sqrt{3}}$  in the principal octant. This leaves only the weight free

for one degree of freedom. The contribution is

$$w\left[4\left(\frac{1}{\sqrt{3}}\right)^{2k}+4\left(-\frac{1}{\sqrt{3}}\right)^{2k}\right].$$

This simplifies to

Case2: This consist of 24 ordinates over the unit sphere. One angle is free, the other two are equal. This yields two degrees of freedom, one angle and one weight. The contribution is

$$w \left[ 4 \mu^{2k} + 8 \left( \sqrt{\frac{1-\mu^2}{2}} \right)^{2k} + 4 (-\mu)^{2k} + 8 \left( -\sqrt{\frac{1-\mu^2}{2}} \right)^{2k} \right]$$

which simplifies to

$$8 \operatorname{w} \left[ \mu^{2k} + 2 \left( \frac{1 - \mu^2}{2} \right)^k \right] \, .$$

Case3: This consist of 24 ordinates over the unit sphere. One angles is free and the other two are equal. This yields two degrees of freedom as for case2 above. There are two way to look at the contribution for case3. The first is to note that all angles for this case result if when solving for a case2, the angles are less

than  $\frac{1}{\sqrt{3}}$ . This puts the angle out of the base hextant if this angle is  $\mu$ . But if,

when this results, we concider the angle as  $\eta$  instead,

this ordinate will lie in the base hextant and is infact a case3 ordinate. The other method is to proceed as normal resuling in the following contribution equation

$$\mathbf{w} \left[ 8\,\mu^{2\,k} + 4\,\left(\sqrt{1-2\,\mu^2}\,\right)^{2\,k} + 8\,\left(-\mu\right)^{2\,k} + 4\,\left(-\sqrt{1-2\,\mu^2}\,\right)^{2\,k} \right]$$

which simplifies to

$$8 w [2 \mu^{2k} + (1 - 2 \mu^{2})^{k}]$$
.

I prefer to use the former method and use a contribution equation in the form

$$8 \operatorname{w} \left[ \eta^{2k} + 2 \left( \frac{1 - \eta^2}{2} \right)^k \right] \,.$$

The two equations are equivalent,

as can be easily verified by substituting  $\mu = \sqrt{\frac{1-\eta^2}{2}}$  into the first result.

Case4: This consist of 48 directions over the unit sphere. Both  $\mu$  and  $\eta$  are free parameters as is the weight. This is the most expensive case to inclued and the most difficult to solve. To find the three angles defining this ordinate in the base hextant, we must leave  $\mu$  and  $\eta$  as free parameters and find  $\xi$  using  $\xi^2 =$ 

1 -  $\mu^2 + \eta^2$ . The resulting contribution equation is

$$\mathbf{w} \left[ 8\mu^{2k} + 8\eta^{2k} + 8\left(\sqrt{1-\mu^2-\eta^2}\right)^{2k} + 8(-\mu)^{2k} + 8(-\eta)^{2k} + 8\left(-\sqrt{1-\mu^2-\eta^2}\right)^{2k} \right]$$

which simplifies to

$$16 w \left[ \mu^{2k} + \eta^{2k} + \left( 1 - \mu^2 - \eta^2 \right)^k \right] \,.$$

Off[General::spell1]

caseA[w\_, k\_] := 6 w /; k == 0 caseA[w\_, k\_] := 2 w /; k != 0 caseB[w\_, k\_] := 12 w /; k == 0 caseB[w\_, k\_] := 8  $\frac{w}{2^{\times}}$  /; k != 0 caseC[w\_,  $\mu_{-}$ , k\_] := 24 w /; k == 0 caseC[w\_,  $\mu_{-}$ , k\_] := 8 w ( $\mu^{2^{\times}}$  + (1 -  $\mu^{2}$ )<sup>k</sup>) /; k != 0 case1[w\_, k\_] := 8  $\frac{w}{3^{\times}}$ case2[w\_,  $\mu_{-}$ , k\_] := 8 w ( $\mu^{2^{\times}}$  + 2  $\left(\frac{1 - \mu^{2}}{2}\right)^{k}$ ) (\* When a case 3 ordinate is desired, use the case 2 equation an substitute  $\eta$  for  $\mu$ . \*) case4[w\_,  $\mu_{-}$ ,  $\eta_{-}$ , k\_] := 16 w ( $\mu^{2^{\times}}$  + (1 -  $\mu^{2} - \eta^{2}$ )<sup>k</sup> +  $\eta^{2^{\times}}$ ) (\* The Filter function below searches a list of quadrature output from a Solve function and \*) (\* returns null values for elements in the list with negative weights or angles or imaginary values. \*)

Filter[TL\_, vars\_] := Table[If[(vars /. TL)[[j]] ==
 Table[Select[Re[(vars /. TL)[[i]]], # > 0&], {i, Length[TL]}][[j]], TL[[j]]],
 {j, Length[TL]}]

GetQuad[Equations\_, vars\_] := Filter[NSolve[Equations, vars], vars]

Eqns = Table [Expand[caseA[wA, k] + caseC[wC,  $\mu$ C, k]] ==  $\frac{1}{2 k + 1}$ , {k, 3}] {2 wA + 8 wC ==  $\frac{1}{3}$ , 2 wA + 8 wC - 16 wC  $\mu$ C<sup>2</sup> + 16 wC  $\mu$ C<sup>4</sup> ==  $\frac{1}{5}$ , 2 wA + 8 wC - 24 wC  $\mu$ C<sup>2</sup> + 24 wC  $\mu$ C<sup>4</sup> ==  $\frac{1}{7}$ } Solve [Eqns, {wA, wC,  $\mu$ C}]

SOIVE[EQHS, WA, WC, H

{}

(\* The caseA + caseC combination has no solution \*)

Eqns = Table [Expand[caseB[wB, k] + caseC[wC,  $\mu$ C, k]] ==  $\frac{1}{2 k + 1}$ , {k, 3}] {4 wB + 8 wC ==  $\frac{1}{3}$ , 2 wB + 8 wC - 16 wC  $\mu$ C<sup>2</sup> + 16 wC  $\mu$ C<sup>4</sup> ==  $\frac{1}{5}$ , wB + 8 wC - 24 wC  $\mu$ C<sup>2</sup> + 24 wC  $\mu$ C<sup>4</sup> ==  $\frac{1}{7}$ } Solve[Eqns, {wB, wC,  $\mu$ C}]

{ }

(\* The BC case combination has no solution \*)

Eqns = Table [case1[w1, k] + caseC[wC,  $\mu$ C, k] ==  $\frac{1}{2k+1}$ , {k, 3}]  $\left\{\frac{8 \text{ w1}}{3} + 8 \text{ wC} == \frac{1}{3}, \frac{8 \text{ w1}}{9} + 8 \text{ wC} \left(\mu C^4 + (1 - \mu C^2)^2\right) == \frac{1}{5}, \frac{8 \text{ w1}}{27} + 8 \text{ wC} \left(\mu C^6 + (1 - \mu C^2)^3\right) == \frac{1}{7}\right\}$ 

Solve[Eqns, {w1, wC,  $\mu$ C}]

$$\left\{ \left\{ w1 \rightarrow \frac{9}{280}, wC \rightarrow \frac{13}{420}, \mu C \rightarrow -\sqrt{\frac{1}{26} \left( 13 - \sqrt{65} \right)} \right\}, \left\{ w1 \rightarrow \frac{9}{280}, wC \rightarrow \frac{13}{420}, \mu C \rightarrow \sqrt{\frac{1}{26} \left( 13 - \sqrt{65} \right)} \right\}, \left\{ w1 \rightarrow \frac{9}{280}, wC \rightarrow \frac{13}{420}, \mu C \rightarrow -\sqrt{\frac{1}{26} \left( 13 + \sqrt{65} \right)} \right\}, \left\{ w1 \rightarrow \frac{9}{280}, wC \rightarrow \frac{13}{420}, \mu C \rightarrow \sqrt{\frac{1}{26} \left( 13 + \sqrt{65} \right)} \right\}, \left\{ w1 \rightarrow \frac{9}{280}, wC \rightarrow \frac{13}{420}, \mu C \rightarrow \sqrt{\frac{1}{26} \left( 13 + \sqrt{65} \right)} \right\} \right\}$$

(\* The fourth solution above belongs to the base set, the others coinside with the ordinates

resulting from symmetry operations. \*)

(\* Numerical

conditioning of the quadrature will be best if the weights are simmilar, the decimal values below show this to potentialy be a good quadrature \*)

N[%21[[4]], 16]

 $\{\texttt{w1} \rightarrow \texttt{0.03214285714285714}, \texttt{wC} \rightarrow \texttt{0.03095238095238095}, \texttt{\muC} \rightarrow \texttt{0.9000482411921158}\}$ 

Eqns = Table [caseA[wA, k] + caseB[wB, k] + case1[w1, k] == 
$$\frac{1}{2 k + 1}$$
, {k, 3}]  
 $\left\{\frac{8 w1}{3} + 2 wA + 4 wB == \frac{1}{3}, \frac{8 w1}{9} + 2 wA + 2 wB == \frac{1}{5}, \frac{8 w1}{27} + 2 wA + wB == \frac{1}{7}\right\}$ 

Solve[Eqns, {wA, wB, w1}]

$$\left\{\left\{\mathsf{wA} \rightarrow \frac{1}{21}, \mathsf{wB} \rightarrow \frac{4}{105}, \mathsf{w1} \rightarrow \frac{9}{280}\right\}\right\}$$

### N[%24, 16]

 $\{\{\text{wA} \rightarrow 0.04761904761904762, \text{wB} \rightarrow 0.0380952380952381, \text{w1} \rightarrow 0.03214285714285714\}\}$ 

(\* The above weights are also similar in magnatude. \*)

Eqns = Table [caseA[wA, k] + caseB[wB, k] + caseC[wC,  $\mu$ C, k] + case1[w1, k] ==  $\frac{1}{2k+1}$ , {k, 5}]

$$\left\{ \frac{8 \text{ w1}}{3} + 2 \text{ wA} + 4 \text{ wB} + 8 \text{ wC} = \frac{1}{3}, \frac{8 \text{ w1}}{9} + 2 \text{ wA} + 2 \text{ wB} + 8 \text{ wC} \left(\mu \text{C}^4 + (1 - \mu \text{C}^2)^2\right) = \frac{1}{5}, \frac{8 \text{ w1}}{27} + 2 \text{ wA} + \text{wB} + 8 \text{ wC} \left(\mu \text{C}^6 + (1 - \mu \text{C}^2)^3\right) = \frac{1}{7}, \frac{8 \text{ w1}}{81} + 2 \text{ wA} + \frac{\text{wB}}{2} + 8 \text{ wC} \left(\mu \text{C}^8 + (1 - \mu \text{C}^2)^4\right) = \frac{1}{9}, \frac{8 \text{ w1}}{243} + 2 \text{ wA} + \frac{\text{wB}}{4} + 8 \text{ wC} \left(\mu \text{C}^{10} + (1 - \mu \text{C}^2)^5\right) = \frac{1}{11} \right\}$$

quadABC1 = Solve[Eqns, {wA, wB, wC, w1,  $\mu$ C}]

{}

(\* TheABC1 case has no solutions \*)

Eqns = Table [caseA[wA, k] + caseB[wB, k] + case1[w1, k] + case2[w2,  $\mu$ 2, k] ==  $\frac{1}{2k+1}$ , {k, 5}]

$$\left\{ \frac{8 \text{ w1}}{3} + 8 \text{ w2} + 2 \text{ wA} + 4 \text{ wB} == \frac{1}{3}, \frac{8 \text{ w1}}{9} + 2 \text{ wA} + 2 \text{ wB} + 8 \text{ w2} \left(\mu 2^4 + \frac{1}{2} (1 - \mu 2^2)^2\right) == \frac{1}{5}, \frac{8 \text{ w1}}{27} + 2 \text{ wA} + \text{ wB} + 8 \text{ w2} \left(\mu 2^6 + \frac{1}{4} (1 - \mu 2^2)^3\right) == \frac{1}{7}, \frac{8 \text{ w1}}{81} + 2 \text{ wA} + \frac{\text{wB}}{2} + 8 \text{ w2} \left(\mu 2^8 + \frac{1}{8} (1 - \mu 2^2)^4\right) == \frac{1}{9}, \frac{8 \text{ w1}}{243} + 2 \text{ wA} + \frac{\text{wB}}{4} + 8 \text{ w2} \left(\mu 2^{10} + \frac{1}{16} (1 - \mu 2^2)^5\right) == \frac{1}{11} \right\}$$

quadAB12 = Solve [Eqns, {wA, wB, w1, w2,  $\mu$ 2}]

$$\left\{ \left\{ wA \rightarrow \frac{4}{315}, wB \rightarrow \frac{64}{2835}, w1 \rightarrow \frac{27}{1280}, w2 \rightarrow \frac{14641}{725760}, \mu2 \rightarrow -\frac{3}{\sqrt{11}} \right\}, \\ \left\{ wA \rightarrow \frac{4}{315}, wB \rightarrow \frac{64}{2835}, w1 \rightarrow \frac{27}{1280}, w2 \rightarrow \frac{14641}{725760}, \mu2 \rightarrow \frac{3}{\sqrt{11}} \right\} \right\}$$

(\* The second solution belongs

to the base set. Its numerical value shows it to potentialy be a good quadrature \*)

N[quadAB12[[2]], 16]

 $\{ w \mathbb{A} \rightarrow 0.0126984126984127, \ w \mathbb{B} \rightarrow 0.02257495590828924, \ w \mathbb{1} \rightarrow 0.02109375, \\ w \mathbb{2} \rightarrow 0.02017333553791887, \ \mu \mathbb{2} \rightarrow 0.904534033733291 \}$ 

Eqns = Table [caseA[wA, k] + caseC[wC, 
$$\mu$$
C, k] + case2[w2,  $\mu$ 2, k] ==  $\frac{1}{2 k + 1}$ , {k, 5}  
{8 w2 + 2 wA + 8 wC ==  $\frac{1}{3}$ , 2 wA + 8 w2 ( $\mu$ 2<sup>4</sup> +  $\frac{1}{2}$  (1 -  $\mu$ 2<sup>2</sup>)<sup>2</sup>) + 8 wC ( $\mu$ C<sup>4</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>2</sup>) ==  $\frac{1}{5}$ ,  
2 wA + 8 w2 ( $\mu$ 2<sup>6</sup> +  $\frac{1}{4}$  (1 -  $\mu$ 2<sup>2</sup>)<sup>3</sup>) + 8 wC ( $\mu$ C<sup>6</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>3</sup>) ==  $\frac{1}{7}$ ,  
2 wA + 8 w2 ( $\mu$ 2<sup>8</sup> +  $\frac{1}{8}$  (1 -  $\mu$ 2<sup>2</sup>)<sup>4</sup>) + 8 wC ( $\mu$ C<sup>6</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>4</sup>) ==  $\frac{1}{9}$ ,  
2 wA + 8 w2 ( $\mu$ 2<sup>10</sup> +  $\frac{1}{16}$  (1 -  $\mu$ 2<sup>2</sup>)<sup>5</sup>) + 8 wC ( $\mu$ C<sup>10</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>5</sup>) ==  $\frac{1}{11}$ }

quadAC2 = Solve[Eqns, {wA, wC, w2,  $\mu$ C,  $\mu$ 2}]

(\* Large output deleted \*)

(\* I'll apply the Filter function to better see what I have \*)

Filter[quadAC2, {wA, wC, w2,  $\mu$ C,  $\mu$ 2}]

$$\{ \text{Null, Null, Null, Null, } \{ wA \rightarrow \frac{2(41 - 2\sqrt{22})}{2835}, wC \rightarrow \frac{44 + 13\sqrt{22}}{5670}, \\ w2 \rightarrow \frac{11(55 - 4\sqrt{22})}{22680}, \muC \rightarrow \sqrt{\frac{1}{66}(33 - \sqrt{33(-11 + 4\sqrt{22})})}, \mu2 \rightarrow \sqrt{\frac{1}{33}(11 + 2\sqrt{22})} \}, \\ \text{Null, Null, Null, Null, } \{ wA \rightarrow \frac{2(41 - 2\sqrt{22})}{2835}, wC \rightarrow \frac{44 + 13\sqrt{22}}{5670}, w2 \rightarrow \frac{11(55 - 4\sqrt{22})}{22680}, \\ \muC \rightarrow \sqrt{\frac{1}{66}(33 + \sqrt{33(-11 + 4\sqrt{22})})}, \mu2 \rightarrow \sqrt{\frac{1}{33}(11 + 2\sqrt{22})} \}, \\ \text{Null, Null, Null, Null, Null, Null} \}$$

(\* Let's see what the numerical values are \*)

### N[%%, 16]

 $\begin{aligned} & \{ & \text{Null, Null, } \{ & \text{wA} \rightarrow 0.02230629169689816, \ \text{wC} \rightarrow 0.01851418075444525, \\ & \text{w2} \rightarrow 0.01757591298799687, \ \mu\text{C} \rightarrow 0.5074563057138757, \ \mu\text{2} \rightarrow 0.7858759158676477 \}, \\ & \text{Null, Null, Null, } \{ & \text{wA} \rightarrow 0.02230629169689816, \ \text{wC} \rightarrow 0.01851418075444525, \\ & \text{w2} \rightarrow 0.01757591298799687, \ \mu\text{C} \rightarrow 0.8616774905910132, \ \mu\text{2} \rightarrow 0.7858759158676477 \}, \\ & \text{Null, Null, Null, Null, Null, Null, Null, Null, Null, Null } \end{aligned}$ 

(\* Since  $\mu 2$  is greater then  $\frac{1}{\sqrt{3}}$  this is a case2 ordinate. The second non null solution is in the base set. \*)

Eqns = Table [caseB[wB, k] + caseC[wC, 
$$\mu$$
C, k] + case2[w2,  $\mu$ 2, k] ==  $\frac{1}{2 k + 1}$ , {k, 5}]  
{8 w2 + 4 wB + 8 wC ==  $\frac{1}{3}$ , 2 wB + 8 w2 ( $\mu$ 2<sup>4</sup> +  $\frac{1}{2}$  (1 -  $\mu$ 2<sup>2</sup>)<sup>2</sup>) + 8 wC ( $\mu$ C<sup>4</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>2</sup>) ==  $\frac{1}{5}$ ,  
wB + 8 w2 ( $\mu$ 2<sup>6</sup> +  $\frac{1}{4}$  (1 -  $\mu$ 2<sup>2</sup>)<sup>3</sup>) + 8 wC ( $\mu$ C<sup>6</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>3</sup>) ==  $\frac{1}{7}$ ,  
 $\frac{\text{wB}}{2}$  + 8 w2 ( $\mu$ 2<sup>8</sup> +  $\frac{1}{8}$  (1 -  $\mu$ 2<sup>2</sup>)<sup>4</sup>) + 8 wC ( $\mu$ C<sup>8</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>4</sup>) ==  $\frac{1}{9}$ ,  
 $\frac{\text{wB}}{4}$  + 8 w2 ( $\mu$ 2<sup>10</sup> +  $\frac{1}{16}$  (1 -  $\mu$ 2<sup>2</sup>)<sup>5</sup>) + 8 wC ( $\mu$ C<sup>10</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>5</sup>) ==  $\frac{1}{11}$ }

quadBC2 = Solve[Eqns, {wB, wC, w2,  $\mu$ C,  $\mu$ 2}]

### (\* Large output deleted \*)

### Filter[quadBC2, {wB, wC, w2, $\mu$ C, $\mu$ 2}]

{Null, Null, Null, Null, Null, Null, Null, Null, Null, Null, Null,

$$\left\{ wB \rightarrow \frac{16 \left(-190 + 169 \sqrt{22}\right)}{526995}, wC \rightarrow \frac{169565 - 5933 \sqrt{22}}{9485910}, w2 \rightarrow \frac{11 \left(55 - 4 \sqrt{22}\right)}{22680}, \\ \mu C \rightarrow \sqrt{\frac{4521 - \sqrt{4521 \left(2761 + 52 \sqrt{22}\right)}}{9042}}, \mu 2 \rightarrow \sqrt{\frac{1}{33} \left(11 + 2 \sqrt{22}\right)} \right\}, \text{Null, Null, Null, Null, Null, Null, } \\ \text{Null, } \left\{ wB \rightarrow \frac{16 \left(-190 + 169 \sqrt{22}\right)}{526995}, wC \rightarrow \frac{169565 - 5933 \sqrt{22}}{9485910}, w2 \rightarrow \frac{11 \left(55 - 4 \sqrt{22}\right)}{22680} \right\} \\ \mu C \rightarrow \sqrt{\frac{4521 + \sqrt{4521 \left(2761 + 52 \sqrt{22}\right)}}{9042}}, \mu 2 \rightarrow \sqrt{\frac{1}{33} \left(11 + 2 \sqrt{22}\right)} \right\}$$

### N[%, 16]

(\* This quadrature also has promising weights, the last solution is in the base set. \*)

Eqns = Table [caseC[wC, 
$$\mu$$
C, k] + case4[w4,  $\mu$ 4,  $\eta$ 4, k] ==  $\frac{1}{2 k + 1}$ , {k, 5}]  
{16 w4 + 8 wC ==  $\frac{1}{3}$ , 16 w4 ( $\eta$ 4<sup>4</sup> +  $\mu$ 4<sup>4</sup> + (1 -  $\eta$ 4<sup>2</sup> -  $\mu$ 4<sup>2</sup>)<sup>2</sup>) + 8 wC ( $\mu$ C<sup>4</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>2</sup>) ==  $\frac{1}{5}$ ,  
16 w4 ( $\eta$ 4<sup>6</sup> +  $\mu$ 4<sup>6</sup> + (1 -  $\eta$ 4<sup>2</sup> -  $\mu$ 4<sup>2</sup>)<sup>3</sup>) + 8 wC ( $\mu$ C<sup>6</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>3</sup>) ==  $\frac{1}{7}$ ,  
16 w4 ( $\eta$ 4<sup>8</sup> +  $\mu$ 4<sup>8</sup> + (1 -  $\eta$ 4<sup>2</sup> -  $\mu$ 4<sup>2</sup>)<sup>4</sup>) + 8 wC ( $\mu$ C<sup>6</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>4</sup>) ==  $\frac{1}{9}$ ,  
16 w4 ( $\eta$ 4<sup>10</sup> +  $\mu$ 4<sup>10</sup> + (1 -  $\eta$ 4<sup>2</sup> -  $\mu$ 4<sup>2</sup>)<sup>5</sup>) + 8 wC ( $\mu$ C<sup>10</sup> + (1 -  $\mu$ C<sup>2</sup>)<sup>5</sup>) ==  $\frac{1}{11}$ }

quadC4 = Solve[Eqns, {wC, w4,  $\mu$ C,  $\mu$ 4,  $\eta$ 4}]

(\* Large output deleted \*)

(\* The decimal equivalent after filtering is shown below \*)

N[quadC4, 25]

### Filter[%, {wC, w4, $\mu$ C, $\mu$ 4, $\eta$ 4}]

```
{Null, Null, Null,
Null, Null, Null, Null, Null, Null, Null, Null, Null, Null, Null, Null, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.326499776057011143395600,
  \texttt{w4} \rightarrow \texttt{0.0122967479674796747967479674797}, \ \mu\texttt{4} \rightarrow \texttt{0.2856038324721905188708181} + -\texttt{0.} \times \texttt{10}^{-\texttt{56}} \texttt{I},
  n4 \rightarrow 0.757316019511794446172265 + -0. \times 10^{-57} \text{ I}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.326499776057011143395600,
  \texttt{w4} \rightarrow \texttt{0.0122967479674796747967479674797}, \ \mu4 \rightarrow \texttt{0.587284341241964586253482} + -\texttt{0.} \times \texttt{10}^{-\texttt{60}} \texttt{I},
  n4 \rightarrow 0.757316019511794446172265 + -0. \times 10^{-57} \text{ I}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.326499776057011143395600,
  w4 \rightarrow 0.01229674796747967479674797, \mu4 \rightarrow 0.587284341241964586253482+0. \times 10^{-60} I,
  \eta 4 \rightarrow 0.2856038324721905188708181 + -0. \times 10^{-50} \, \mathrm{I}\}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.326499776057011143395600,
  w4 \rightarrow 0.01229674796747967479674797, \mu4 \rightarrow 0.757316019511794446172265 + 0. \times\,10^{-60} I,
  n4 \rightarrow 0.2856038324721905188708181 + -0. \times 10^{-50} \text{ I}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.326499776057011143395600,
  w4 \rightarrow 0.01229674796747967479674797, \mu4 \rightarrow 0.2856038324721905188708181 + 0. \times 10^{-60} I,
  \eta 4 \rightarrow 0.587284341241964586253482 + -0. \times 10^{-51} \, {\rm I}\}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.326499776057011143395600,
  w4 \rightarrow 0.01229674796747967479674797, \mu4 \rightarrow 0.757316019511794446172265 + 0. \times 10^{-55} I,
  \eta 4 \rightarrow 0.587284341241964586253482 + -0. \times 10^{-51} \, \mathrm{I}\},
 Null, Null,
 Null, Null, Null, Null, Null, Null, Null, Null, Null, Null, Null, Null, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \mu C \rightarrow 0.945197279003024623550787,
  \texttt{w4} \rightarrow \texttt{0.0122967479674796747967479674797}, \ \mu\texttt{4} \rightarrow \texttt{0.2856038324721905188708181} + \texttt{-0.} \times \texttt{10}^{-\texttt{56}} \texttt{I},
  \eta 4 \rightarrow 0.757316019511794446172265 + -0. \times 10^{-57} \text{ I}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.945197279003024623550787,
  w4 \rightarrow 0.01229674796747967479674797, \mu4 \rightarrow 0.587284341241964586253482+-0.\times\,10^{-60} I,
  \eta 4 \rightarrow 0.757316019511794446172265 + -0. \times 10^{-57} \, {\rm I}\}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.945197279003024623550787,
  w4 \rightarrow 0.01229674796747967479674797, \mu4 \rightarrow 0.587284341241964586253482 + 0. \times 10^{-60} I,
  \eta 4 \rightarrow 0.2856038324721905188708181 + -0. \times 10^{-50} I}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.945197279003024623550787,
  w4 \rightarrow 0.01229674796747967479674797, \mu4 \rightarrow 0.757316019511794446172265 + 0. \times 10^{-60} I,
   n4 \rightarrow 0.2856038324721905188708181 + -0. \times 10^{-50} I}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.945197279003024623550787,
  \texttt{w4} \rightarrow \texttt{0.0122967479674796747967479674797}, \ \mu 4 \rightarrow \texttt{0.2856038324721905188708181} + \texttt{0.} \times \texttt{10}^{-\texttt{60}} \texttt{I},
   n4 \rightarrow 0.587284341241964586253482 + -0. \times 10^{-51} \text{ I}, Null, Null,
 Null, {wC \rightarrow 0.01707317073170731707317073, \muC \rightarrow 0.945197279003024623550787,
  \texttt{w4} \rightarrow \texttt{0.01229674796747967479674797}, \ \texttt{\mu4} \rightarrow \texttt{0.757316019511794446172265} + \texttt{0.} \times \texttt{10}^{-\texttt{55}} \texttt{I},
   n4 \rightarrow 0.587284341241964586253482 + -0. \times 10^{-51} \text{ I} \}
```

(\* The filter didi no eliminate the imaginary portions above because the are artifacts of rounding error and not truely imaginary values. Higher precision math reduces the imaginary component further. Carfull inspection shows that all the above solutions are reflections of the last one, which is in the base set. \*)  $\begin{aligned} & \text{Eqns = Table} \Big[ \text{casel} [w1, k] + \text{caseC} [wC, \mu C, k] + \text{case2} [w2, \mu 2, k] == \frac{1}{2 \, k + 1}, \ \{k, 5\} \Big] \\ & \left\{ \frac{8 \, \text{w1}}{3} + 8 \, \text{w2} + 8 \, \text{wC} == \frac{1}{3}, \ \frac{8 \, \text{w1}}{9} + 8 \, \text{w2} \left( \mu 2^4 + \frac{1}{2} \left( 1 - \mu 2^2 \right)^2 \right) + 8 \, \text{wC} \left( \mu C^4 + \left( 1 - \mu C^2 \right)^2 \right) == \frac{1}{5}, \\ & \frac{8 \, \text{w1}}{27} + 8 \, \text{w2} \left( \mu 2^6 + \frac{1}{4} \left( 1 - \mu 2^2 \right)^3 \right) + 8 \, \text{wC} \left( \mu C^6 + \left( 1 - \mu C^2 \right)^3 \right) == \frac{1}{7}, \\ & \frac{8 \, \text{w1}}{81} + 8 \, \text{w2} \left( \mu 2^8 + \frac{1}{8} \left( 1 - \mu 2^2 \right)^4 \right) + 8 \, \text{wC} \left( \mu C^8 + \left( 1 - \mu C^2 \right)^4 \right) == \frac{1}{9}, \\ & \frac{8 \, \text{w1}}{243} + 8 \, \text{w2} \left( \mu 2^{10} + \frac{1}{16} \left( 1 - \mu 2^2 \right)^5 \right) + 8 \, \text{wC} \left( \mu C^{10} + \left( 1 - \mu C^2 \right)^5 \right) == \frac{1}{11} \end{aligned}$ 

quad1C2 = Solve[Eqns, {w1, wC, w2,  $\mu$ C,  $\mu$ 2}]

(\* Large output deleted \*)

N[quad1C2, 40]

Filter[%, {w1, wC, w2, µC, µ2}]

```
{Null, Null, Null, {w1 \rightarrow 0.03139662070471866287073752871645825335788+0. ×10<sup>-52</sup> I, wC \rightarrow 0.03118946912313230360638905694603843725377+0. ×10<sup>-53</sup> I, w2 \rightarrow 0.00001165730862814210336510014847547829359805+0. ×10<sup>-53</sup> I, \muC \rightarrow 0.440124599527176955730943844695059472227+-0. ×10<sup>-68</sup> I, \mu2 \rightarrow 1.482898731354957472518351725181578455870+-0. ×10<sup>-68</sup> I}, Null, Null, Null, {w1 \rightarrow 0.03139662070471866287073752871645825335788+0. ×10<sup>-52</sup> I, wC \rightarrow 0.03118946912313230360638905694603843725377+0. ×10<sup>-53</sup> I, w2 \rightarrow 0.00001165730862814210336510014847547829359805+0. ×10<sup>-53</sup> I, w2 \rightarrow 0.00001165730862814210336510014847547829359805+0. ×10<sup>-53</sup> I, w2 \rightarrow 0.0897936710960768162219787279180959387596+-0. ×10<sup>-68</sup> I, \mu2 \rightarrow 1.482898731354957472518351725181578455870+-0. ×10<sup>-68</sup> I, w2 \rightarrow 0.0001165730862814210336510014847547829359805+0. ×10<sup>-53</sup> I, w2 \rightarrow 0.100001165730862814210336510014847547829359805+0. ×10<sup>-53</sup> I, w2 \rightarrow 0.48298731354957472518351725181578455870+-0. ×10<sup>-68</sup> I, m2 \rightarrow 1.482898731354957472518351725181578455870+-0. ×10<sup>-68</sup> I, mull, Null, Nu
```

(\* From the above two sets of output,

we see that there are no acceptabe 1C2 quadratures. The filtered output shows quadratrues with cosines greater than one and from the unfiltered solutions we see there are quadratres with valid angles but negative weights. \*)

Eqns = Table [caseA[wA, k] + caseB[wB, k] + case4[w4,  $\mu$ 4,  $\eta$ 4, k] ==  $\frac{1}{2 k + 1}$ , {k, 5}] {16 w4 + 2 wA + 4 wB ==  $\frac{1}{2}$ , 2 wA + 2 wB + 16 w4 ( $\eta$ 4<sup>4</sup> +  $\mu$ 4<sup>4</sup> + ( $1 - \eta$ 4<sup>2</sup> -  $\mu$ 4<sup>2</sup>)<sup>2</sup>) ==  $\frac{1}{5}$ ,

$$2 \text{ wA} + \text{wB} + 16 \text{ w4} \left(\eta 4^{6} + \mu 4^{6} + (1 - \eta 4^{2} - \mu 4^{2})^{3}\right) == \frac{1}{7},$$

$$2 \text{ wA} + \frac{\text{wB}}{2} + 16 \text{ w4} \left(\eta 4^{8} + \mu 4^{8} + (1 - \eta 4^{2} - \mu 4^{2})^{4}\right) == \frac{1}{9},$$

$$2 \text{ wA} + \frac{\text{wB}}{4} + 16 \text{ w4} \left(\eta 4^{10} + \mu 4^{10} + (1 - \eta 4^{2} - \mu 4^{2})^{5}\right) == \frac{1}{11}$$

quadBC2 = Solve[Eqns, {wA, wB, w4,  $\mu$ 4,  $\eta$ 4}]

(\* Large output deleted \*)

N[%, 40]

### Filter[\$, {wA, wB, w4, $\mu$ 4, $\eta$ 4}]

{Null, Null, Null

(\* there are no valid AB4 quadratures \*)

(\* The next series of quadratures are of order 7 \*)

Eqns = Table

 $caseA[wA, k] + caseB[wB, k] + caseC[wC, \muC, k] + case1[w1, k] + case2[w2, \mu2, k] = \frac{1}{2k+1}$ 

 $\{k, 7\}$ 

$$\begin{cases} \frac{8 \text{ w1}}{3} + 8 \text{ w2} + 2 \text{ wA} + 4 \text{ wB} + 8 \text{ wC} == \frac{1}{3}, \\ \frac{8 \text{ w1}}{9} + 2 \text{ wA} + 2 \text{ wB} + 8 \text{ w2} \left(\mu 2^4 + \frac{1}{2} (1 - \mu 2^2)^2\right) + 8 \text{ wC} \left(\mu C^4 + (1 - \mu C^2)^2\right) == \frac{1}{5}, \\ \frac{8 \text{ w1}}{27} + 2 \text{ wA} + \text{wB} + 8 \text{ w2} \left(\mu 2^6 + \frac{1}{4} (1 - \mu 2^2)^3\right) + 8 \text{ wC} \left(\mu C^6 + (1 - \mu C^2)^3\right) == \frac{1}{7}, \\ \frac{8 \text{ w1}}{81} + 2 \text{ wA} + \frac{\text{wB}}{2} + 8 \text{ w2} \left(\mu 2^8 + \frac{1}{8} (1 - \mu 2^2)^4\right) + 8 \text{ wC} \left(\mu C^8 + (1 - \mu C^2)^4\right) == \frac{1}{9}, \\ \frac{8 \text{ w1}}{243} + 2 \text{ wA} + \frac{\text{wB}}{4} + 8 \text{ w2} \left(\mu 2^{10} + \frac{1}{16} (1 - \mu 2^2)^5\right) + 8 \text{ wC} \left(\mu C^{10} + (1 - \mu C^2)^5\right) == \frac{1}{11}, \\ \frac{8 \text{ w1}}{729} + 2 \text{ wA} + \frac{\text{wB}}{8} + 8 \text{ w2} \left(\mu 2^{12} + \frac{1}{32} (1 - \mu 2^2)^6\right) + 8 \text{ wC} \left(\mu C^{12} + (1 - \mu C^2)^6\right) == \frac{1}{13}, \\ \frac{8 \text{ w1}}{2187} + 2 \text{ wA} + \frac{\text{wB}}{16} + 8 \text{ w2} \left(\mu 2^{14} + \frac{1}{64} (1 - \mu 2^2)^7\right) + 8 \text{ wC} \left(\mu C^{14} + (1 - \mu C^2)^7\right) == \frac{1}{15} \end{cases}$$

quadABC12 = Solve[Eqns, {wA, wB, wC, w1, w2,  $\mu$ C,  $\mu$ 2}]

(\* Large output deleted \*)

N[quadABC12, 16]

### Filter[%, {wA, wB, w1, wC, w2, $\mu$ C, $\mu$ 2}]

 $\begin{aligned} &\{ \text{Null, Null, Null, } \{ \text{wA} \rightarrow 0.009048188830155413, \\ &\text{wB} \rightarrow 0.02103246043742795, \\ &\text{wl} \rightarrow 0.01827941392341811, \\ &\text{wC} \rightarrow 0.006451491538566835, \\ &\text{w2} \rightarrow 0.0163437597273743, \\ &\mu\text{C} \rightarrow 0.2979519566503113, \\ &\mu\text{C} \rightarrow 0.8753170875981718 \}, \\ &\text{Null, Null, Null, } \{ \text{wA} \rightarrow 0.009048188830155413, \\ &\text{wB} \rightarrow 0.02103246043742795, \\ &\text{wl} \rightarrow 0.01827941392341811, \\ &\text{wC} \rightarrow 0.006451491538566835, \\ &\text{w2} \rightarrow 0.0163437597273743, \end{aligned}$ 

 $\mu \rm C \rightarrow 0.9545808669401723, \ \mu \rm 2 \rightarrow 0.8753170875981718\},$  Null, Null

(\* The second of the above valid quadratures is in the base set, the other is a reflection. \*)

Eqns = Table [caseA[wA, k] + caseC[wC,  $\mu$ C, k] + case2[w2,  $\mu$ 2, k] + case2[w3,  $\mu$ 3, k] ==  $\frac{1}{2k+1}$ , {k, 7}]

$$\left\{ 8 \text{ w2} + 8 \text{ w3} + 2 \text{ wA} + 8 \text{ wC} = \frac{1}{3}, \\ 2 \text{ wA} + 8 \text{ w2} \left( \mu 2^4 + \frac{1}{2} \left( 1 - \mu 2^2 \right)^2 \right) + 8 \text{ w3} \left( \mu 3^4 + \frac{1}{2} \left( 1 - \mu 3^2 \right)^2 \right) + 8 \text{ wC} \left( \mu C^4 + \left( 1 - \mu C^2 \right)^2 \right) = \frac{1}{5}, \\ 2 \text{ wA} + 8 \text{ w2} \left( \mu 2^6 + \frac{1}{4} \left( 1 - \mu 2^2 \right)^3 \right) + 8 \text{ w3} \left( \mu 3^6 + \frac{1}{4} \left( 1 - \mu 3^2 \right)^3 \right) + 8 \text{ wC} \left( \mu C^6 + \left( 1 - \mu C^2 \right)^3 \right) = \frac{1}{7}, \\ 2 \text{ wA} + 8 \text{ w2} \left( \mu 2^8 + \frac{1}{8} \left( 1 - \mu 2^2 \right)^4 \right) + 8 \text{ w3} \left( \mu 3^8 + \frac{1}{8} \left( 1 - \mu 3^2 \right)^4 \right) + 8 \text{ wC} \left( \mu C^6 + \left( 1 - \mu C^2 \right)^4 \right) = \frac{1}{9}, \\ 2 \text{ wA} + 8 \text{ w2} \left( \mu 2^{10} + \frac{1}{16} \left( 1 - \mu 2^2 \right)^5 \right) + 8 \text{ w3} \left( \mu 3^{10} + \frac{1}{16} \left( 1 - \mu 3^2 \right)^5 \right) + 8 \text{ wC} \left( \mu C^{10} + \left( 1 - \mu C^2 \right)^5 \right) = \frac{1}{11}, \\ 2 \text{ wA} + 8 \text{ w2} \left( \mu 2^{12} + \frac{1}{32} \left( 1 - \mu 2^2 \right)^6 \right) + 8 \text{ w3} \left( \mu 3^{12} + \frac{1}{32} \left( 1 - \mu 3^2 \right)^6 \right) + 8 \text{ wC} \left( \mu C^{12} + \left( 1 - \mu C^2 \right)^6 \right) = \frac{1}{13}, \\ 2 \text{ wA} + 8 \text{ w2} \left( \mu 2^{14} + \frac{1}{64} \left( 1 - \mu 2^2 \right)^7 \right) + 8 \text{ w3} \left( \mu 3^{14} + \frac{1}{64} \left( 1 - \mu 3^2 \right)^7 \right) + 8 \text{ wC} \left( \mu C^{14} + \left( 1 - \mu C^2 \right)^7 \right) = \frac{1}{15} \right\}$$

quadAC23 = Solve[Eqns, {wA, wC, w2, w3,  $\mu$ C,  $\mu$ 2,  $\mu$ 3}]

\$Aborted

```
(* The computer was not able to solve this exactly,
I will now try the numerical function *)
```

```
quadAC23 = NSolve[Eqns, \{wA, wC, w2, w3, \mu C, \mu 2, \mu 3\}]
```

\$Aborted

(\* The computer was not able to find this solution directly either. I will lend some assistance \*)

```
Eqns14 = Table
```

caseA[wA, k] + caseC[wC,  $\mu$ C, k] + case2[w2,  $\mu$ 2, k] + case2[w3,  $\mu$ 3, k] ==  $\frac{1}{2k+1}$ , {k, 4}

$\left\{8 \text{ w2} + 8 \text{ w3} + 2 \text{ wA} + 8 \text{ wC} = \frac{1}{3},\right.$
$2 \text{ wA} + 8 \text{ w2} \left(\mu 2^4 + \frac{1}{2} \left(1 - \mu 2^2\right)^2\right) + 8 \text{ w3} \left(\mu 3^4 + \frac{1}{2} \left(1 - \mu 3^2\right)^2\right) + 8 \text{ wC} \left(\mu C^4 + \left(1 - \mu C^2\right)^2\right) = \frac{1}{5},$
$2 \text{ wA} + 8 \text{ w2} \left(\mu 2^{6} + \frac{1}{4} \left(1 - \mu 2^{2}\right)^{3}\right) + 8 \text{ w3} \left(\mu 3^{6} + \frac{1}{4} \left(1 - \mu 3^{2}\right)^{3}\right) + 8 \text{ wC} \left(\mu C^{6} + \left(1 - \mu C^{2}\right)^{3}\right) = \frac{1}{7},$
$2 \text{ wA} + 8 \text{ w2} \left(\mu 2^{8} + \frac{1}{8} \left(1 - \mu 2^{2}\right)^{4}\right) + 8 \text{ w3} \left(\mu 3^{8} + \frac{1}{8} \left(1 - \mu 3^{2}\right)^{4}\right) + 8 \text{ wC} \left(\mu C^{8} + \left(1 - \mu C^{2}\right)^{4}\right) = \frac{1}{9}$

Solve[Eqns14, {wA, wC, w2, w3}]

 $\{wA, wC, w2, w3\} = \{wA, wC, w2, w3\} / . \%[[1]];$ 

Expand[wA];

Together[%];

Cancel[%];

```
 \left( - (1 - 2\mu C^2)^2 (3 + 9\mu 3^6 - 29\mu C^2 + 29\mu C^4 + \mu 3^2 (-1 + 54\mu C^2 - 54\mu C^4) + \mu 3^4 (-11 - 21\mu C^2 + 21\mu C^4) \right) + 9\mu 2^6 (-(1 - 2\mu C^2)^2 + \mu 3^4 (-39 + 210\mu C^2 - 210\mu C^4) + 21\mu 3^6 (1 - 5\mu C^2 + 5\mu C^4) + \mu 3^2 (19 - 105\mu C^2 + 105\mu C^4) \right) + \mu 2^2 ((1 - 2\mu C^2)^2 (1 - 54\mu C^2 + 54\mu C^4) + 9\mu 3^6 (19 - 105\mu C^2 + 105\mu C^4) + \mu 3^4 (-281 + 1716\mu C^2 - 2556\mu C^4 + 1680\mu C^6 - 840\mu C^8) + \mu 3^2 (109 - 653\mu C^2 + 989\mu C^4 - 672\mu C^6 + 336\mu C^8)) - \mu 2^4 ((1 - 2\mu C^2)^2 (-11 - 21\mu C^2 + 21\mu C^4) + 27\mu 3^6 (13 - 70\mu C^2 + 70\mu C^4) - 3\mu 3^4 (207 - 1225\mu C^2 + 1645\mu C^4 - 840\mu C^6 + 420\mu C^8) + \mu 3^2 (281 - 1716\mu C^2 + 2556\mu C^4 - 1680\mu C^6 + 840\mu C^8)) \right) \right) 
 \left( \left( 630 (-1 + \mu 2^2) (-1 + \mu 3^2)\mu C^2 (-1 + \mu C^2) (9\mu 2^4\mu 3^2 (-1 + \mu 3^2) + (-1 + \mu 3^2) (1 - 2\mu C^2)^2 + \mu 2^2 (-9\mu 3^4 + (1 - 2\mu C^2)^2 + 4\mu 3^2 (2 - 3\mu C^2 + 3\mu C^4)) \right) \right)
```

wA = %27;

Expand[wC];

Together[%];

Cancel[%];

PowerExpand[%];

Simplify[%]

 $(3 + 2 \mu 3^{2} - 9 \mu 3^{4} - 9 \mu 2^{4} (1 - 18 \mu 3^{2} + 21 \mu 3^{4}) + 2 \mu 2^{2} (1 - 54 \mu 3^{2} + 81 \mu 3^{4})) /$  $(2520 \mu C^{2} (-1 + \mu C^{2}) (9 \mu 2^{4} \mu 3^{2} (-1 + \mu 3^{2}) + (-1 + \mu 3^{2}) (1 - 2 \mu C^{2})^{2} +$  $\mu 2^{2} (-9 \mu 3^{4} + (1 - 2 \mu C^{2})^{2} + 4 \mu 3^{2} (2 - 3 \mu C^{2} + 3 \mu C^{4}))))$ 

wC = %33;

Expand[w2];

Together[%];

Cancel[%38];

```
 \left( -9\,\mu 3^{6} + (1 - 2\,\mu C^{2})^{2} - 3\,\mu 3^{2} \left( 5 - 24\,\mu C^{2} + 24\,\mu C^{4} \right) + \mu 3^{4} \left( 23 - 84\,\mu C^{2} + 84\,\mu C^{4} \right) \right) / \left( 630 \left( -1 + \mu 2^{2} \right) \left( \mu 2^{2} - \mu 3^{2} \right) \left( 9\,\mu 2^{4}\,\mu 3^{2} \left( -1 + \mu 3^{2} \right) + \left( -1 + \mu 3^{2} \right) \left( 1 - 2\,\mu C^{2} \right)^{2} + \mu 2^{2} \left( -9\,\mu 3^{4} + \left( 1 - 2\,\mu C^{2} \right)^{2} + 4\,\mu 3^{2} \left( 2 - 3\,\mu C^{2} + 3\,\mu C^{4} \right) \right) \right) \right)
```

w2 = %;

### Expand[w3];

Together[%];

Cancel[%];

PowerExpand[%];

Simplify[%]

$$\left(9\,\mu 2^{6} - (1 - 2\,\mu C^{2})^{2} + \mu 2^{4} (-23 + 84\,\mu C^{2} - 84\,\mu C^{4}) + 3\,\mu 2^{2} (5 - 24\,\mu C^{2} + 24\,\mu C^{4})\right) / \left(630\,(\mu 2^{2} - \mu 3^{2}) (-1 + \mu 3^{2}) (9\,\mu 2^{4}\,\mu 3^{2} (-1 + \mu 3^{2}) + (-1 + \mu 3^{2}) (1 - 2\,\mu C^{2})^{2} + \mu 2^{2} (-9\,\mu 3^{4} + (1 - 2\,\mu C^{2})^{2} + 4\,\mu 3^{2} (2 - 3\,\mu C^{2} + 3\,\mu C^{4})\right)\right)$$

w3 = %;

```
lhs5 = caseA[wA, 5] + caseC[wC, \muC, 5] + case2[w2, \mu2, 5] + case2[w3, \mu3, 5];
```

Expand[%];

Together[%];

Cancel[%];

```
 \left( \left( -23 + 25\,\mu^{3^{2}} - 5\,\mu^{3^{4}} + 3\,\mu^{3^{6}} \right) \left( 1 - 2\,\mu^{2^{2}} \right)^{2} - 3\,\mu^{2^{6}} \left( 9\,\mu^{3^{6}} - \left( 1 - 2\,\mu^{2^{2}} \right)^{2} + \mu^{3^{4}} \left( -23 + 84\,\mu^{2^{2}} - 84\,\mu^{2^{4}} \right) + 3\,\mu^{3^{2}} \left( 5 - 24\,\mu^{2^{2}} + 24\,\mu^{2^{4}} \right) \right) + \mu^{2^{4}} \left( -5\,\left( 1 - 2\,\mu^{2^{2}} \right)^{2} + \mu^{3^{4}} \left( 65 + 636\,\mu^{2^{2}} - 636\,\mu^{2^{4}} \right) + 3\,\mu^{3^{6}} \left( 23 - 84\,\mu^{2^{2}} + 84\,\mu^{2^{4}} \right) + 3\,\mu^{3^{2}} \left( -43 - 124\,\mu^{2^{2}} + 124\,\mu^{2^{4}} \right) \right) + \mu^{2^{2}} \left( 25\,\left( 1 - 2\,\mu^{2^{2}} \right)^{2} - 9\,\mu^{3^{6}} \left( 5 - 24\,\mu^{2^{2}} + 24\,\mu^{2^{4}} \right) + 3\,\mu^{3^{4}} \left( -43 - 124\,\mu^{2^{2}} + 124\,\mu^{2^{4}} \right) \right) \right) \right) \right) 
 \left( 252\,\left( 9\,\mu^{2^{4}}\,\mu^{3^{2}} \left( -1 + \mu^{3^{2}} \right) + \left( -1 + \mu^{3^{2}} \right) \left( 1 - 2\,\mu^{2^{2}} \right)^{2} + \mu^{2^{2}} \left( -9\,\mu^{3^{4}} + \left( 1 - 2\,\mu^{2^{2}} \right)^{2} + 4\,\mu^{3^{2}} \left( 2 - 3\,\mu^{2^{2}} + 3\,\mu^{2^{4}} \right) \right) \right) \right) \right)
```

```
1hs5 = %;
```

```
lhs6 = caseA[wA, 6] + caseC[wC, \muC, 6] + case2[w2, \mu2, 6] + case2[w3, \mu3, 6];
```

Expand[%];

Together[%];

Cancel[%];

PowerExpand[%];

Simplify[%]

```
 \left( \left(1 - 2\,\mu\text{C}^2\right)^2 \left(27\,\mu\text{3}^6 + 33\,\mu\text{3}^8 + 2\left(-97 - 6\,\mu\text{C}^2 + 6\,\mu\text{C}^4\right) + \mu\text{3}^2 \left(245 - 8\,\mu\text{C}^2 + 8\,\mu\text{C}^4\right) - 3\,\mu\text{3}^4 \left(37 - 12\,\mu\text{C}^2 + 12\,\mu\text{C}^4\right) \right) - 3\,\mu\text{2}^8 \left(9\,\mu\text{3}^6 - \left(1 - 2\,\mu\text{C}^2\right)^2 + \mu\text{3}^4 \left(-23 + 84\,\mu\text{C}^2 - 84\,\mu\text{C}^4\right) + 3\,\mu\text{3}^2 \left(5 - 24\,\mu\text{C}^2 + 24\,\mu\text{C}^4\right) \right) - 3\,\mu\text{2}^6 \left(9 + 11\,\mu\text{3}^2\right) \left(9\,\mu\text{3}^6 - \left(1 - 2\,\mu\text{C}^2\right)^2 + \mu\text{3}^4 \left(-23 + 84\,\mu\text{C}^2 - 84\,\mu\text{C}^4\right) + 3\,\mu\text{3}^2 \left(5 - 24\,\mu\text{C}^2 + 24\,\mu\text{C}^4\right) \right) + \mu^2^2 \left(\left(1 - 2\,\mu\text{C}^2\right)^2 \left(245 - 8\,\mu\text{C}^2 + 8\,\mu\text{C}^4\right) - 99\,\mu\text{3}^8 \left(5 - 24\,\mu\text{C}^2 + 24\,\mu\text{C}^4\right) - 12\,\mu\text{3}^6 \left(31 - 151\,\mu\text{C}^2 + 151\,\mu\text{C}^4\right) + \mu^3^2 \left(676 + 2220\,\mu\text{C}^2 - 3948\,\mu\text{C}^4 + 3456\,\mu\text{C}^6 - 1728\,\mu\text{C}^8\right) + 54\,\mu\text{3}^4 \left(-1 - 162\,\mu\text{C}^2 + 210\,\mu\text{C}^4 - 96\,\mu\text{C}^6 + 48\,\mu\text{C}^8\right) \right) + \frac{3}{2}\,\mu^2^4 \left(-6\,\mu\text{3}^6 \left(-7 - 6\,\mu\text{C}^2 + 6\,\mu\text{C}^4\right) - \left(1 - 2\,\mu\text{C}^2\right)^2 \left(37 - 12\,\mu\text{C}^2 + 12\,\mu\text{C}^4\right) + 11\,\mu\text{3}^8 \left(23 - 84\,\mu\text{C}^2 + 84\,\mu\text{C}^4\right) + 18\,\mu\text{3}^2 \left(-1 - 162\,\mu\text{C}^2 + 210\,\mu\text{C}^4 - 96\,\mu\text{C}^6 + 48\,\mu\text{C}^8\right) - 4\,\mu\text{3}^4 \left(60 - 991\,\mu\text{C}^2 + 1243\,\mu\text{C}^4 - 504\,\mu\text{C}^6 + 252\,\mu\text{C}^8\right) \right) \right) \right)
```

lhs7 = %;

```
lhs7 = caseA[wA, 7] + caseC[wC, \muC, 7] + case2[w2, \mu2, 7] + case2[w3, \mu3, 7];
```

Expand[%];

Together[%];

Cancel[%];

$$- \left(9\,\mu 2^{10} \left(9\,\mu 3^{6} - (1 - 2\,\mu C^{2})^{2} + \mu 3^{4} \left(-23 + 84\,\mu C^{2} - 84\,\mu C^{4}\right) + 3\,\mu 3^{2} \left(5 - 24\,\mu C^{2} + 24\,\mu C^{4}\right)\right) + \mu 2^{8} \left(10 + 9\,\mu 3^{2}\right) \left(9\,\mu 3^{6} - (1 - 2\,\mu C^{2})^{2} + \mu 3^{4} \left(-23 + 84\,\mu C^{2} - 84\,\mu C^{4}\right) + 3\,\mu 3^{2} \left(5 - 24\,\mu C^{2} + 24\,\mu C^{4}\right)\right) + \mu 2^{6} \left(7 + 10\,\mu 3^{2} + 9\,\mu 3^{4}\right) \left(9\,\mu 3^{6} - (1 - 2\,\mu C^{2})^{2} + \mu 3^{4} \left(-23 + 84\,\mu C^{2} - 84\,\mu C^{4}\right) + 3\,\mu 3^{2} \left(5 - 24\,\mu C^{2} + 24\,\mu C^{4}\right)\right) + (1 - 2\,\mu C^{2})^{2} \left(47 - 7\,\mu 3^{6} - 10\,\mu 3^{8} - 9\,\mu 3^{10} + 12\,\mu C^{2} - 12\,\mu C^{4} - 8\,\mu 3^{2} \left(9 - \mu C^{2} + \mu C^{4}\right) + \mu 3^{4} \left(51 - 36\,\mu C^{2} + 36\,\mu C^{4}\right) + \mu 2^{2} \left(-8 \left(1 - 2\,\mu C^{2}\right)^{2} \left(9 - \mu C^{2} + \mu C^{4}\right) + 27\,\mu 3^{10} \left(5 - 24\,\mu C^{2} + 24\,\mu C^{4}\right) + 3\,\mu 3^{8} \left(47 - 228\,\mu C^{2} + 228\,\mu C^{4}\right) + \mu 3^{6} \left(95 - 464\,\mu C^{2} + 464\,\mu C^{4}\right) + \mu 3^{4} \left(-349 + 4348\,\mu C^{2} - 6940\,\mu C^{4} + 5184\,\mu C^{6} - 2592\,\mu C^{8}\right) + \mu 3^{2} \left(50 - 1872\,\mu C^{2} + 3600\,\mu C^{4} - 3456\,\mu C^{6} + 1728\,\mu C^{8}\right)\right) - \mu 2^{4} \left(-3 \left(1 - 2\,\mu C^{2}\right)^{2} \left(17 - 12\,\mu C^{2} + 12\,\mu C^{4}\right) - 4\,\mu 3^{6} \left(-5 - 24\,\mu C^{2} + 24\,\mu C^{4}\right) + 9\,\mu 3^{10} \left(23 - 84\,\mu C^{2} + 84\,\mu C^{4}\right) + \mu 3^{8} \left(95 - 192\,\mu C^{2} + 192\,\mu C^{4}\right) - 4\,\mu 3^{4} \left(155 - 1376\,\mu C^{2} + 2132\,\mu C^{4} - 1512\,\mu C^{6} + 756\,\mu C^{8}\right) + \mu 3^{2} \left(349 - 4348\,\mu C^{2} + 6940\,\mu C^{4} - 5184\,\mu C^{6} + 2592\,\mu C^{8}\right)\right)\right) \right)$$

lhs7 = %;

NSolve 
$$\left[ \{ 1hs5 == \frac{1}{11}, 1hs6 == \frac{1}{13}, 1hs7 == \frac{1}{15} \}, \{ \mu C, \mu 2, \mu 3 \} \right]$$

\$Aborted

Solve 
$$\left[1hs5 = \frac{1}{11}, \{\mu C\}\right]$$

```
(* Large output deleted *)
```

```
Length[%]
```

4

(\* The four solutions above correspond to the base angle and its reflections, I will continue only using the last solution above \*)

```
\mu C = \mu C / . \%78[[4]]
```

Expand[%];

Together[%];

```
(\sqrt{(-1+23 \mu 3^2 - 55 \mu 3^4 +
                                    33\,\mu 3^6 + 33\,\mu 2^6\,\,(1 - 18\,\mu 3^2 + 21\,\mu 3^4) + \mu 2^2\,\,(23 - 448\,\mu 3^2 + 1023\,\mu 3^4 - 594\,\mu 3^6) + 1023\,\mu 3^6 - 594\,\mu 3^6 - 594\,\mu 3^6 - 596\,\mu 3^6 - 596\,\mu
                                   11\,\mu2^4\,\left(-5+93\,\mu3^2-159\,\mu3^4+63\,\mu3^6\right)\,-\,\mu2\,\mu3\,\sqrt{(71-1807\,\mu3^2+8006\,\mu3^4-14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14190\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+14100\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^6+1400\,\mu3^
                                                                 11187\,\mu3^8 - 3267\,\mu3^{10} + 1089\,\mu2^{10}\,\left(-3 + 52\,\mu3^2 - 18\,\mu3^4 - 204\,\mu3^6 + 189\,\mu3^8\right) \,+
                                                                  33 \mu 2^8 (339 – 6287 \mu 3^2 + 9822 \mu 3^4 + 7698 \mu 3^6 – 18513 \mu 3^8 + 6237 \mu 3^{10}) –
                                                                 2\,\mu 2^4\,\left(-4003+83437\,\mu 3^2-287774\,\mu 3^4+360954\,\mu 3^6-162063\,\mu 3^8+9801\,\mu 3^{10}\right)\,-
                                                                 22 \mu 2^6 (645 – 12706 \mu 3^2 + 32814 \mu 3^4 – 19704 \mu 3^6 – 11547 \mu 3^8 + 10098 \mu 3^{10}) +
                                                                \mu 2^2 \left(-1807 + 40008 \,\mu 3^2 - 166874 \,\mu 3^4 + 279532 \,\mu 3^6 - 207471 \,\mu 3^8 + 56628 \,\mu 3^{10}\right)\right)\Big) \Big/
        (\sqrt{2} \sqrt{(-1 + 23 \mu 3^2 - 55 \mu 3^4 + 33 \mu 3^6 + 33 \mu 2^6 (1 - 18 \mu 3^2 + 21 \mu 3^4) + 
                                          \mu 2^2 \left( 23 - 448 \,\mu 3^2 + 1023 \,\mu 3^4 - 594 \,\mu 3^6 \right) + 11 \,\mu 2^4 \left( -5 + 93 \,\mu 3^2 - 159 \,\mu 3^4 + 63 \,\mu 3^6 \right) \right) \right)
μC = %;
Expand[lhs6];
 Together[%];
  PowerExpand[%];
  Together[%];
  Cancel[%];
  PowerExpand[%92];
  Simplify[%]
   (* Large ugly equation deleted *)
  lhs6 = \%;
  Expand[lhs7];
  Together[%];
  Cancel[%];
  PowerExpand[%];
   Simplify[%];
   1hs7 = %
    (* Large ugly equation deleted *)
 NSolve \left[ \left\{ 1hs6 = \frac{1}{13}, 1hs7 = \frac{1}{15} \right\}, \left\{ \mu 2, \mu 3 \right\} \right]
```

Out of memory. Exiting.

```
quadAC23 = FindRoot[Eqns, {wA, .0130608}, {wC, .0154866},
{w2, .00673874}, {w3, .0161761}, {\muC, .410254}, {\mu2, .848421}, {\mu3, .300144},
AccuracyGoal \rightarrow 24, WorkingPrecision \rightarrow 34]
```

 $\{ wA \rightarrow 0.01306075218457543404037831995817255, wC \rightarrow 0.01548662322913343575994122509932307, w2 \rightarrow 0.00673874365125243712696660016440459, w3 \rightarrow 0.01617611174013693526966426141339587, \muC \rightarrow 0.410253515086337171560174981725413, \mu2 \rightarrow 0.848421498634701466179649259967932, \mu3 \rightarrow 0.3001438436359286871701637504638956 \}$ 

(\* The FindRoot function is a root solving function. The initial guesses used are from a previous effort on a slower computer. This is one possible solution to the system of equations. \*)

# **Appendix C: Valid Quadrature Base Sets**

# N = 3

Cases	weight	mu	eta	xi
AB				
A	0.0476190476190476	1	0	0
B	0.0380952380952381	0.7071067811865475	0	0.7071067811865475
1	0.0321428571428571	0.5773502691896258	0.5773502691896258	0.5773502691896258
C1				
C	0.0309523809523810	0.9000482411921158	0	0.4357902747044488
1	0.0321428571428571	0.5773502691896258	0.5773502691896258	0.5773502691896258

# N = 5

Cases	weight	mu	eta	xi
AB12				
A	0.0126984126984127	1	0	0
В	0.0225749559082892	0.7071067811865475	0	0.7071067811865475
1	0.0210937500000000	0.5773502691896258	0.5773502691896258	0.5773502691896258
2	0.0201733355379189	0.9045340337332910	0.3015113445777637	0.3015113445777637
AC2				
A	0.0223062916968982	1	0	0
С	0.0185141807544452	0.8616774905910132	0	0.5074563057138757
2	0.0175759129879969	0.7858759158676477	0.4372636760921183	0.4372636760921183
BC2				
В	0.0182978666108076	0.7071067811865475	0	0.7071067811865475
С	0.0149418203732660	0.9526970315441012	0	0.3039216446504885
2	0.0175759129879969	0.7858759158676477	0.4372636760921183	0.4372636760921183
C4				
C	0.0170731707317073	0.9451972790030246	0	0.3264997760570111
4	.01229674796747967	0.7573160195117944	0.5872843412419646	0.2856038324721905

N = 7

Cases	weight	mu	eta	xi
ABC12				
A	0.0090481888301554	1	0	0
C	0.0064514915385668	0.9545808669401723	0	0.2979519566503114
В	0.0210324604374280	0.7071067811865475	0	0.7071067811865475
1	0.0182794139234181	0.5773502691896258	0.5773502691896258	0.5773502691896258
2	0.0163437597273743	0.8753170875981718	0.7854361833270270	0.7854361833270270
AC23				
A	0.0130607521845754	1	0	0
C	0.0154866232291334	0.911971520037388	0	0.4102535150863372
2	0.0067387436512524	0.8484214986347015	0.374286628571238	0.374286628571238
3	0.0161761117401369	0.3001438436359287	0.674504882535127	0.674504882535127

machine proton × generalized input module run on 10/ 5/97 with solver version 05-05\*product release 3.0 0.0163437597273742918 0.0163437597273742918 0.0022620472075388518 0.01827941392341814 ievt=0 ibl=1 ibr=0 ibt=0 ibb=1 norm=.125 fluxp=1 angp=0 epsi=1.-6 This is the M7 Quadrature set with 5 dummy ordinates added im=1 jm=1 km=1 it=16 jt=16 kt=16 maxscm= 218100 maxlcm=64800 /\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\* В Г О С К V \*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\* \*\*\*\*\*\*\*\*\*\*\*\*\*\*\*\* ...listing of cards in the input stream... in cubeshield geometry. THREEDANT THESIS TEST PROBLEM 2: (x, y, z) cube source 0.0032257457692834124 0.0032257457692834124 0.0032257457692834124 igeom=x-y-z ngroup=1 isn=12 niso=1 mt=1 nzone=1 0.010516230218714844 iso2 1.0 sourcf=8r1 f0;7y1;f0;7y1;7y16;f0;15y1;7y16 /\*\* to fill in the required array values. /isotope 2 microscopic cross sections isotope 1 microscopic cross šections 0.25 0.00 0.75 0.50 balp=1 ibback=0 ibfrnt=1 sourcp=2 mix2 wgt=0.0022620472075388518 0.0032257457692834124 0.0163437597273742918 matls=mix1 iso1 1.0 /; 0.010516230218714844 zints=16 yints=16 0.50 0.00 0.75 0 ymesh=-0 2.0
zmesh=-0 2.0 assign=matls 0 zones=f1 /0.25 2 \*\*/ Ļ ų 27. 26. 29. 30. 32. 35. 36. 28. 31. 33. 39. 40. 41. 37. \* \* \* × \* \*

*	*	42.	0.0032257457692834124	0.010516230218714844	0.0032257457692834124	*
*	*	43.	0.0022620472075388518	0 0 0 0		*
*	*	44.	00000000000000000000000000000000000000	0.297951956650311267	0.34192104071	*
*	*	45.	0.0000000000000000000000000000000000000	0.707106781188	0.57735026919	*
*	*	46.	0.0000000000000000000000000000000000000	0.954580866940172434	0.875317087598171994	*
*	*	47.	0.34192104071	0.0000000000000000000000000000000000000	1.0000000000000000000000000000000000000	*
*	*	48.	0.9545808669401724	0.70710678118655	0.29795195665031112	*
*	*	49.	0.0000000000000000000000000000000000000	.1.1.1.1.1		*
*	*	50.	eta=0.00000000000000000000	0.0000000000000000000000000000000000000	0.3419210407	*
*	*	51.	0.297951956650311267	0.0000000000000000000000000000000000000	0.57735026919	*
*	*	52.	0.707106781188	0.0000000000000000000000000000000000000	0.3419210407	¥
*	*	53.	0.875317087598171994	0.954580866940172434	0.0000000000000000000000000000000000000	*
*	*	54.	0.29795195665031107	0.707106781186545	0.95458086694017243	*
*	*	55.	1.0000000000000000000000000000000000000	.1.1.1.1.1		*
*	*	56.	ιt			*
*	*	57.				*
*	*	58.				*
*	*	59.	/************************ B T O	T O C K VI *******************	* * * * * * *	* .
*	*	60.	pted=1			*
*	*	61.	edoutf=2			*
*	*	62.	t			*
*	*					*
*	**	******	*****	*******	*******	*****
1 DETATI		ות יידומייי	181.8481			
-	1100					

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""""""""""""""""""""""""""""""""""""""	的复数 化化合物 化化合物 化化合物 化化合物 化化合物 化化合物 化化合物 化化化合物 化化化化化 化化化化化化化化
flux and eigenvalue convergence as monitored by threedant	ored by threedant
***************************************	* * * * * * * * * * * * * * * * * * * *
***************	
*key start iteration monitor * ***********************************	
<ul> <li>(sec) no. inners sub-outers sourc multip lambda-1 flux</li> <li>.87 0</li> </ul>	flux change fiss change converged
inner iteration summary for outer iteration no.	1
iter per max flux	
group group change mesh 1 8 .54E-06 1, 8, 16	
* cpu time outer diffusion max ptwise	wise max ptwise inners

* (S6 * 11.	(sec) no. inners 11.49 1 8	sub-outers so 0 1.	sourc multip lambda-1 1.00000000 -1.00000E+00	flux change fiss ( 5.40926E-07 0.0000	fiss change converged 0.00000E+00 yes		
* * -	\$\$\$\$\$	\$\$ all convergence criteria	criteria satisfied	\$\$\$\$\$\$			
* * + + pai	particle balance = -3.52024E-08		total inners all out	outers = 8			
, , , , , , , , , , , , , , , , , , ,	* ************************************	*****	******	***************************************	*****	*******	*****
****	********	***********	***********	***************************************	*********	******	************
* * ·			group edit	and balances upon	convergence		
* * * * *	******	*****	***********	* ************************************	*****	******	**********
* * *		title THREEDANT	THESIS TEST	PROBLEM 2: (x, y, z) sp	sphere source		
* * *			system balance	ance tables(neutrons	ons only)		
* * * * * * * * * * * * * * * * * * *	* ************************************	******* 1]e *******					
db * *	source	fission source	absorption	in scatter	self scatter	out scatter	net leakage
 * * *	1.2500000E-01	0.0000000000000000000000000000000000000	6.5240937E-02	0.0000000E+00	1.3048187E-01	8.3266727E-16	5.9759067E-02
* * * * tot	1.250000E-01	0.000000E+00	6.5240937E-02	0.000000E+00	1.3048187E-01	8.3266727E-16	5.9759067E-02
db * * *		right leakage horizontl leakage	top leakage	vertical leakage	back leakage	fr-back leakage	particle balance
ب * * •	1.9919689Е-02	1.9919689E-02	1.9919689Е-02	1.9919689Е-02	1.9919689E-02	1.9919689E-02	-3.5202419Е-08
+ * * * to t	1.9919689Е-02	1.9919689Е-02	1.9919689Е-02	1.9919689Е-02	1.9919689Е-02	1.9919689Е-02	-3.5202419Е-08
db * * *	left leakage	bottom leakage	front leakage	nprod spectrum			
	0.0000000E+00	0.000000E+00	0.000000E+00	0.0000000E+00			
* tot +	0.000000E+00	0.00000000000000	0.0000000E+00	0.0000000E+00			
* * *	* ********************	*********	*******	***************************************	**********	********	*********

.04 seconds. ----1 and plane .08 .09 flux components for group, .34 10.24 0.0000000E+00 8.3266727E-16 1.9919689E-02 neutron -3.5202419E-08 5.9759067E-02 1.9919689E-02 1.9919689E-02 1.9919689E-02 1.9919689E-02 1.2500000E-01 0.0000000E+00 6.5240937E-02 1.3048187E-01 1.9919689E-02 Multigrid average convergence rate by group... WU. timing info...tswep,tdsa,trelx,tput3,tintrp= 75.54 neutron neutron Total= neutron Multigrid work units... integral summary information summary eigenvalue not used integral-horizontal lkage-i integral-vertical lkage-i integral-fr-back lkage-i integral-particle bal-i By group... integral-right lkage-i integral-absorption-i integral-back lkage-i integral-self-scak-i integral-net lkage-i integral-top lkage-i integral-out-scak-i integral-fission-i integral-in-scak-i 75.54 integral-source-i .5720 н

\* \*

... interface file rtflux written..

threedant iteration time, mins 1.9200E-01

*		edit rı	run on 10/ 5/97 with solver version 05-05-95*product release 3.0 machine F	e proton
*		*****	***************************************	* * * * * * * * * *
***************************************	*********	*******	~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~~	********
* * * * * * * * * * * * * * * * * * * *		*****	***************************************	*****
* =			edit output	
× ***********************************	********	******	***************************************	*******
* .				
× *			block vi - edit specification data	
*				
* ****************				
*key start edit output * **************************				
**********************************	********	****	***************************************	
*				
<pre>* * cross section balancing (balxs.ne.0)</pre>	ng (balxs.n	e.0)		
* or				
<pre>* transport correction  * will NOT be reflected .</pre>	(trcor=diag in edits	, cesaro,	o, or bhs)	
< *				
****************	****	*****	***************************************	
* .			input control integers	
.* * *	pted zned	0/1 0/1	no/yes - point edits desired no/yes - zone edits desired	
* * +	ajed	0/1	direct/adjoint edit(use rtflux/atflux file)	
	igrped	0/1/2/3	print totals only/print broad groups only/same as 1/print all groups and	totals
* *	byvolp rzflux	0/1	no/yes - multiply point reaction rates by mesh volumes no/ves - write the rzflux file (zone average flux file)	
	rzmflx	0/1	- write the rzmflx file (zone average	
* *			floating parameters	
*			•	
* 0.000000E+00 * 2.100000E+02 *	00 power 02 mevper	0/p mev per	no/normalize all results, including flux files, to p megawatts per fission (default: 210 mev)	
*				

*	.energy	energy related edit information	
* * * * * * *	1 nur 0 nur 1 tot 1 tot	number of fine neutron groups number of fine gamma groups total number of fine groups total number of broad groups	
	. space	space related edit information	
40.	4096 nur 0 nur 1 0/1	number of points to edit number of zones to edit 0/1 no/yes density factors were input 0/1/2/3 no/solver/edit/both zones mass edit (requires atomic weights to be present)	input s mass edit be present)
* ************************************	**** ***** ****** ****** ****** ****** ****	**********************************	* ************************************
**************************************	****** -edtogx ******	**************************************	
*******************	* * * * * * * * * * * * * * * * * * * *	***************************************	***************************************
* *		run híahliahts	
*********			
*key start run nigniignus ~ ~			
* *	*****	***************************************	**
*		all modules are tentatively go.	*
*		interface file geodst written.	* .
* *	* *	cross sections from cards. *kwikrd overridden for ifido=2*	* *
*	بط *		*
*		interface file asgmat written	*
*		xs files macrxs, snxedt written.	*
*	*	solinp	* .
*	* ÷	interface file editit written.	<b>*</b> -
*	* +	start solver execution.	* 1
+ +	* *	all convergence criteria met.	* *
* *	< *	interface file sncons written.	< *

\* × \* start edit execution. edits completed.

\* \* ×

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# storage and timing history

*		scm	scm	Lcm	LCM	cpu	sysio
* *	module	words	limit 	words	limit 	seconds	seconds
*		0	0	0	0	11.7	0.
*	100	7717	218100	0	0	8.	0.
*	101	0	0	0	0	.2	0.
*	102	1036	218100	m	64800	0.	0.
*	103	26	218100	0	0	0.	0.
ىد	104	2018	218100	0	0	0.	••
J.	105	0	0	0	0	0.	0.
L.	106	0	0	0	0	0.	۰.
	107	50	218100	7	64800	0.	0.
	108	1502	218100	305	64800	•	••
	109	7722	218100	0	0	0.	0.
	112	0	0	0	0	••	0.
	200	132425	218100	62877	64800	10.7	0.
	201	0	0	0	0	0.	0.
	202	0	0	0	0	••	••
	203	3823	218100	0	0	0.	0.
	204	0	0	0	0	••	•
×	205	0	0	0	0		0.
×	206	0	0	0	0	10.6	0.
*	207	0	0	0	0	•	0.
÷	208	0	0	0	0	••	0.
*	210	0	0	0	0	0.	0.
*	211	0	0	0	0	0.	0.
*	300	4214	218100	0	0	.1	•
*	301	33022	218100		64800	•	0.
JL.	302	0	0	0	0	••	•
	400	0	0	0	0	•	•

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13. ABSTRACT (Maximum 200 words) Several new quadrature sets for use in the discrete ordinates method of solving the Boltzmann neutral particle transport						
equation are derived. These symmetric quadratures extend the traditional symmetric quadratures by allowing ordinates						
perpendicular to one or two of the coordinate axes. Comparable accuracy with fewer required ordinates is obtained.						
Quadratures up to seventh order are presented. The validity and efficiency of the quadratures is then tested and compared						
with the Sn level symmetric quadratures relative to a Monte Carlo benchmark solution. The criteria for comparison include						
with the Sn level symmetric quadratures relative to a Monte Carlo benchmark solution. The criteria for comparison include current through the surface, scalar flux at the surface, volume average scalar flux, and time required for convergence.						
current through the surface, scalar flux at the surface, volume average scalar flux, and time required for convergence. Appreciable computational cost was saved when used in an unstructured tetrahedral cell code using highly accurate						
characteristic methods. However, no appreciable savings in computation time was found using the new quadratures compared with traditional Sn methods on a regular Cartesian mesh using the standard diamond difference method. These						
compared with traditional Sn methods on a regular Cartesian mesh using the standard diamond difference method. These muddratures are recommended for use in three-dimensional calculations on an unstructured mesh.						
quadratures are recommended for use in three-dimensional calculations on an unstructured mesh.						
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